

Tuesday, March 19, 2024

Spontaneous Emission (continued)

recall: $\frac{d}{dt} C_e(t) = - \sum_{\gamma} \frac{|\Omega_{\gamma}|^2}{4} \int_0^t e^{-\gamma_{\gamma}(t-t')} C_e(t') dt'$

convert to integral?

$\delta_{\gamma} = \omega_{\gamma} - \omega_{eg}$

photon states $(\omega_{\gamma}, k_{\gamma}, \text{polarization})$

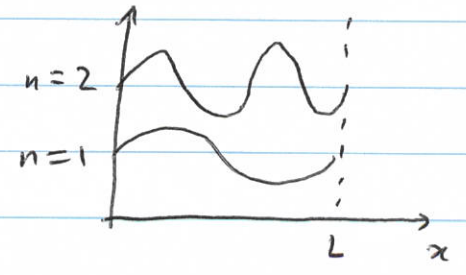
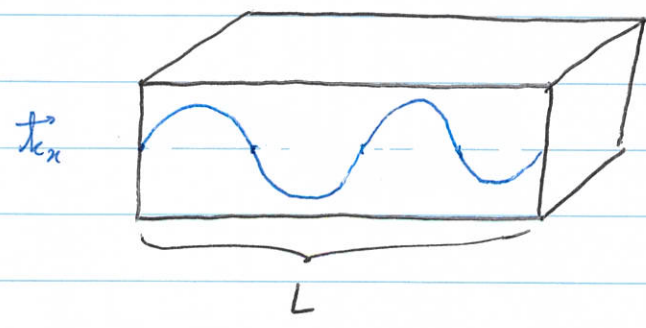
objective: rewrite above equation as $\frac{d}{dt} C_e(t) = - \left[\frac{\gamma}{2} \right] C_e$

where $\gamma = \text{decay rate} = 1/\tau$

$\tau = \text{decay time}$

Converting \sum_{γ} to an integral (sum over γ -photon states)

Let's count the number of photon spatial modes in a box universe with periodic boundary conditions



$\Rightarrow k_x = \frac{2\pi}{L} n_x \Rightarrow dn_x = \frac{L}{2\pi} dk_x$

number of modes between k_x & $k_x + dk_x$

$$\Rightarrow dN = \left(\frac{L}{2\pi}\right)^3 dk^3 = dn_x dn_y dn_z$$

2 polarizations

$$dN = 2 \frac{V}{(2\pi)^3} k^2 \sin\theta dk d\theta d\varphi$$

\uparrow $k = \frac{\omega_r}{c}$ \uparrow $\frac{d\omega_r}{c}$

$$\Rightarrow dN_r = 2 \frac{V}{(2\pi)^3} \frac{\omega_r^2}{c^3} \sin\theta d\omega_r d\theta d\varphi$$

convert to integral

Thus: $\frac{d}{dt} C_e(t) = - \sum_{r \text{ states}} \frac{|\mathcal{R}_r|^2}{4} \int_0^t e^{-i\delta_r(t-t')} C_e(t') dt'$

recall: $\mathcal{R}_r = -\frac{1}{\hbar} |\vec{\mu}_{eg}| |\vec{E}_r| \sin\theta \Rightarrow |\mathcal{R}_r|^2 = \frac{|\mu_{eg}|^2}{\hbar} \frac{\omega_r \sin^2\theta}{\epsilon_0 V}$

$$\Rightarrow \frac{d}{dt} C_e(t) = - \int_0^\infty d\omega_r \left\{ \frac{2V}{(2\pi)^3} \frac{\omega_r^2}{c^3} \frac{|\mu_{eg}|^2}{4\hbar} \frac{\omega_r}{\epsilon_0 V} \int_0^\pi d\theta \sin^3\theta \int_0^{2\pi} d\varphi \right\}$$

$$\times \int_0^t e^{-i\delta_r(t-t')} C_e(t') dt'$$

$\delta_r = \omega_r - \omega_{eg}$

Strongly peaked at $\begin{cases} \omega_r = \omega_{eg} \\ t = t' \end{cases}$

$$\rightarrow \pi \delta(\delta_r) \tilde{\delta}(t-t') = \mathcal{P} [1/\delta]$$

(Lamb shift (neglect))

$$\Rightarrow \frac{d}{dt} C_e(t) = - \frac{\cancel{2\pi} \cancel{\hbar} \omega_{eg}^3}{(2\pi)^{\cancel{3/2}} c^3} \frac{\mu_{eg}^2}{\cancel{4\hbar}} \frac{1}{\cancel{\epsilon_0} \cancel{V}} \frac{\cancel{4}}{3} \cancel{2\pi} \cancel{\hbar} C_e(t)$$

$$= -\frac{1}{2} \underbrace{\left(\frac{\mu_{eg}^2 \omega_{eg}^3}{3\pi \epsilon_0 \hbar c^3} \right)}_{\gamma = 1/\tau} C_e(t)$$

= spontaneous emission rate

$$\Rightarrow \boxed{\frac{d}{dt} C_e(t) = -\frac{1}{2} \gamma C_e(t)} \quad \text{decay of } C_e \text{ is exponential with time constant} = \frac{2}{\gamma}$$

In "Schrodinger format", the equation becomes

$$i\hbar \frac{d}{dt} C_e(t) = -i\hbar \frac{\gamma}{2} C_e(t)$$

We would like to write the Schrodinger equation of the dressed atom with spontaneous emission as:

$$i\hbar \frac{d}{dt} \begin{pmatrix} C_g(t) \\ C_e(t) \end{pmatrix} = \hbar \begin{pmatrix} \delta & \Omega/2 \\ \Omega^*/2 & 0 \end{pmatrix} \begin{pmatrix} C_g(t) \\ C_e(t) \end{pmatrix} - \begin{pmatrix} 0 (?) \\ i\hbar \frac{\gamma}{2} C_e(t) \end{pmatrix}$$

We are not sure what to put here because as spontaneous emission depletes the excited state $|e\rangle$, atoms go into the ground state $|g\rangle$, but not coherently [the environment chooses a direction for the emitted photon]
 "environment performs a measurement"

$$\Rightarrow i\hbar \frac{d}{dt} \begin{pmatrix} c_g(t) \\ c_e(t) \end{pmatrix} = \hbar \begin{pmatrix} \epsilon & \Omega/2 \\ \Omega^*/2 & -i\gamma/2 \end{pmatrix} \begin{pmatrix} c_g(t) \\ c_e(t) \end{pmatrix}$$

$\underbrace{\hspace{10em}}_H$

imaginary term

↳ non-Hermitian

⇒ probability is not conserved !!



The Schrodinger equation can only handle coherent superpositions of states, not statistical mixtures.

Two methods for dealing with statistical mixtures:

- Density matrix

- Monte Carlo wavefunctions

or

"Quantum trajectories"

Density Matrix

(Von Neumann, Landau, Bloch)

definition: The density matrix ρ for a system with wavefunction $|\psi\rangle$ is given by

$$\rho = |\psi\rangle\langle\psi|$$

If $|\psi\rangle = c_g|g\rangle + c_e|e\rangle$, then

$$\rho = |\psi\rangle\langle\psi| = \begin{pmatrix} c_g c_g^* |g\rangle\langle g| & c_g c_e^* |g\rangle\langle e| \\ c_e c_g^* |e\rangle\langle g| & c_e c_e^* |e\rangle\langle e| \end{pmatrix}$$

For a population in a statistical mixture of states $|\psi_i\rangle$ with probabilities p_i , then

$$\rho = \sum_i p_i |\psi_i\rangle\langle\psi_i|$$

← allows for quantum superpositions and statistical mixtures simultaneously.

Properties:

Expectation value: $\langle A \rangle = \langle \psi | A | \psi \rangle = \text{Tr}(\rho A)$

time evolution:

$$i\hbar \frac{d}{dt} \rho = [\mathcal{H}, \rho]$$

notation:

$$\rho = \begin{bmatrix} \rho_{gg} & \rho_{ge} \\ \rho_{eg} & \rho_{ee} \end{bmatrix} = \begin{bmatrix} c_g c_g^* & c_g c_e^* \\ c_e c_g^* & c_e c_e^* \end{bmatrix}$$

Resonance Fluorescence - optical Bloch Equations

The great advantage of the density matrix approach is that the decoherence that accompanies spontaneous emission can be added in easily

coherent evolution: $i\hbar \frac{d}{dt} \rho = [H, \rho]$

with $H = \begin{pmatrix} E_g & \hbar\Omega \cos(\omega_L t) \\ \hbar\Omega^* \cos(\omega_L t) & E_e \end{pmatrix}$
 (don't forget RWA)

or $H_{\text{dressed}} = \begin{pmatrix} S & \hbar\Omega/2 \\ \hbar\Omega^*/2 & 0 \end{pmatrix}$

$$\frac{d}{dt} \rho_{gg} = -\frac{i}{\hbar} [H, \rho]_{gg} + \gamma \rho_{ee}$$

$$\frac{d}{dt} \rho_{ee} = -\frac{i}{\hbar} [H, \rho]_{ee} - \gamma \rho_{ee}$$

$$\frac{d}{dt} \rho_{eg} = -\frac{i}{\hbar} [H, \rho]_{eg} - \frac{1}{2} \gamma \rho_{eg}$$

$$\frac{d}{dt} \rho_{ge} = -\frac{i}{\hbar} [H, \rho]_{ge} - \frac{1}{2} \gamma \rho_{ge}$$

\Rightarrow solve for $\rho_{gg}, \rho_{ee}, \rho_{eg}, \rho_{ge}$