

Thursday, March 7, 2024

Spontaneous Emission

I - Basic Physics

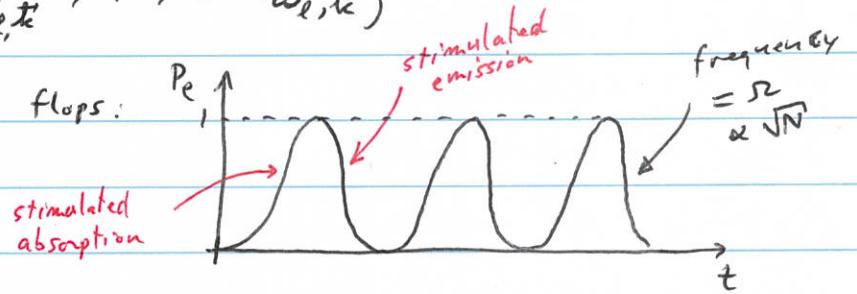
Recall the dressed atom Hamiltonian:

$$H = H_{\text{atom}} + H_{\text{Em field}} + H_{\text{interaction}}$$

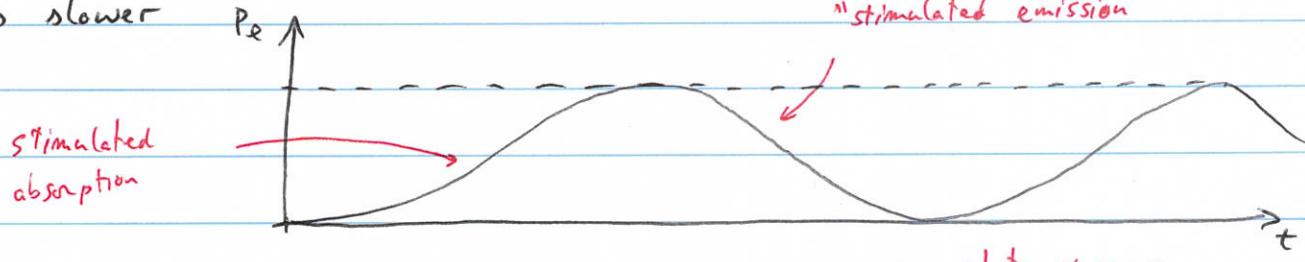
$$= \begin{pmatrix} E_g & 0 \\ 0 & E_e \end{pmatrix} + \hbar \omega_{\text{c}, t_0} \begin{pmatrix} N+1 & 0 \\ 0 & N \end{pmatrix} + \hbar \begin{pmatrix} 0 & \mathcal{R}/2 \\ \mathcal{R}^*/2 & 0 \end{pmatrix}$$

basis: $\{|g\rangle|N+1\rangle_{\omega_c, t_0}, |e\rangle|N\rangle_{\omega_c, t_0}\}$

Atom-field system Rabi flops:



Suppose that we have a $N=1$ laser, then the Rabi-flapping is slower



The basis is $\{|g\rangle|1\rangle_{\omega_c, t_0}, |e\rangle|0\rangle_{\omega_c, t_0}\}$

energetically, we can include $|e\rangle|0\rangle_{\omega_c, t_0}$ as well

So the basis is really

$$\{ |g\rangle|1\rangle_{\omega_e, t_e}, |e\rangle|0\rangle_{\omega_e, t_e}, |e\rangle|0\rangle_{\omega_e, t_e' \neq t_e}, \dots, |g\rangle|1\rangle_{\omega_e, t_e'} \}$$

during "stimulated emission" photon could go to $t_e' \neq t_e$

infinite number

II Spontaneous Emission

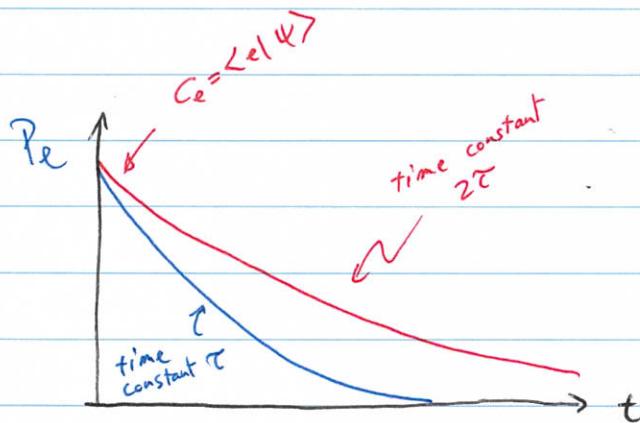
In other words, an excited atom can undergo "stimulated emission" by the photon vacuum in any direction.

- An atom in an excited state will tend to decay into a superposition of many $|g\rangle|1\rangle$ "one" photon states (all pointing in different directions).
- The more states into which an atom can "Rabi flop" or leak into, the faster the decay.

→ The decay is exponential with time constant τ .

For a 2-level atom

$$\tau = \frac{1}{\gamma} = \frac{3\pi \epsilon_0 \hbar c^3}{|\langle e | R | g \rangle|^2 \omega_{eg}^3} \quad \approx 27 \text{ ns for Rb}$$



→ # density of photon t_e' states scales like ω_{eg}^3

$$\frac{dP_e}{dt} = -\gamma P_e$$

$$\frac{dC_e}{dt} = -\frac{\gamma}{2} C_e$$

(III) Derivation

[originally derived by Wigner & Weisskopf]

[method: show that decay is exponential, i.e. $\frac{dc_e}{dt} = -\frac{\gamma}{2} c_e$

Dressed atom wavefunction for decaying excited states

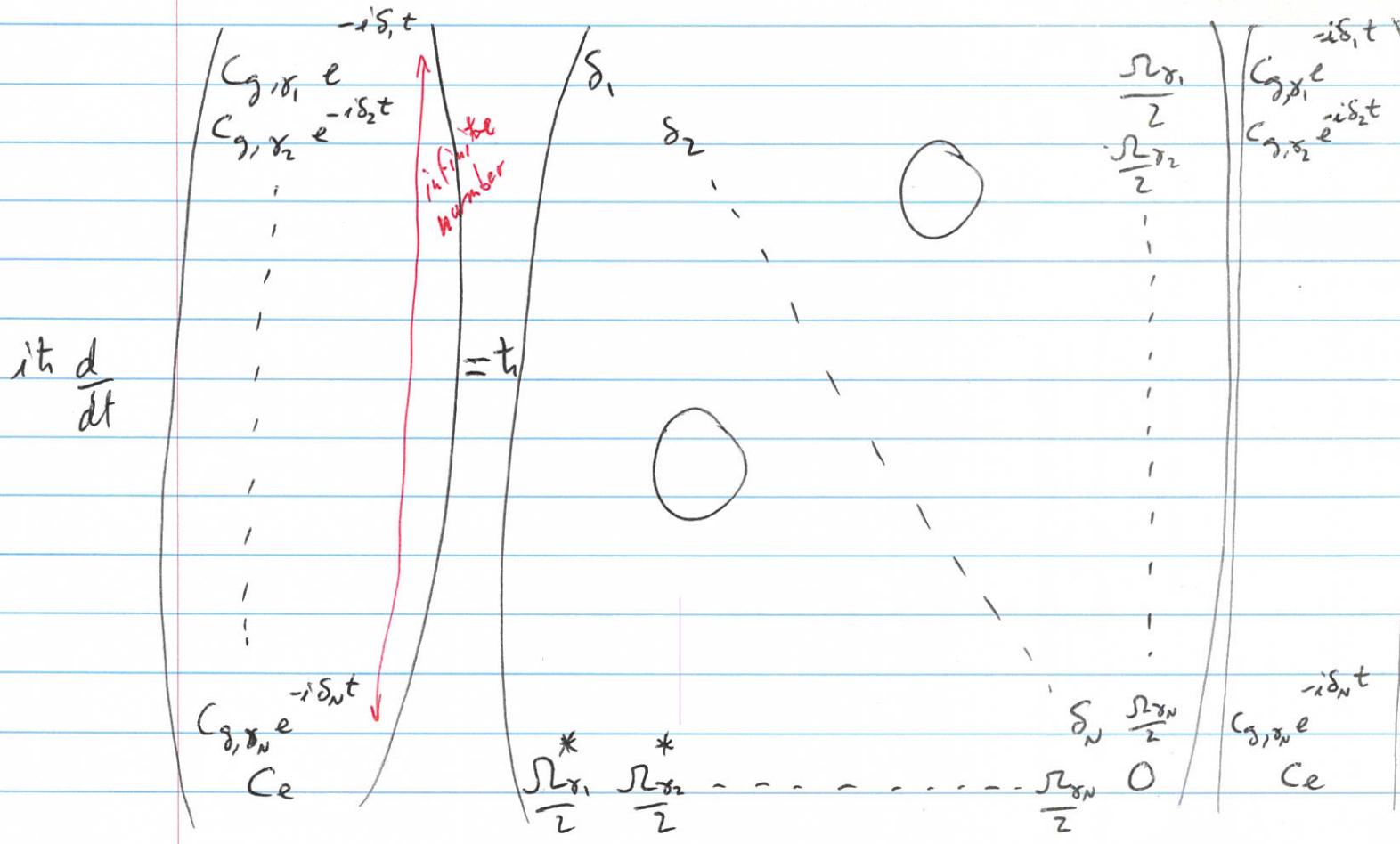
$$|\Psi(t)\rangle = c_e(t)|e\rangle + \sum_{\gamma, \text{ photon states}} c_{g,\gamma}(t) e^{-i\delta_\gamma t} |g\rangle |1\rangle_\gamma$$

with $S_\gamma = \omega_\gamma - \omega_{eg}$

we would like to write $\delta_\gamma = 0$

to conserve energy, but since the excited state exists for a finite time \rightarrow then there ~~exists~~ is an energy spread.

The Schrödinger equation becomes:



$$\Rightarrow i \frac{d}{dt} C_e(t) = \sum_{\gamma \text{ states}} C_{g,\gamma}(t) \frac{\mathcal{R}_\gamma^*}{2} e^{-i\delta_\gamma t} \quad (1)$$

$$\left\{ \begin{array}{l} i \frac{d}{dt} C_{g,\gamma}(t) \cdot e^{-i\delta_\gamma t} + \delta_\gamma \cdot C_{g,\gamma}(t) e^{-i\delta_\gamma t} = \delta_\gamma C_{g,\gamma} e^{-i\delta_\gamma t} + \frac{\mathcal{R}_\gamma}{2} C_e(t) \\ \hookrightarrow i \frac{d}{dt} C_{g,\gamma}(t) = \frac{\mathcal{R}_\gamma}{2} C_e(t) e^{+i\delta_\gamma t} \end{array} \right. \quad (2)$$

↑ infinite number
of equations
(1 for each γ state)

next, integrate (2)

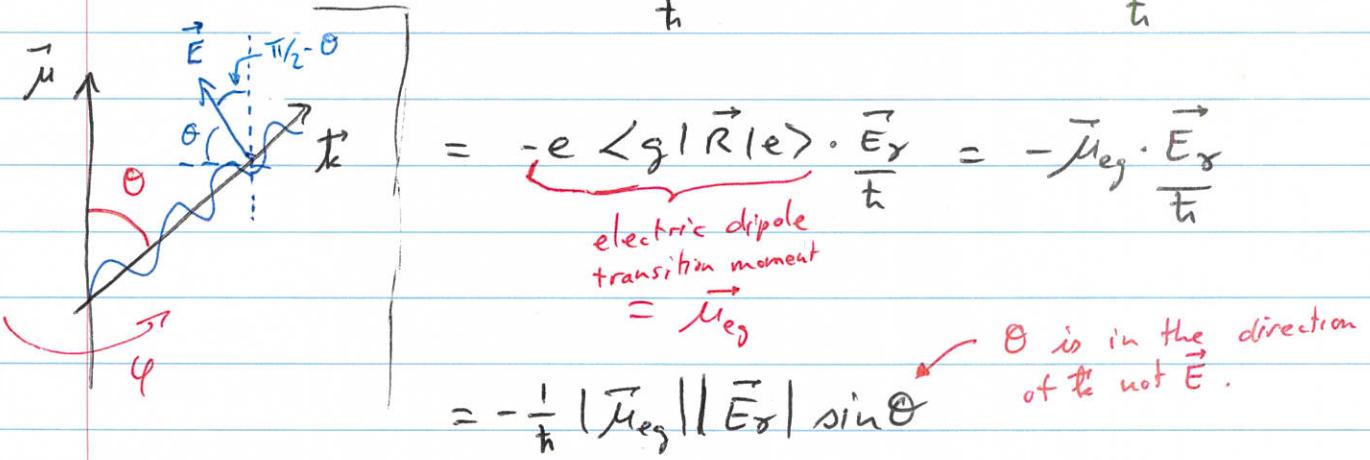
$$C_{g,\gamma}(t) = -i \frac{\mathcal{R}_\gamma}{2} \int_0^t e^{i\delta_\gamma t'} C_e(t') dt'$$

substitute into (1) :

$$\frac{d}{dt} C_e(t) = - \sum_{\gamma} \underbrace{\frac{(\mathcal{R}_\gamma)^2}{4}}_{\text{evaluate this part in the hope of pulling } C_e(t') \text{ term out of integral.}} \int_0^t e^{-i\delta_\gamma(t-t')} C_e(t') dt'$$

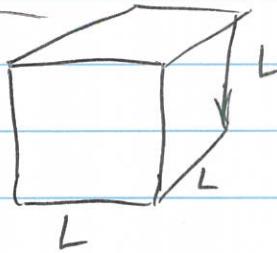
evaluate this part in the hope of pulling $C_e(t')$ term out of integral.

$$\text{Recall: } \mathcal{R}_\gamma = \frac{\langle g | H_{\text{int}} | e \rangle}{\hbar} = \frac{\langle g | -e \vec{R} \cdot \vec{E}_\gamma | e \rangle}{\hbar}$$



Cartoon quantization of the EM-field

Universe = Box with side L



$$\Rightarrow \text{volume} = V = L^3$$

$$\text{Energy of a photon in Box} = \hbar \omega_x = \frac{1}{2} \int \left(\epsilon_0 \vec{E}^2 + \frac{1}{\mu_0} \vec{B}^2 \right) dV$$

a little bit
of algebra

$$\Rightarrow |\vec{E}_x|_{\text{RMS}} = \sqrt{\frac{\hbar \omega_x}{\epsilon_0 V}}$$

$$\text{So far Energy} = \frac{\hbar \omega_x}{2} \Rightarrow |\vec{E}_{x, N=0}|_{\text{RMS}} = |\vec{E}_{\text{vacuum}}|_{\text{RMS}}$$

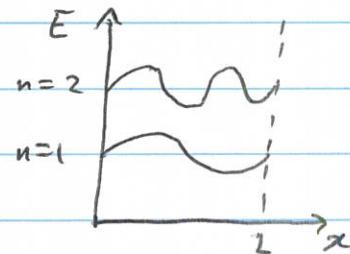
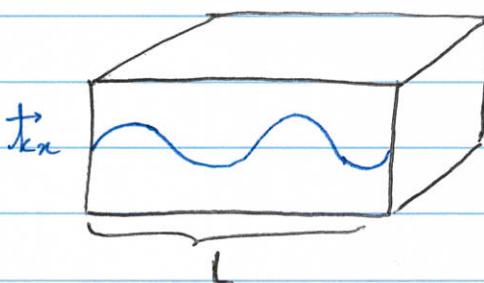
ideal vacuum has energy $\frac{\hbar \omega}{2}$

$$= \sqrt{\frac{\hbar \omega_x}{2 \epsilon_0 V}}$$

$$\Rightarrow |\vec{E}_{\text{vacuum}}| = \sqrt{\frac{\hbar \omega_x}{\epsilon_0 V}} \equiv |\vec{E}_x|$$

stopped here

Let's count the number of photon spatial modes in the box universe with periodic boundary conditions



$$\Rightarrow k_x = \frac{2\pi n_x}{L} \Rightarrow dn_x = \frac{L}{2\pi} dk_x$$

number of modes between k & $k+dk$

$$\Rightarrow dN = \left(\frac{L}{2\pi}\right)^3 dk^3 = dx dy dz$$

↓
2 polarizations

$$dN = 2 \frac{V}{(2\pi)^3} k^2 \sin \theta dk d\theta d\phi$$

\downarrow \downarrow
 ω_γ $\frac{d\omega_\gamma}{c}$

$$\Rightarrow dN_\gamma = 2 \frac{V}{(2\pi)^3} \frac{\omega^2}{c^3} \sin \theta dw d\theta d\phi$$

convert to integral

$$\text{Thus: } \frac{d}{dt} C_e(t) = - \sum_{\text{states}} \frac{1}{4} |S_\gamma|^2 \int_0^t e^{-i\delta_\gamma(t-t')} C_e(t') dt'$$

$$\frac{d}{dt} C_e(t) = - \int_0^\infty d\omega_\gamma \left\{ \frac{\omega_\gamma^2}{c^3} \frac{2V}{(2\pi)^3} \frac{|\mu_{eg}|^2 |E_\gamma|^2}{4 t^2} \int_0^\pi d\theta \sin^3 \theta \right\} d\omega_\gamma$$

$\frac{\omega_\gamma}{E_\gamma} \frac{\pi}{4\sqrt{3}}$ $\delta_\gamma = \omega_\gamma - \omega_{eg}$

$$x \int_0^t e^{-i\delta_\gamma(t-t')} C_e(t') dt'$$

$$\approx -\frac{1}{2} \left[\frac{\mu_{eg}^2 \omega_{eg}^3}{3\pi^2 \epsilon_0 h c^3} \right] C_e(t)$$

$\gamma = \frac{1}{2}$
 $= \text{spontaneous emission rate}$

Strongly peaked at $\begin{cases} \omega_\gamma = \omega_{eg} \\ t = t' \end{cases}$

$\pi \delta(\delta_\gamma) \tilde{\delta}(t-t') - P\left[\frac{i}{8}\right]$
 Lamb shift (neglect)

$$\Rightarrow \boxed{\frac{d}{dt} C_e(t) = -\frac{1}{2} \gamma C_e(t)}$$

decay is exponential with time constant $= \frac{2}{\gamma}$