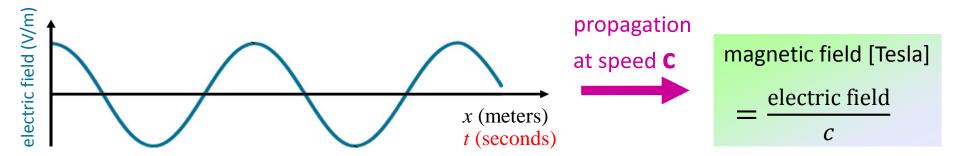
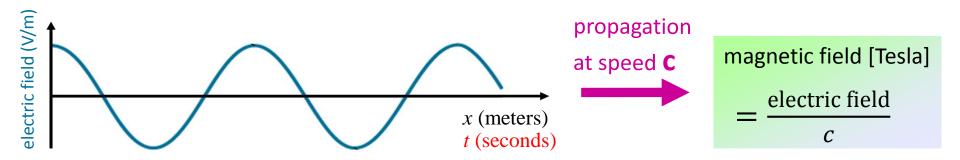
Today's Topics

Wednesday, September 18, 2019 (Week 3, lecture 9) – Chapter 5.

- 1. Light basics: EM waves & photons
- 2. Electronic structure of atoms
- 3. Spectroscopy
- 4. Doppler effect

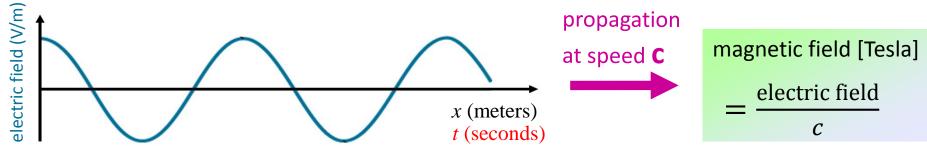




Intensity = power/area =
$$I = \frac{1}{2}c\varepsilon_0 E^2$$
 [W/m²] $E = \text{electric field [V/m]}$ $\varepsilon_0 = \text{permittivity of vaction}$

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 $= 8.85 \times 10^{-12} \, \text{C}^2/(\text{N} \cdot \text{m}^2)$



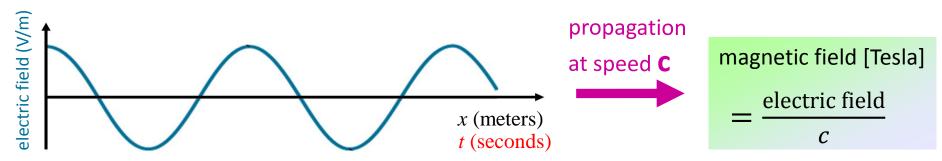
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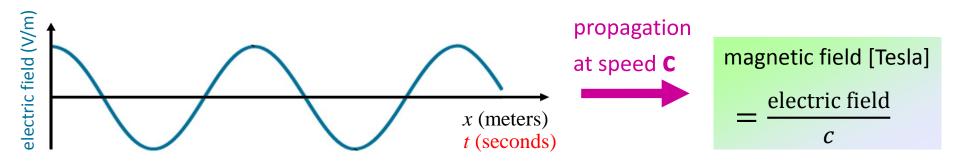


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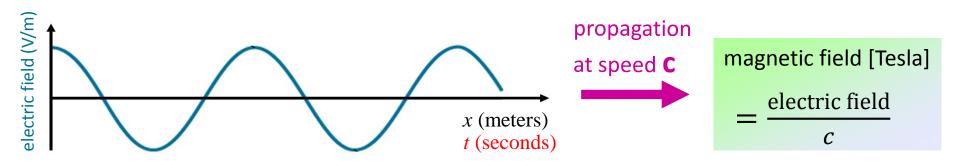
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Answer: oscillate charge (or accelerate it).

→ generates "dipole radiation."







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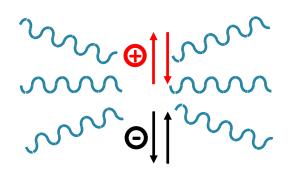
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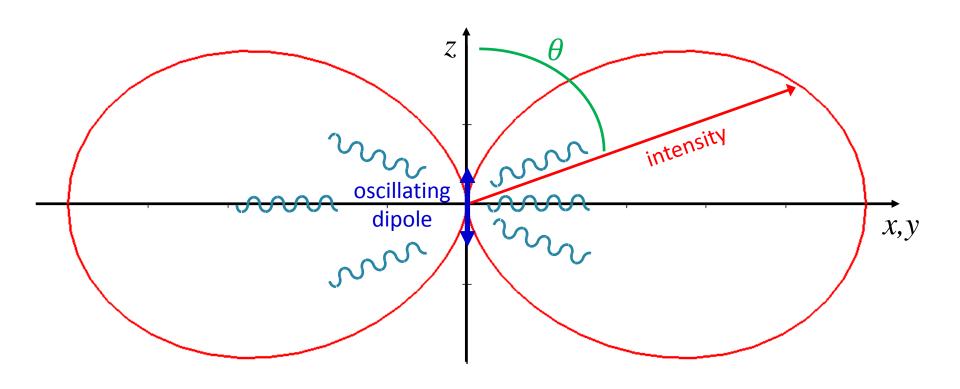
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Dipole Radiation Pattern

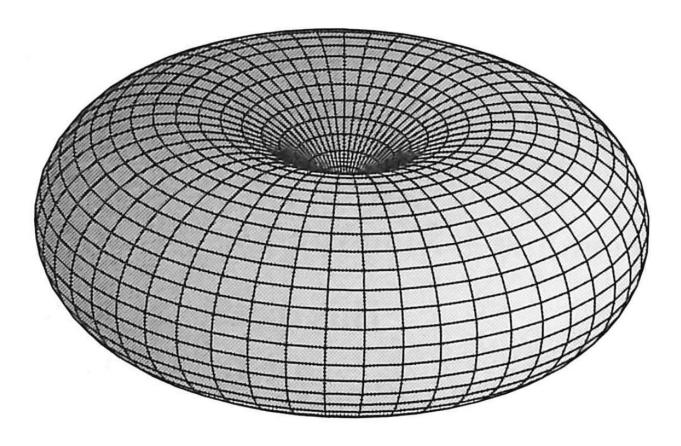
dipole moment = p_0 = charge \times separation



$$Intensity = \frac{\pi^2 p_0^2}{2\epsilon_0 c^3} \cdot f^4 \cdot \frac{\sin^2 \theta}{r^2} \left[\propto f^4 \frac{1}{r^2} \right]$$

r =distance from dipole f =frequency

Dipole Radiation Pattern



[Figure 11.4, Introduction to Electrodynamics, by D. Griffiths, 4th Ed.]

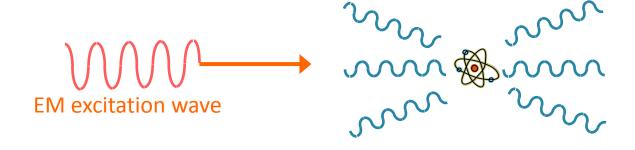
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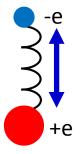
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Dipole Radiation Example #1 Atomic fluorescence & photon scattering

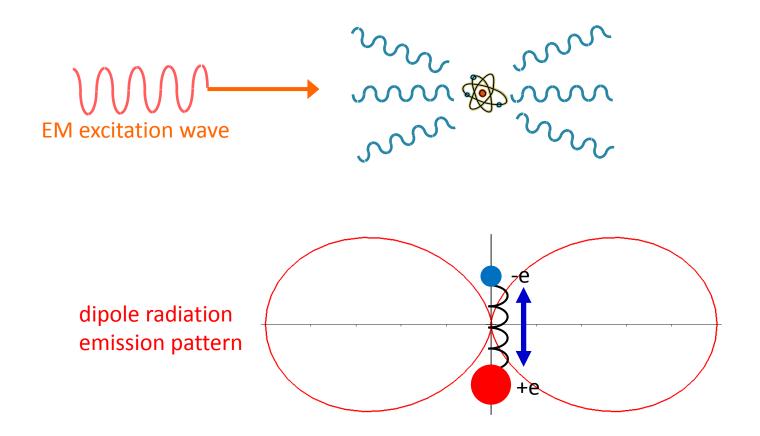
Rayleigh scattering: an atom behaves like a perfect electric dipole when excited by an EM wave.





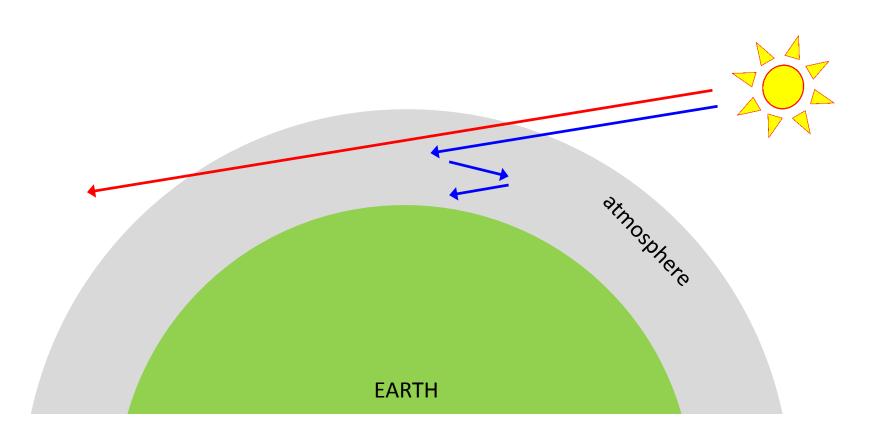
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Dipole Radiation Example #2 Blue Sky

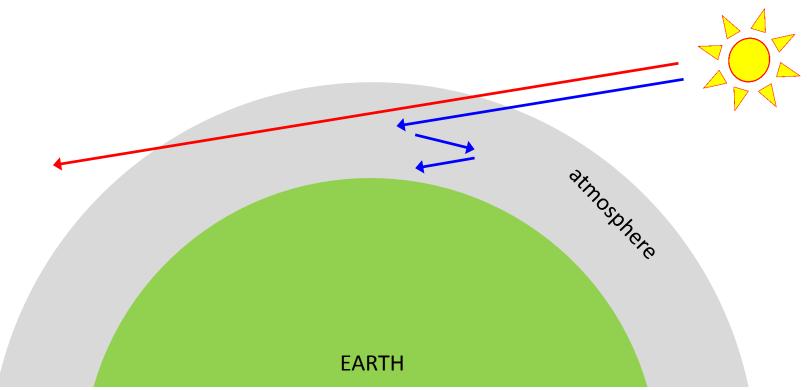
Blue light scatters at a higher rate than red light → Sky looks blue.



Dipole Radiation Example #2 Blue Sky

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$$Intensity \propto f^4 \propto \frac{1}{\lambda^4}$$

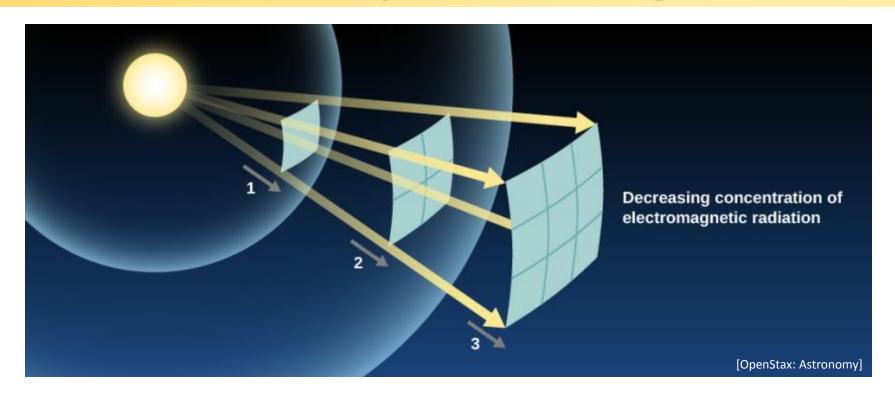


Dipole Radiation Example #2 Blue Sky

Blue light scatters at a higher rate than red light → Sky looks blue.

Intensity
$$\propto f^4 \propto \frac{1}{\lambda^4}$$
 \Longrightarrow $\lambda_{\text{blue}} = 450 \text{ nm}$ $\frac{I_{blue}}{I_{red}} = \left(\frac{650}{450}\right)^4 \approx 4.3$

Inverse Square Law for Light



- \succ As light radiates away from its source, it spreads out such that its intensity decreases as the **square** of the **distance** d from its source.
- > Intensity $\propto 1/d^2$.

Photon energy =
$$E_{\gamma} = hf$$

Important: Photon is massless
$$M_{\gamma} = 0$$

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$$p_{\gamma} = \frac{h}{\lambda} = \frac{E_{\gamma}}{c}$$

Note:
$$p_{\gamma} \neq M_{\gamma}c \ (=0)$$

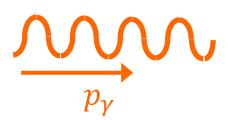
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Shine light on an atom



$$p_{atom} = 0$$

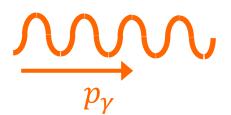
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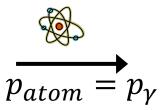
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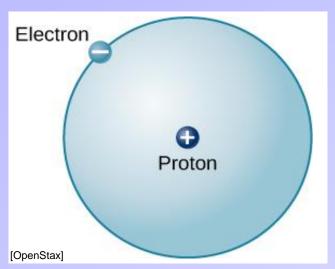
Atom after absorption of 1 photon



(conservation of momentum)

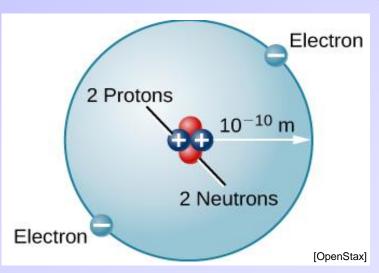
Basic Structure of Atoms

Hydrogen: ¹H



Helium: ⁴He

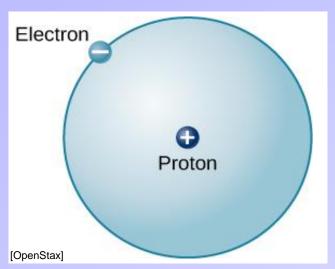
4 = 2 protons + 2 neutrons



- Atom consist of positively charged nucleus orbited by negatively charged electrons (for neutral atoms: total charge = zero).
- > Electron number, orbits, and properties determine the **chemistry** of the atom.

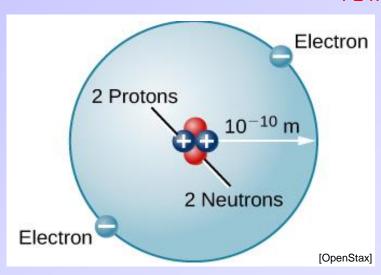
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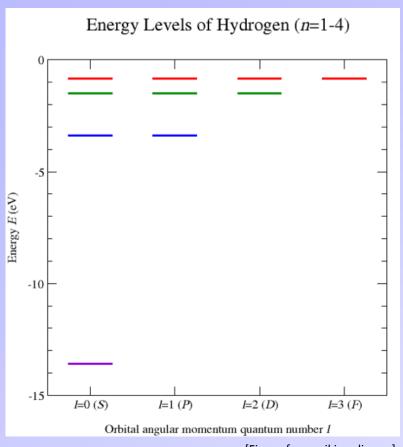
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- Atom consist of positively charged nucleus orbited by negatively charged electrons (for neutral atoms: total charge = zero).
- Electron number, orbits, and properties determine the chemistry of the atom.
- Nucleus consists of positively charged protons and neutral neutrons.
- For neutral atoms: Number of protons = number of electrons.
- ➤ Neutrons help bind protons together. Number of neutrons ≥ number of protons.

Electronic

Structure of Atoms



[Figure from wikimedia.org]

N = 4

N = 5

Blue-green spectral line orbit line

Violet spectral line

N = 3

Red spectral line

Red spectral line

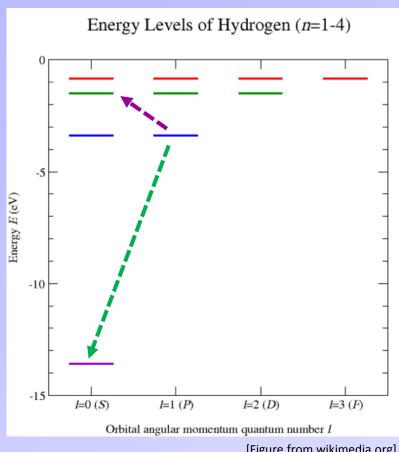
[OpenStax]

Electrons have discrete allowed energies and orbits.

Note: $1 \text{ eV} = 1.602 \times 10^{-19} \text{ J}$

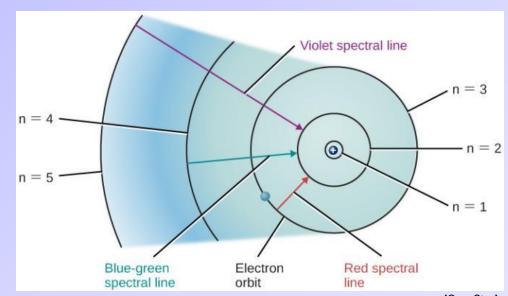
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[Figure from wikimedia.org]

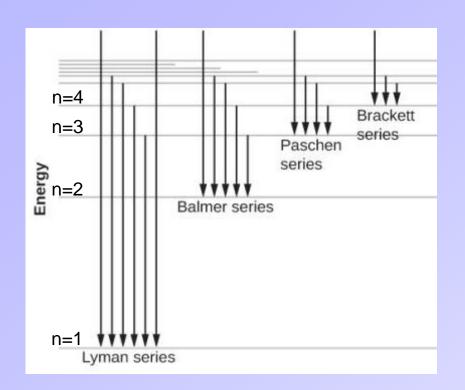
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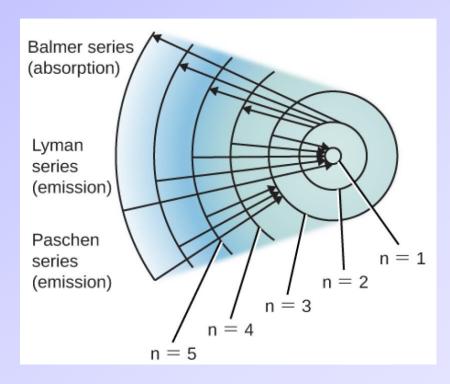


[OpenStax]

- Electrons have discrete allowed energies and orbits.
- Transitions between two energy levels requires emission or absorption of a photon that bridges the energy gap.
- **Discrete** emission and absorption **spectra**.

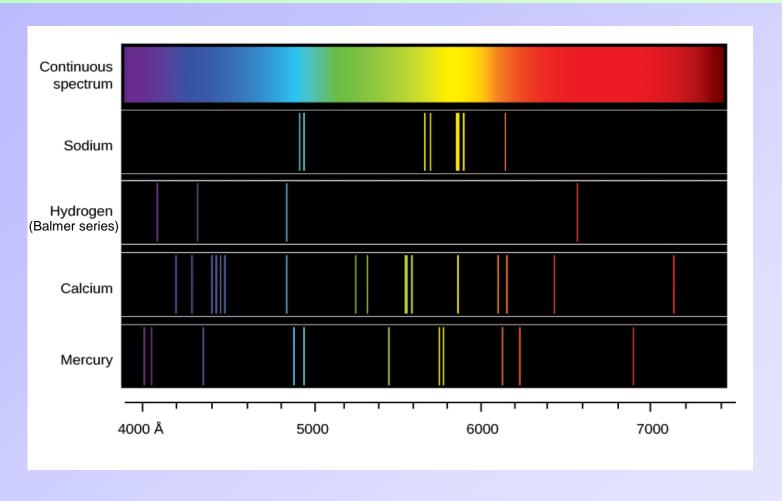
Emission Spectrum of Hydrogen





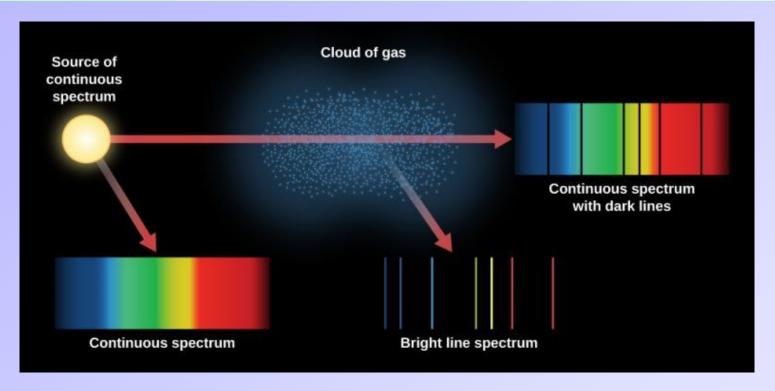
- Hydrogen has a number of emission and absorption spectral series that depend on the start/end point of the transition.
- Other elements are qualitatively similar.
- Also true for molecules, but their spectra are more complicated.

Emission Spectra "Fingerprints"



If you build a catalog of spectral lines, then you can determine the elements that are present from the spectrum.

Emission & Absorption Processes



Three types of spectra

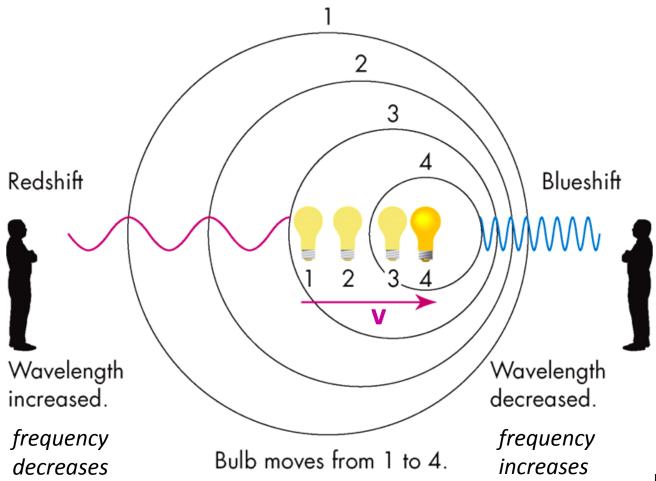
- Continuous spectrum, e.g. a thermal blackbody source.
- 2. Emission spectrum (discrete): if light excites atoms, then the atomic emission will be at discrete frequencies.
- 3. Absorption spectrum (discrete): if a continuous spectrum excites atoms, then the absorption of photons will remove light at discrete frequencies ("shadow lines").

Doppler Effect

A moving source cannot change the speed of its emitted light, but it does change its frequency & wavelength.

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[image source: J. Nelson]

Doppler Effect

A moving source cannot change the speed of its emitted light, but it does change its frequency & wavelength.

Works for sound too !!!

Doppler Shift Calculation

Doppler frequency shift:
$$\frac{\Delta f}{f} = -\frac{\Delta \lambda}{\lambda} = \frac{v}{c} \quad \text{with } f' = f + \Delta f$$
 frequency of stationary source of moving source

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If source is moving towards you, then light is blue shifted.

$$v>0$$
 $\Delta f>0$, f' goes up $\Delta \lambda < 0$

If source is moving away from you, then light is red shifted.

$$v < 0$$
 $\Delta f < 0$, f' goes down $\Delta \lambda > 0$