Week 3 Light & Matter

1. Electromagnetic waves & photons

2. Spectroscopy and atoms

3. Particles, nuclei, and fusion

REMINDER: Midterm #1 is on Monday, September 14.

Today's Topics

Friday, September 4, 2020 (Week 3, lecture 8) – Chapter 5.

A. Electromagnetic waves

- B. Electromagnetic spectrum
- C. Blackbody radiation
- D. Photons

Speed of Light

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The speed of light <u>does NOT depend</u> on the observer:

- If observer A is at rest and measures the speed of light of their laser pointer, then they will measure $c = 3.0 \times 10^8$ m/s.
- If observer B is moving at 290,000 km/s, then they will measure the speed of light of <u>observer A's laser pointer</u> to be c = 3.0 × 10⁸ m/s.

Speed of Light in Matter

The speed of light *in matter is slower* than in vacuum

Speed of light in air = 99.97% of c

Speed of light in water = 75% of c

Speed of light in glass = 67% of c

Speed of light in **diamond = 41% of c**

Speed of light in silicon $\simeq 25\%$ of c



[123RF.com]

Note: In engineered atomic gases, light can be brought ~ 10 m/s and even stopped. (Novikova Lab at W&M)

Light:

Particle or Wave?



Electromagnetic Waves

James Clerk Maxwell (1831-1879) worked on electricity and magnetism:

- They are different facets of the same phenomenon.
- Light is a wave of electric & magnetic fields.



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oscillating electric field

oscillating magnetic field
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Wave Properties



Frequency: $f = \frac{1}{T}$ = oscillations per second

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Wave Properties



Wave Addition: Constructive Interference



Wave Addition: Constructive Interference



Wave Addition: Destructive Interference



Wave Addition: Destructive Interference





Screen





Screen





[Young's double slit experiment, by Thomas Young (1773-1829) in 1801-1803.]



[[]data by M. Frayser, W&M 2018]

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Electromagnetic Spectrum

- Visible light represents only a small portion of electromagnetic waves.
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Astronomers use all Wavelengths

Crab Nebula (M1)

- Exploding star remnant (superonova).
- Recorded by Chinese astronomers and others (1054 AD).
- Located at about 6500 ly in our galaxy (Taurus constellation).
- This composite image is by the Hubble Space Telescope (visible light).



Crab Nebula with Radio-Waves



Radio (Very Large Array)

[VLA/NRAO/AUI/NSF]

Crab Nebula with Infrared Light



[NASA/Spitzer/JPL-Caltech]

Crab Nebula with Visible Light



[NASA, ESA, and Hubble (STScI)]

Crab Nebula with Ultraviolet Light



[XMM-Newton/ESA]

Crab Nebula with X-Rays



X-ray (Chandra)

[NASA/Chandra/CXC]

Absorption by Earth's Atmosphere



Thermal Light Sources Blackbody Radiation

- The oldest and simplest way to make light is by heating something up (filament, gas, wood, etc).
- Hotter = brighter, colder = dimmer.
- Hotter = white-blue, colder = dim red.
- Color of thermal source \rightarrow temperature.



incandescent lightbulb

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Ideal thermal source of light

Blackbody Radiation (1)



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Blackbody Radiation (2)

- Total output power (per unit area)
 = area under the curve
 - = Luminosity (L)
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$$L = \sigma T^4$$

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Stefan-Boltzman Law:
$$L = \sigma T^4$$
 Increasing temperature, increases output power a lot *Stefan-Boltzman constant:* $\sigma = 5.67 \times 10^{-8} \frac{W}{m^2 K^4}$

Light is also a particle: the Photon

- Max Planck (1858-1947) figured out that light also behaves as a *particle* using blackbody radiation.
- Albert Einstein (1879-1955) also figured out that light behaves as a particle based on the photo-electric effect.



Max Planck

- Light particle = photon =packet of EM energy
- Energy = hf (f is the frequency) h = Planck's constant = 6.626 × 10⁻³⁴ J·s
- Discovery of the photon helped initiate Quantum Mechanics.



Albert Einstein

Repeat Interference Experiment one photon at a time



Repeat Interference Experiment one photon at a time



[A. L. Weiss and T. L. Dimitrova, Swiss Physics Society, 2009.]

