Today's Topics

Monday, September 9, 2019 (Week 2, lecture 5) – Chapter 3.

0. Kepler's 3rd law

- 1. Galileo & Newton
- 2. Newton's laws
- 3. Conservation laws
- 5. Gravity

Participation Quiz

2 minutes

0 point – no answer 1 point – wrong answer 2 points – correct answer



$0.930 imes 10^3 \mathrm{~m}$

$93.0 imes10^9~{ m cm}$

$93.0 imes 10^4 \mathrm{~m}$

Kepler's 3rd Law

T = orbital period in units of Earth years

a = semimajor axis in AU

$T^2 = a^3$

Examples

- 1. Martian orbit.
- 2. Orbital speed vs orbital radius.

Galileo Galilei: Birth of Classical Mechanics

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- Contributed to physics, astronomy, optics, engineering.
- Confronted Catholic Inquisition over **heliocentrism**.



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Physics contributions – classical mechanics

- Galilean Relativity
 - \rightarrow Objects in uniform motion tend to stay in motion.
- Objects fall with a parabolic trajectory.



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Astronomy contributions

- > Key developer of the **telescope** for astronomy.
- Discovered the moons of Jupiter.
- Discovered the *phases of Venus* (similar to Moon phases).
- Proponent of heliocentric view.



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Acceleration

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Examples:

- 1. A car's acceleration is advertised as "0-100 km/h in 5 seconds."
- 2. Acceleration due to gravity is g = 9.8 m/s per second= 9.8 m/s²

Constant speed: x = vt

[x = position, v = speed, t = time (elapsed)]

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Dropped Object



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[a = acceleration]

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Dropped Object



Galilean Relativity

<u>Definition</u> An **inertial frame** is a *coordinate system* moving at constant velocity. [constant velocity = constant speed & constant direction]

→ Inertial frame = space that travels with you, e.g. car, airplane, rocket, etc ...
 → Note: an accelerating/rotating system is NOT an inertial frame.

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Galilean relativity posits that in any <u>inertial frame</u>:

"you cannot tell that you are moving based on local measurement."

 \rightarrow i.e. an inertial frame locally behaves as if it is at rest (locally).

 \rightarrow corollary: an object in uniform motion will tend to stay in uniform motion.

Examples:

- 1. Car: You cannot tell that a car is moving (when at constant velocity) unless you look out window.
- 2. Airplane: You cannot tell an airplane is moving (when at constant velocity) unless you look out window (or hit turbulence).

Isaac Newton: Founder of Classical Mechanics



Newton (1689) [by G. Kneller]

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Classical Mechanics

- "Newton's Laws" of classical mechanics.
- Law of universal gravitation.
- Newton's laws are used for calculating planetary & stellar motion. (+ Einstein's "Special Relativity")

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Astronomy

- **Optics:** white light &colors, refraction.
- Invented the reflecting telescope.

Newton's Laws

of Classical Mechanics

1st Law: An object moves at constant velocity if there is no net force acting on it.

[fine print: in an inertial reference frame]

2nd Law: Force = mass × acceleration.

3rd Law: For any force, there is always an equal and opposite reaction force.

Newton's 1st Law

An object moves at constant velocity if there is no net force acting on it.

[fine print: in an inertial reference frame]

Note: This law is a variation on the <u>Galilean relativity</u> statement.

Newton's 2nd Law

Force = Mass × Acceleration

or

F = max F = net forcem = massa = acceleration

[fine print: in an inertial reference frame]

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Note 1: This equation is mostly useful if you know the net force applied.

Note 2: If the acceleration is zero, then the net force is zero.

For any force, there is always an equal and opposite reaction force

$$F_{A \to B} = -F_{B \to A}$$

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Example:

$$A = box$$
$$F_A \equiv gravity$$

B = platform

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Newton's 3rd Law: Rocket Thrust

A rocket accelerates by pushing on its exhaust.



A rocket does NOT push on the air to accelerate. A rocket does NOT push on its platform to accelerate.

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Conservation of Momentum

momentum = mass × velocity

total momentum = sum of the momenta of all the sub-parts of a system

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Conservation Law

The total momentum of a closed system never changes.

no external objects enter no external forces

Momentum Conservation: Rocket Thrust

 $Momentum_{rocket} + Momentum_{exhaust} = 0$



Conservation of Energy

Kinetic Energy =
$$E_k = \frac{1}{2}mv^2$$

 $m = mass$
 $v = speed$

Potential Energy = "stored" energy

example: gravitational potential energy

Total Energy = sum of the energies of all the sub-parts of a system

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Properties

- Falls off as $1/r^2$.
- Proportional to M_A .
- Proportional to M_B .
- G = Newton's constant = $6.67430(15) \times 10^{-11}$ $m^3/Kg \cdot s^2$