### **Today's Topics**

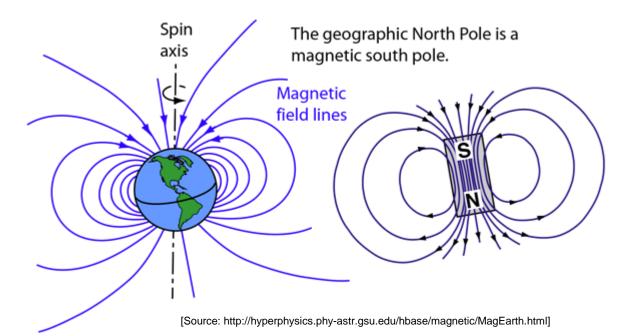
Friday, October 11, 2019 (Week 6, lecture 18) – Chapter 8.

## 1. Magnetosphere

2. Atmosphere

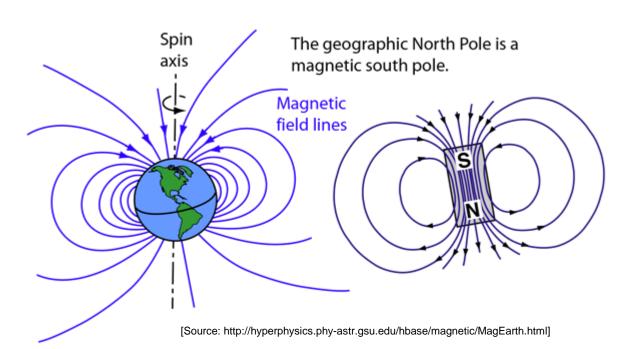
## **Earth's Magnetic Field**

- Earth has a <u>magnetic field</u> generated by electrical current in its core.
- The magnetic is <u>not</u> aligned with Earth's rotation axis.
- ➢ 0.3-0.5 Gauss at surface.

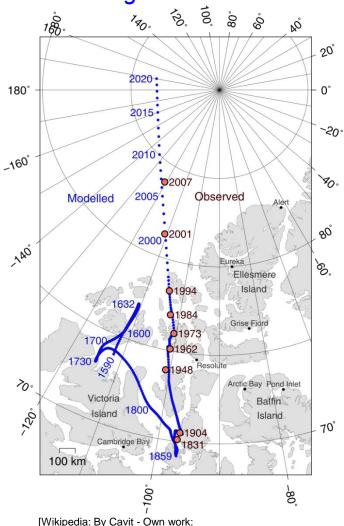


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- The magnetic pole drifts over time and flips on a time scale of 0.1 – 1 million years.

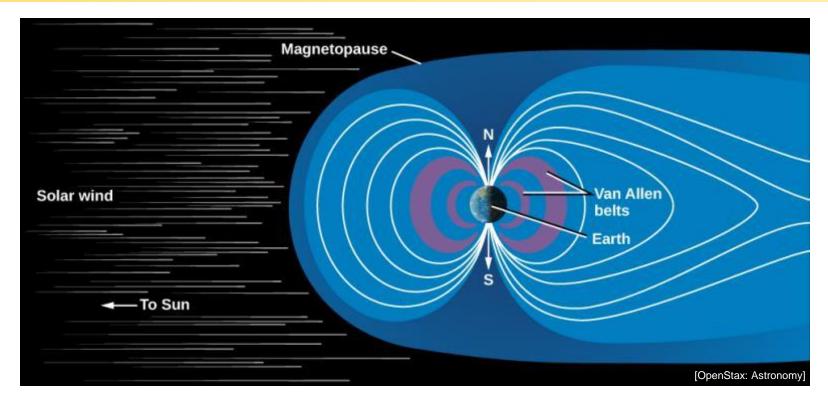


#### Earth's magnetic north vs time



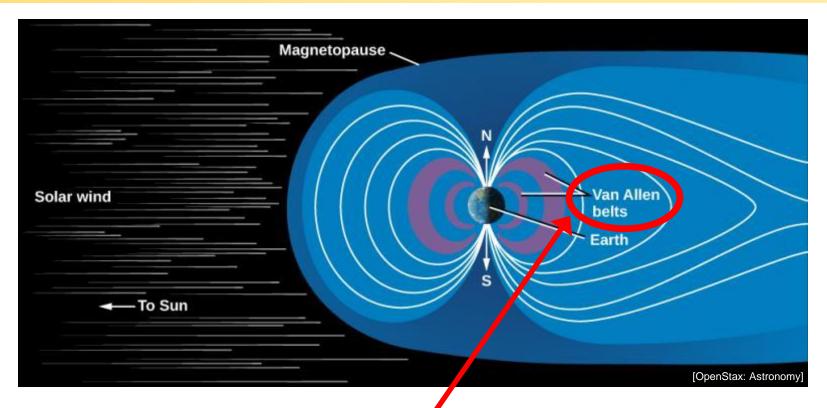
Observed pole positions taken from Newitt et al., "Location of the North Magnetic Pole in April 2007", Earth Planets Space, 61, 703–710, 2009 Modelled pole positions taken from the National Geophysical Data Center, "Wandering of the Geomagnetic Poles"Map created with GMT, CC BY 4.0]

## **Magnetosphere**



- Earth's magnetic field screens the planet from charged particles emitted by the Sun (i.e. solar wind).
- The Earth's magnetic field deflects the charged particles into spiral trajectories and slows them down.

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Charged particles are trapped by magnetic field in the Van Allen radiation belts.

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## **Aurora Borealis**



#### **Basic physics**

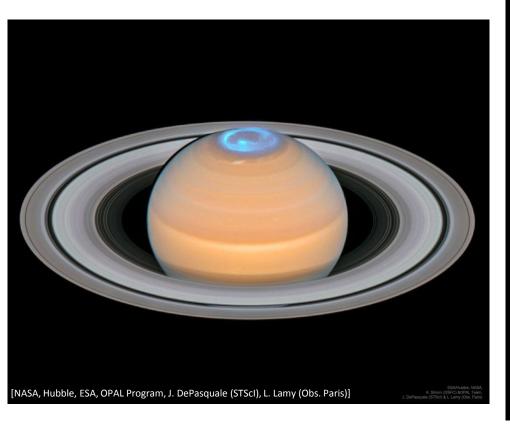
- Solar wind charged particles are directed by the Earth's magnetic field into the atmosphere.
- Atmosphere molecules/atoms are ionized, excited, and generate light (red: H, green: O).

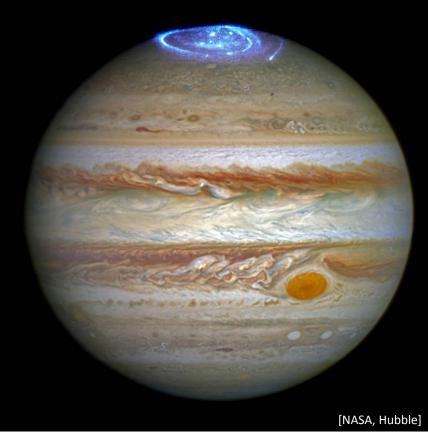
### **Aurora Australis**

NASA IMAGE satellite Aurora Australis (south pole)

### **Aurora on Jupiter and Saturn**

Hubble Space telescope images: UV image overlaid on an optical image





## **Earth's Atmosphere**

The atmosphere forms the outer layer of the planet.

- $\rightarrow$  It's what you see from outer space.
- $\rightarrow$  It's where most of life is (plus oceans).
- $\rightarrow$  It's where we live.

### **Primary Composition**

78 % nitrogen gas  $(N_2)$ 21 % oxygen gas  $(O_2)$ 1 % argon gas (Ar) 0.04% carbon dioxide gas  $(CO_2)$ 0-4% water vapor  $(H_2O)$  – variable

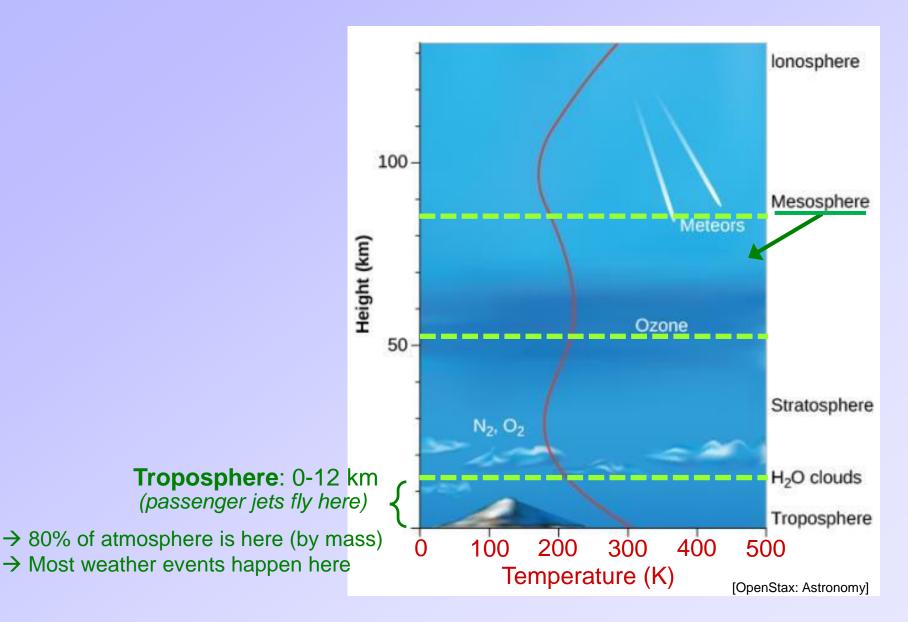
#### **Trace gases**

Neon (Ne) Helium (He) Methane ( $CH_4$ ) Krypton (Kr) Ozone ( $O_3$ )

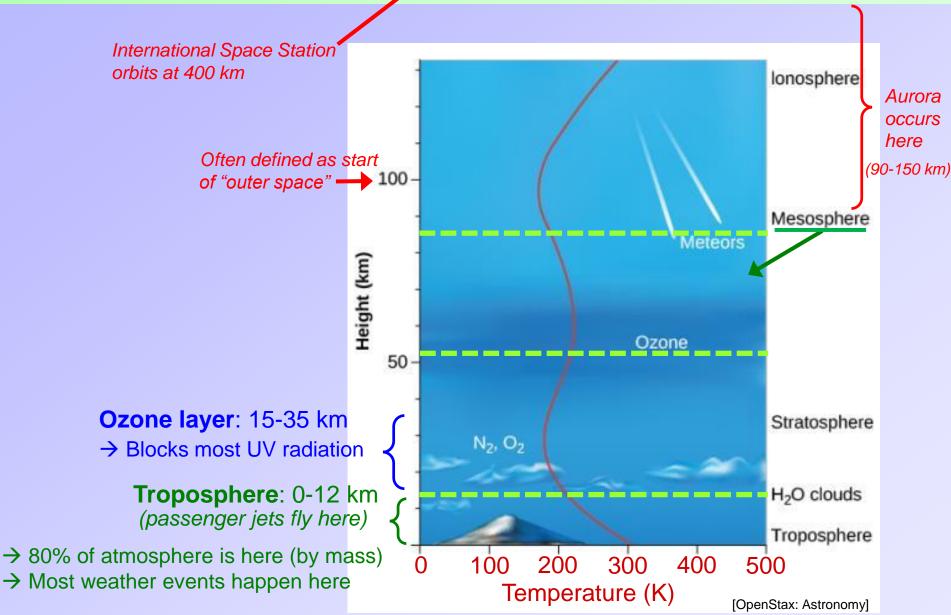


Earth as seen from Apollo 17

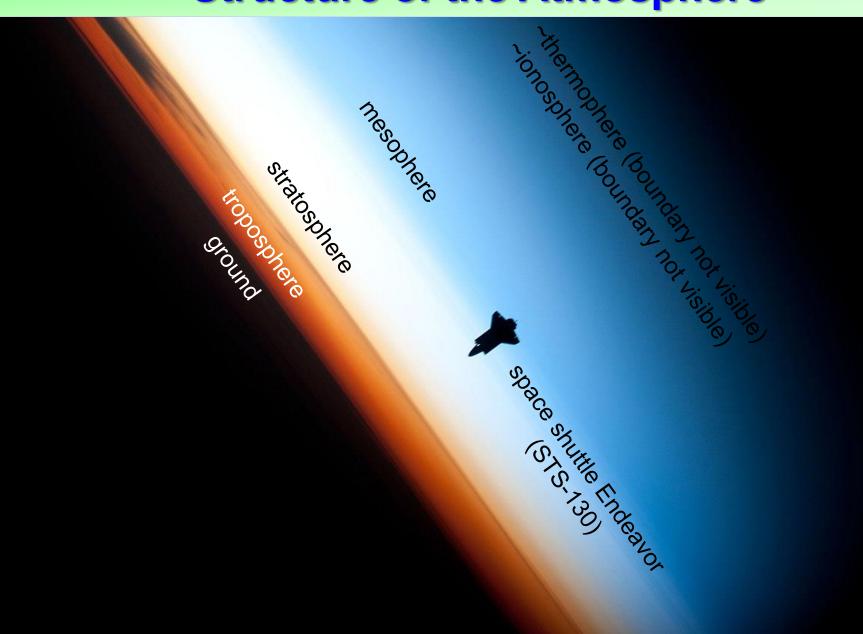
## **Structure of the Atmosphere**



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[NASA, International Space Station]

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Radioactive decay of potassium-40.
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Tom Greene, CW foundation



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### Water (H<sub>2</sub>O): 0-4 %

Water vapor comes primarily from *ocean evaporation*. → Water is mildly reactive, excellent catalyst.

#### **Ozone (O<sub>3</sub>): trace quantities**

Ozone is created-destroyed by UV light:  $O_2 + UV \rightarrow 2O$ ,  $O_2 + O \rightarrow O_3$ ,  $O_3 + UV \rightarrow O_2 + O$ 











## **Comparison with Venus & Mars**

	[NASA] Venus	Earth	IESAJ Mars
Nitrogen	3.5 %	78 %	2.6 %
Oxygen	trace	21 %	0.17 %
Argon	0.005 %	1 %	1.9 %
Carbon dioxide	96.5 %	0.04 %	95 %
water	trace	0-4 %	0.03 %

Earth's atmosphere is **<u>not</u>** CO<sub>2</sub> because of *life* 

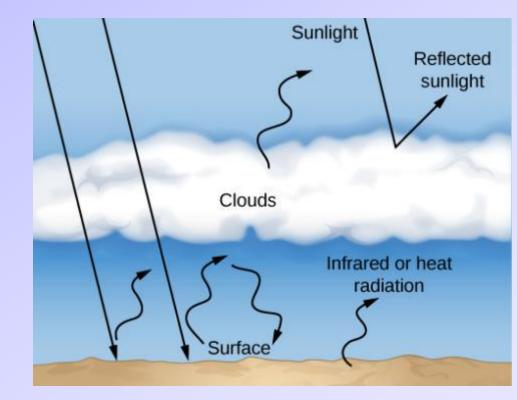
The Martian and Venusian atmospheres are dominated by carbon dioxide.

## **Atmospheric Temperature**

## **The Greenhouse Effect**

### How it works:

- Sunlight penetrates atmosphere and *heats surface*.
- The heated surface re-radiates in the infrared.



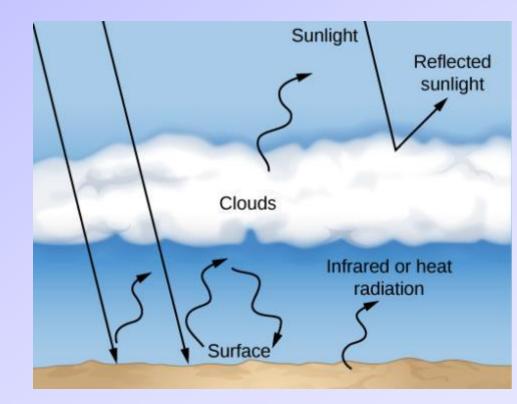
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Note: **Clouds** also help block the re-radiation of infrared radiation.



# **Atmospheric Temperature**

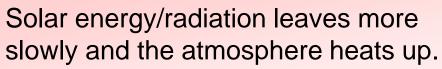
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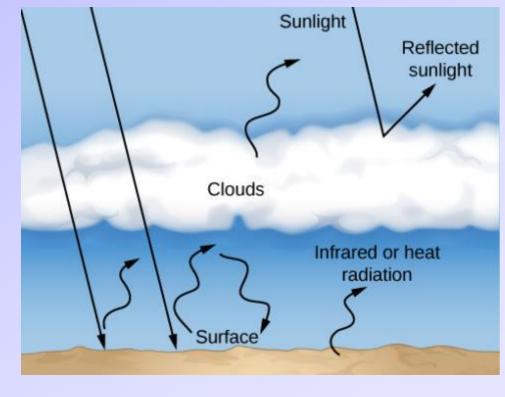
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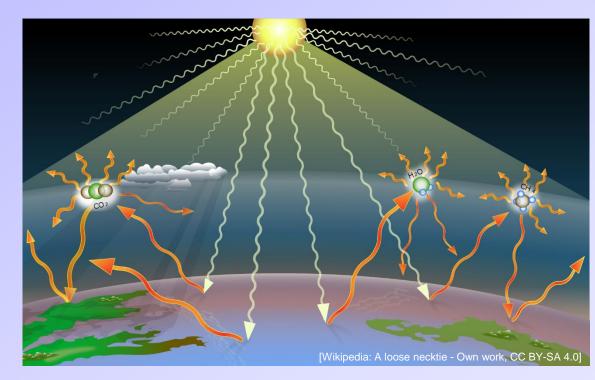




## **Greenhouse Gases**

### Primary Greenhouse Gases on Earth

- Water, H<sub>2</sub>O (+ clouds)
   → contribution: 36-72 %
- Carbon dioxide, CO<sub>2</sub>
   → contribution: 9-26 %
- Methane, CH<sub>4</sub>
   → contribution: 4-9 %
- Ozone, O<sub>3</sub>
  - $\rightarrow$  contribution: 3-7 %



#### Other greenhouse gases

- Nitrous oxide (N<sub>2</sub>O)
- Chlorofluorocarbons (CFCs).

## **The Greenhouse Effect**

### comparison with Mars & Venus

	[NASA] Venus	Image: Constraint of the second se	ESAJ Mars	
Temperature <b>with</b> greenhouse effect	470° C	15° C	– 50° C	
Temperature <b>without</b> <u>greenhouse effect</u> (estimate)	- 40° C	– 16° C	– 56° C	
	low temperature predic because of <u>high albedo</u> i.e. it's fairly reflective	o. Gre	Greenhouse effect is small because Mars has a thin atmosphere.	