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Monte Carlo Wavefunctions - Quantum Trajectories

The "Schrodinger equation" for a 2-level atom in a laser field with spontaneous emission is

$$i\hbar \frac{d}{dt} \begin{pmatrix} C_g \\ C_e \end{pmatrix} = \hbar \begin{pmatrix} \delta/2 & \Omega/2 \\ \Omega^*/2 & -\delta/2 \end{pmatrix} \begin{pmatrix} C_g \\ C_e \end{pmatrix} - \begin{pmatrix} 0 \\ -i\hbar \frac{\gamma}{2} C_e \end{pmatrix}$$

$$= \hbar \begin{pmatrix} \delta/2 & \Omega/2 \\ \Omega^*/2 & -\delta/2 - i\frac{\gamma}{2} \end{pmatrix} \begin{pmatrix} C_g \\ C_e \end{pmatrix}$$

effectively, $H = \hbar \begin{pmatrix} \delta/2 & \Omega/2 \\ \Omega^*/2 & \delta/2 - i\frac{\gamma}{2} \end{pmatrix}$

non-Hermitian

↳ does not conserve probability

(complex eigenenergies)

↳ time evolution: $e^{-i \frac{(E_e)}{\hbar} t} \pm \frac{C_e}{\hbar}$

How do the states evolve in time?

Before a photon is emitted:

$$|\Psi_0(t)\rangle = \left[C_g(t) |g\rangle_a |N+1\rangle_l + C_e(t) |e\rangle_a |N\rangle_l \right] \otimes |0\rangle_{\text{all other vacuum photon states}}$$

$C_g(t)$ & $C_e(t)$ are given by the previous "Schrodinger" equation.

After a photon is emitted at time $t = t_g$, then

$$|\Psi_1(t=t_g)\rangle = |g\rangle \otimes \left[\sum_{k, \epsilon} \alpha_{k, \epsilon} |k, \epsilon, 1\rangle \right]$$

all possible "i" photon states

photon polarization
momentum
accounts for dipole spatial emission distribution

Actually, the system (atom + laser + universe) could be in a superposition of these 2 states:

(leaked photon not yet detected)

$$|\Psi(t)\rangle = \underbrace{\sqrt{1-\beta^2}}_{\text{for short times}} \underbrace{|\Psi_0(t)\rangle}_{\text{no photon emitted}} + \beta \underbrace{|\Psi_1(t)\rangle}_{\text{1 photon leaks into an infinite superposition of photon momentum states}}$$

Loss of probability from H ends up here

$$\beta = \underbrace{\sqrt{\gamma |C_e|^2 dt}}_{\text{for short times}}$$

Environment Detector

But we are not dealing with a closed system, but an open quantum system. The "leaked photon" could be detected at any moment by the environment [environment is like a giant detector].

Let's assume that the "environment detector" is 100% efficient and makes a measurement in a time dt immediately after a photon leaks out \rightarrow wavefunction collapses

Immediately after collapse:

"photon not emitted state"

keeps evolving as before, but needs to be renormalized, since system is definitely here.

$$|\psi(t+dt)\rangle = \begin{cases} |\psi_0(t+dt)\rangle \text{ with probability } p = 1 - \beta^2 \\ \text{or} \\ |\psi_1(t+dt)\rangle = |g\rangle |k, \epsilon, l\rangle \text{ with probability } p = \beta^2 \end{cases}$$

photon collapses to some k, ϵ state, but we don't pay attention to it.

\hookrightarrow process repeats \rightarrow propagate a time dt again.

This story suggests the following algorithm for calculating the time evolution of the "atom + laser + universe/environment detector" system --- though we only pay attention to the atomic state.

Step 1: at time = t the state of the system is

$$|\psi(t)\rangle = |\psi_0(t)\rangle \\ = c_g(t)|g\rangle + c_e(t)|e\rangle$$

Step 2: at time = $t+dt$ we have 2 possible trajectories for the wavefunction

$$|\psi(t+dt)\rangle = \begin{cases} \text{case A: } \frac{1}{\sqrt{1-dp}} [c_g(t+dt)|g\rangle + c_e(t+dt)|e\rangle] \\ \text{with probability } 1-dp \\ \text{OR} \\ \text{case B: } |g\rangle \text{ with probability } dp \end{cases}$$

↖ re-normalization

with $dp = \gamma |c_e|^2 dt = \beta^2$

How do you pick between case A & case B?

↳ "flip a coin": choose a random number $\epsilon \in [0,1]$

$$\begin{cases} \text{if } \epsilon > dp \Rightarrow |\psi(t+dt)\rangle = |\text{case A}\rangle \\ \text{if } \epsilon < dp \Rightarrow |\psi(t+dt)\rangle = |\text{case B}\rangle = |g\rangle \end{cases}$$

Condition on dt: $dt = \Delta t \ll \frac{1}{\Omega}, \frac{1}{\gamma}, \frac{1}{\delta}$

↑
not infinitesimal (though that would be ideal)



Random number generator must work well for very small number i.e. $\epsilon \approx 0$.

Electronic structure of Alkali Atoms

(I) Fine structure Hamiltonian (i.e. includes spin-orbit coupling)

$$H_0 = \frac{p^2}{2m_e} - \frac{e^2}{R} + H_{LS}$$

$V(R)$

with $H_{LS} = -\vec{\mu}_e \cdot \vec{B}'$

$\frac{e}{m_e} \vec{S}$

↑ effective magnetic field seen by the electron because it is moving through an electric field.

[in frame of e^- , "relativistic" effect]

$$\vec{B}' = -\frac{1}{c^2} \vec{v}_e \times \vec{E}$$

$\vec{v}_e = \frac{\vec{p}}{m_e}$

$-\frac{1}{e} \frac{1}{r} \frac{dV}{dr} \vec{r}$

$\vec{p} \times \vec{r} = -\vec{L}$

thus $H_{LS} = \frac{1}{2} \frac{e^2}{m_e^2 c^2} \frac{1}{R^3} \vec{L} \cdot \vec{S} = E_{LS} \vec{L} \cdot \vec{S}$

due to
Thomas precession
↳ rotating relativistic
frame effect

$$= \frac{E_{LS}}{2} (\vec{J}^2 - \vec{L}^2 - \vec{S}^2)$$

$$= \frac{E_{LS}}{2} [J(J+1) - L(L+1) - S(S+1)]$$

Recall:

$$\vec{J} = \vec{L} + \vec{S}$$

$$J^2 = L^2 + S^2 + 2\vec{L} \cdot \vec{S}$$

Possible values for J : $J = L+S, L+S-1, \dots, |L-S|$

Example: for $L=1, S=1/2 \Rightarrow J = 3/2$ or $1/2$