

# Today's Topics

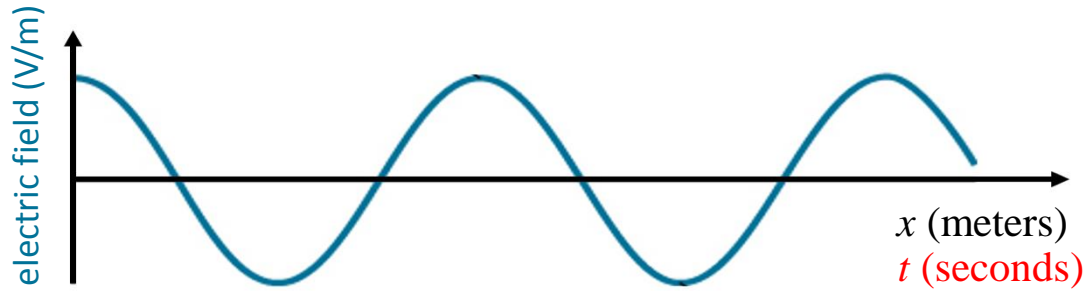
Monday, September 7, 2020 (Week 3, lecture 9) – Chapter 5.

A. Light basics: EM waves & photons

B. Electronic structure of atoms & spectroscopy

D. Doppler effect, part 1

# EM Wave Basics



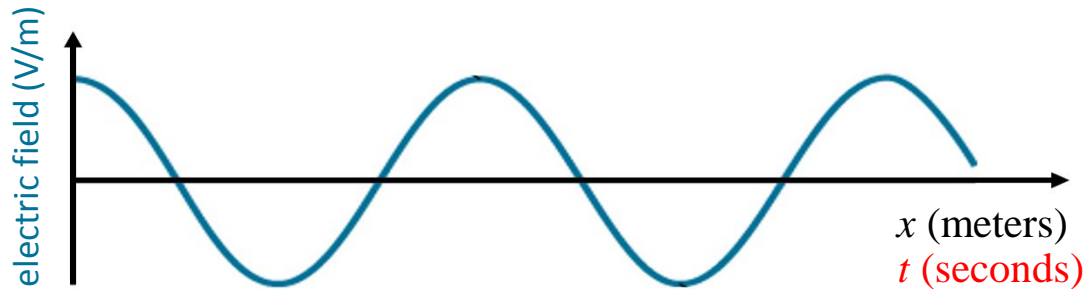
propagation  
at speed  $c$



magnetic field [Tesla]

$$= \frac{\text{electric field}}{c}$$

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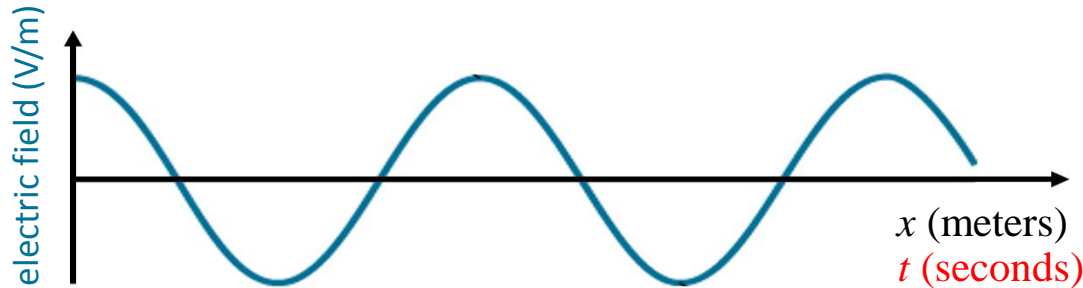
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*"brightness" of light (not source)*

$E$  = electric field [V/m]

$\epsilon_0$  = permittivity of vacuum  
 $= 8.85 \times 10^{-12} \text{ C}^2/(\text{N}\cdot\text{m}^2)$

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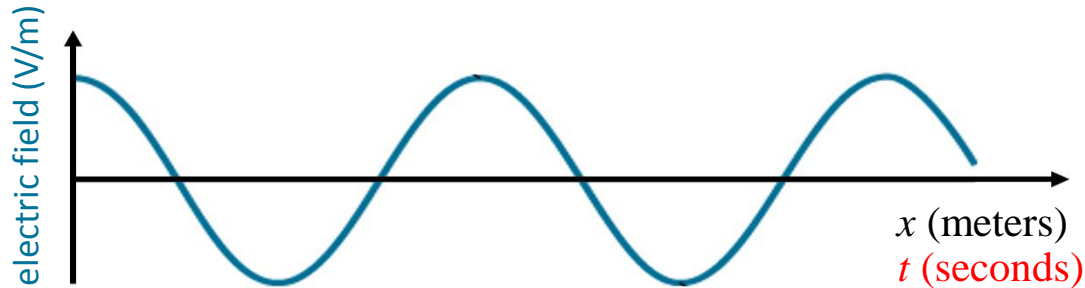
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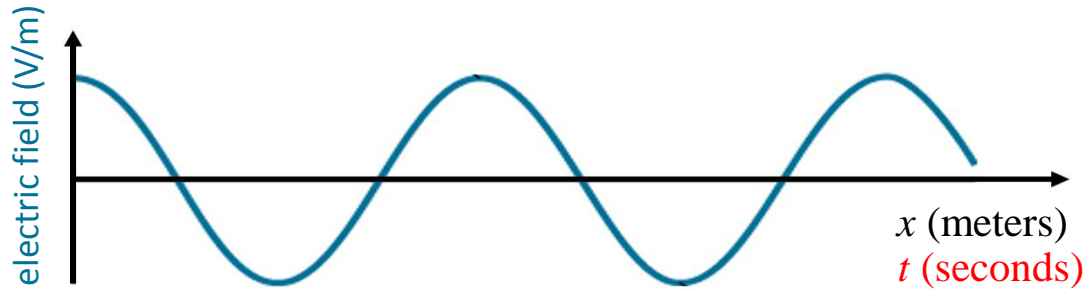
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**Question:** How do you generate light?

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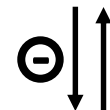
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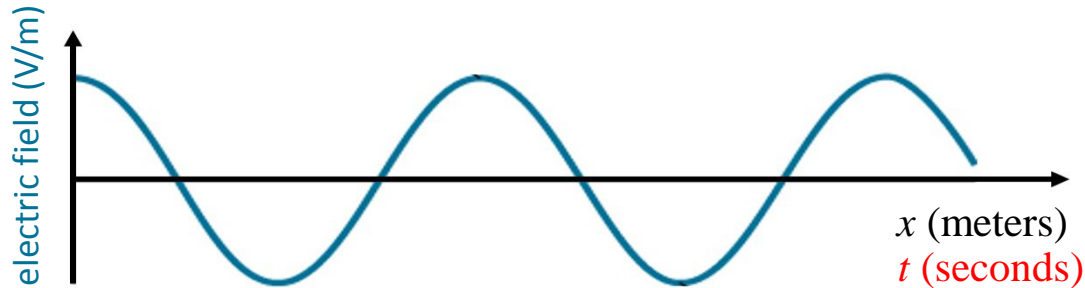
**Question:** How do you generate light?

**Answer:** oscillate charge (or accelerate it).

→ generates "dipole radiation."



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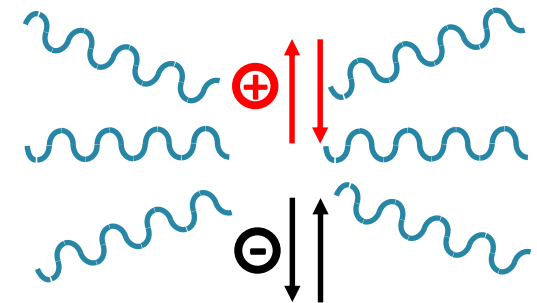
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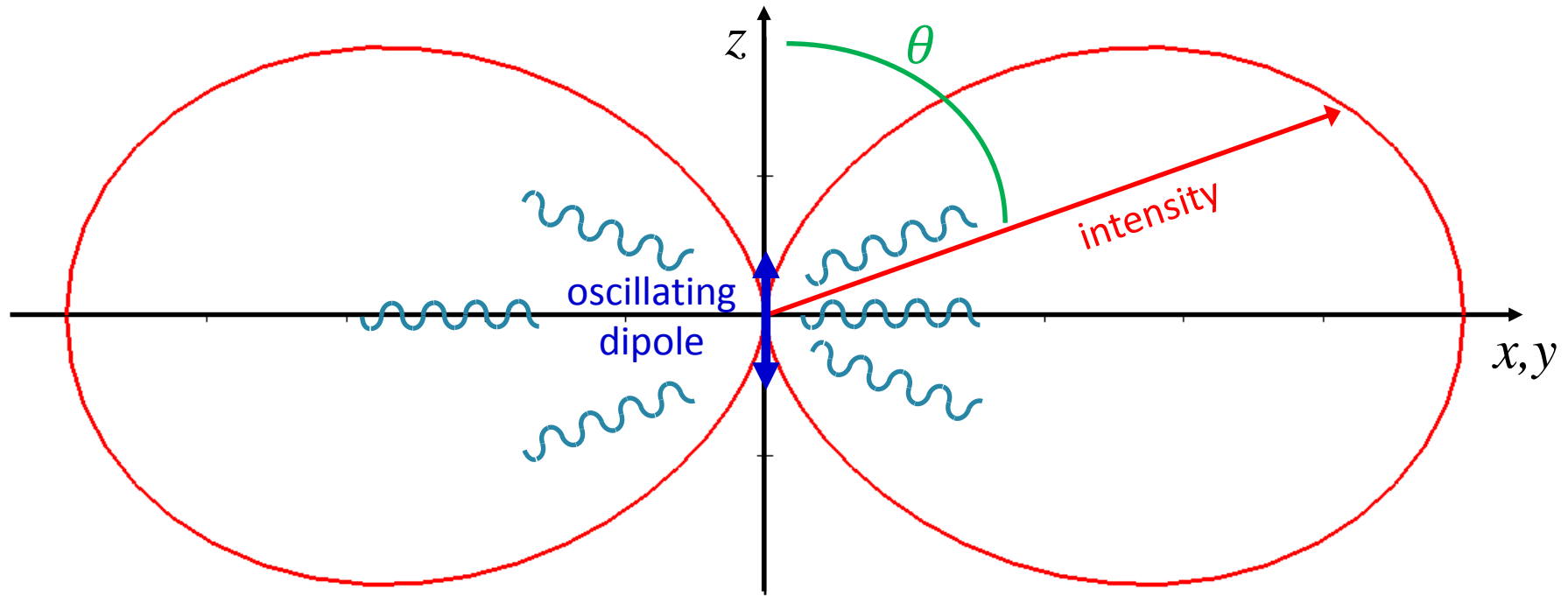
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# Dipole Radiation Pattern

dipole moment =  $p_0 = \text{charge} \times \text{separation}$



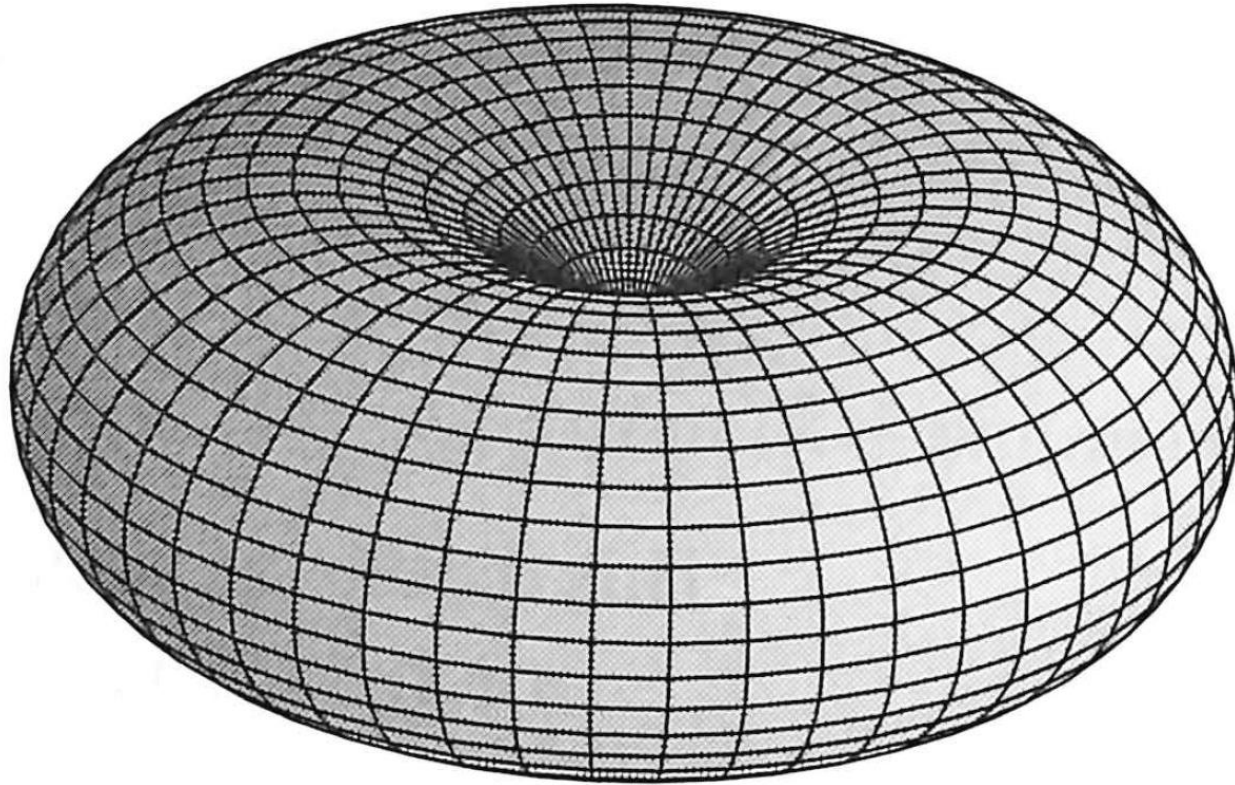
$$\text{Intensity} = \frac{\pi^2 p_0^2}{2\epsilon_0 c^3} \cdot f^4 \cdot \frac{\sin^2 \theta}{r^2} \quad \boxed{\propto f^4 \frac{1}{r^2}}$$

$r$  = distance from dipole

$f$  = frequency



# Dipole Radiation Pattern



[Figure 11.4, *Introduction to Electrodynamics*, by D. Griffiths, 4<sup>th</sup> Ed.]

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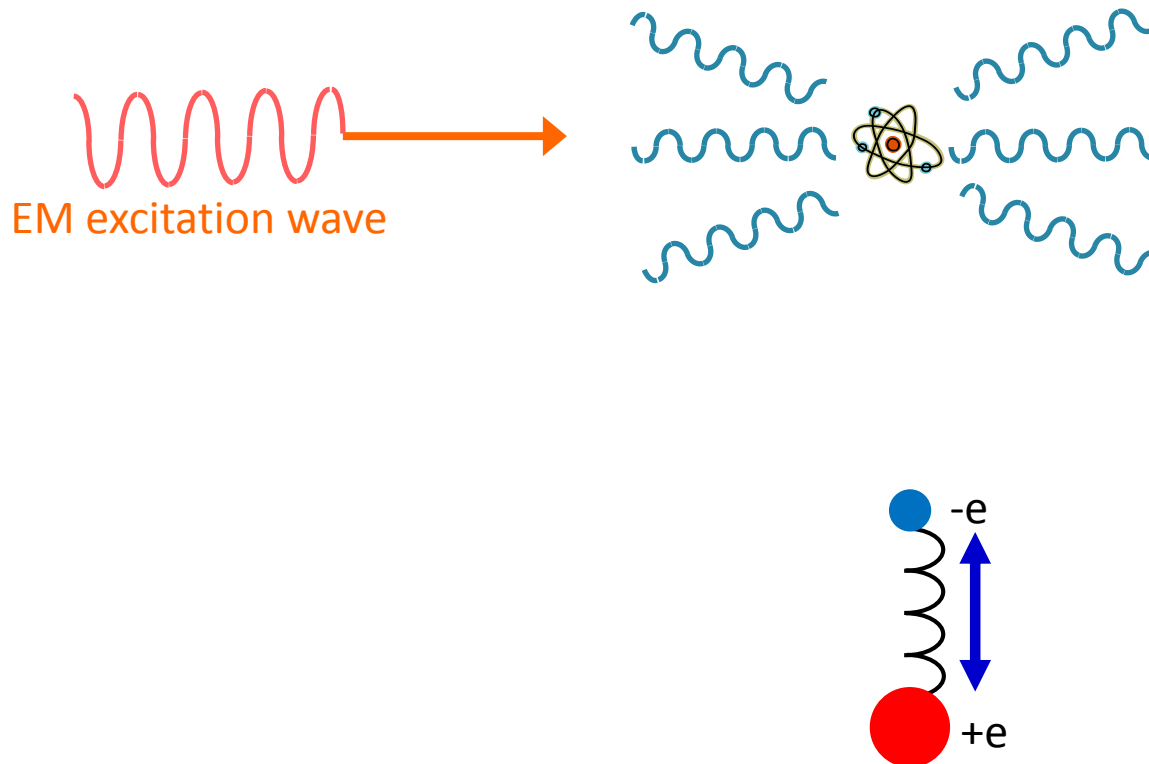
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# Dipole Radiation Example #1

## Atomic fluorescence & photon scattering

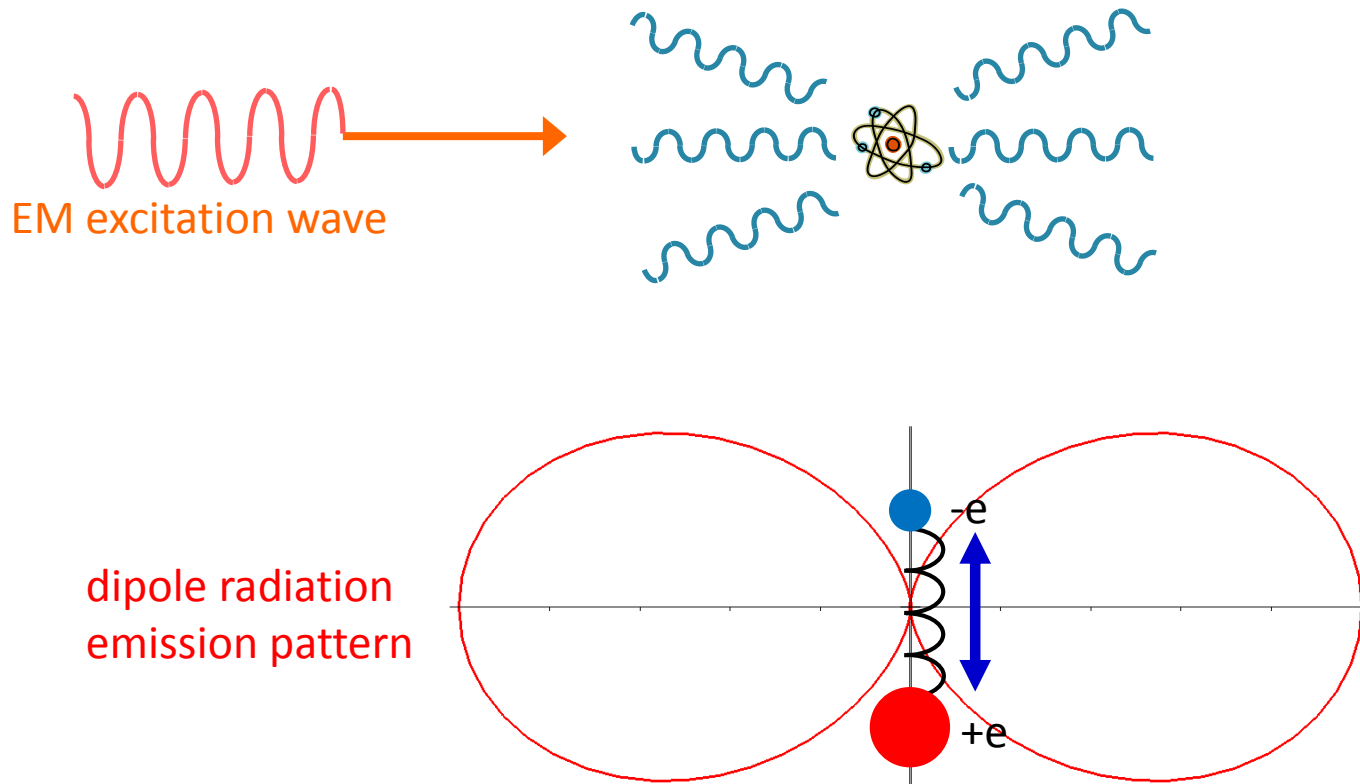
**Rayleigh scattering**: an atom behaves like a perfect electric dipole when excited by an EM wave.



# Dipole Radiation Example #1

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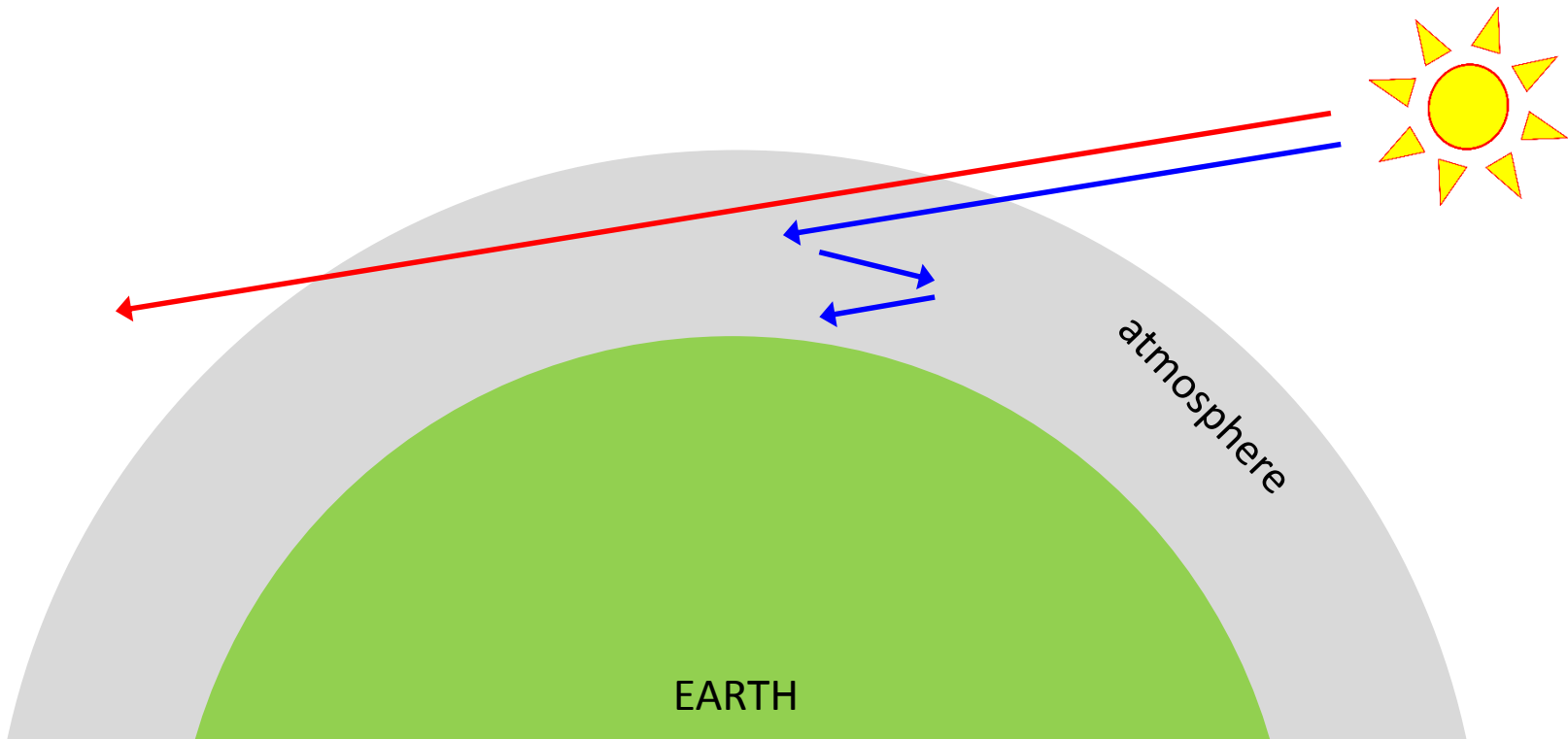
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# Dipole Radiation Example #2

## Blue Sky

**Blue light** scatters at a higher rate than **red light** → Sky looks blue.

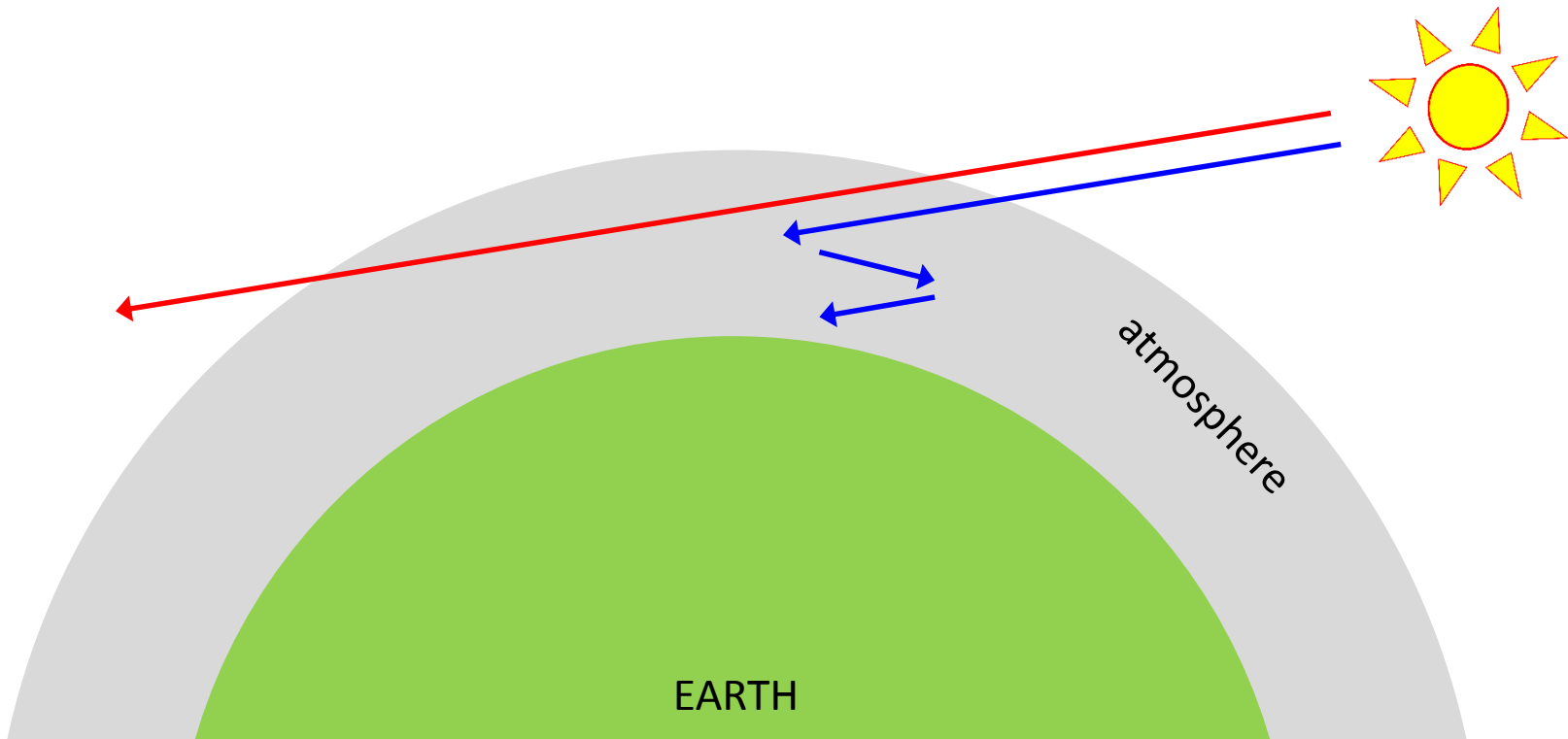


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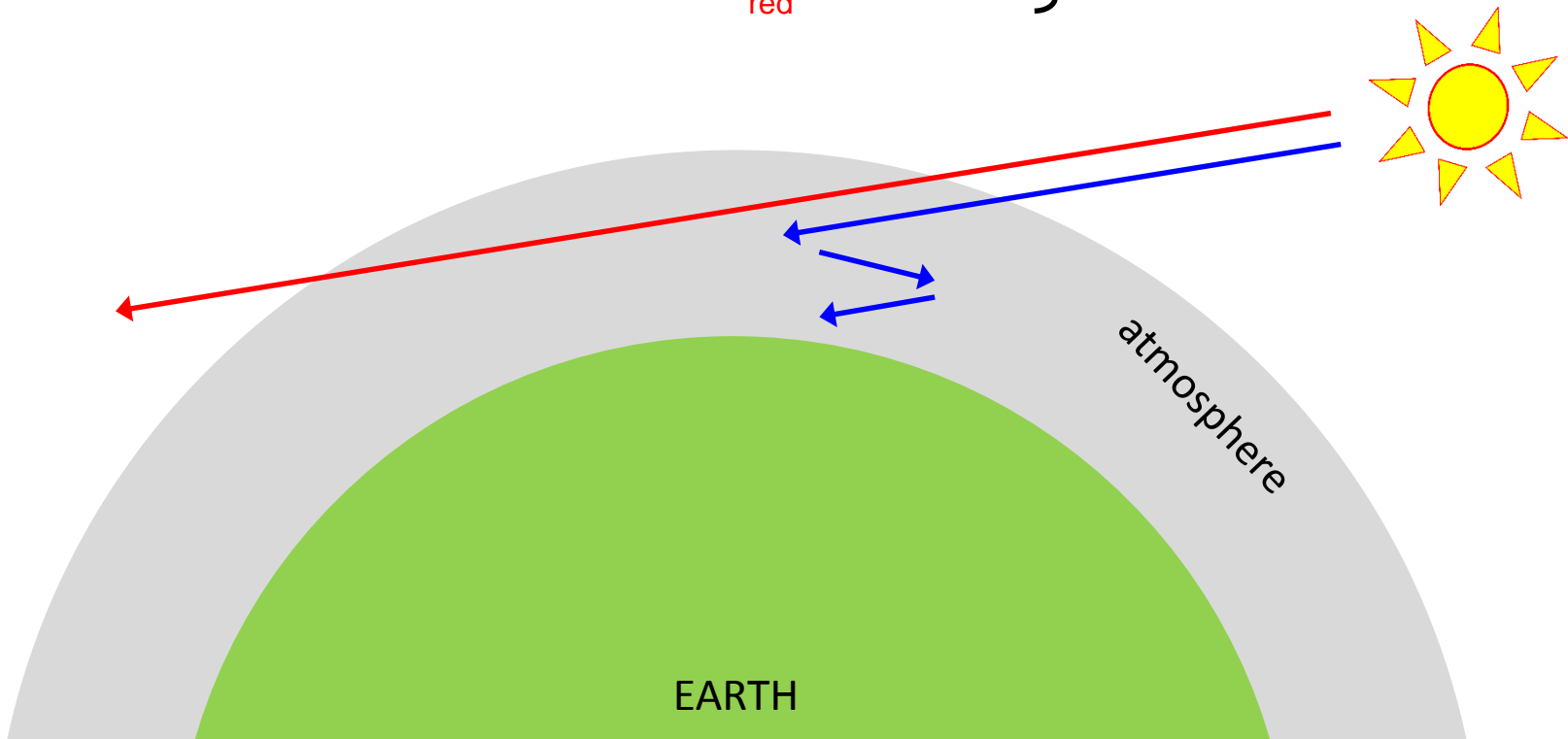


# Dipole Radiation Example #2

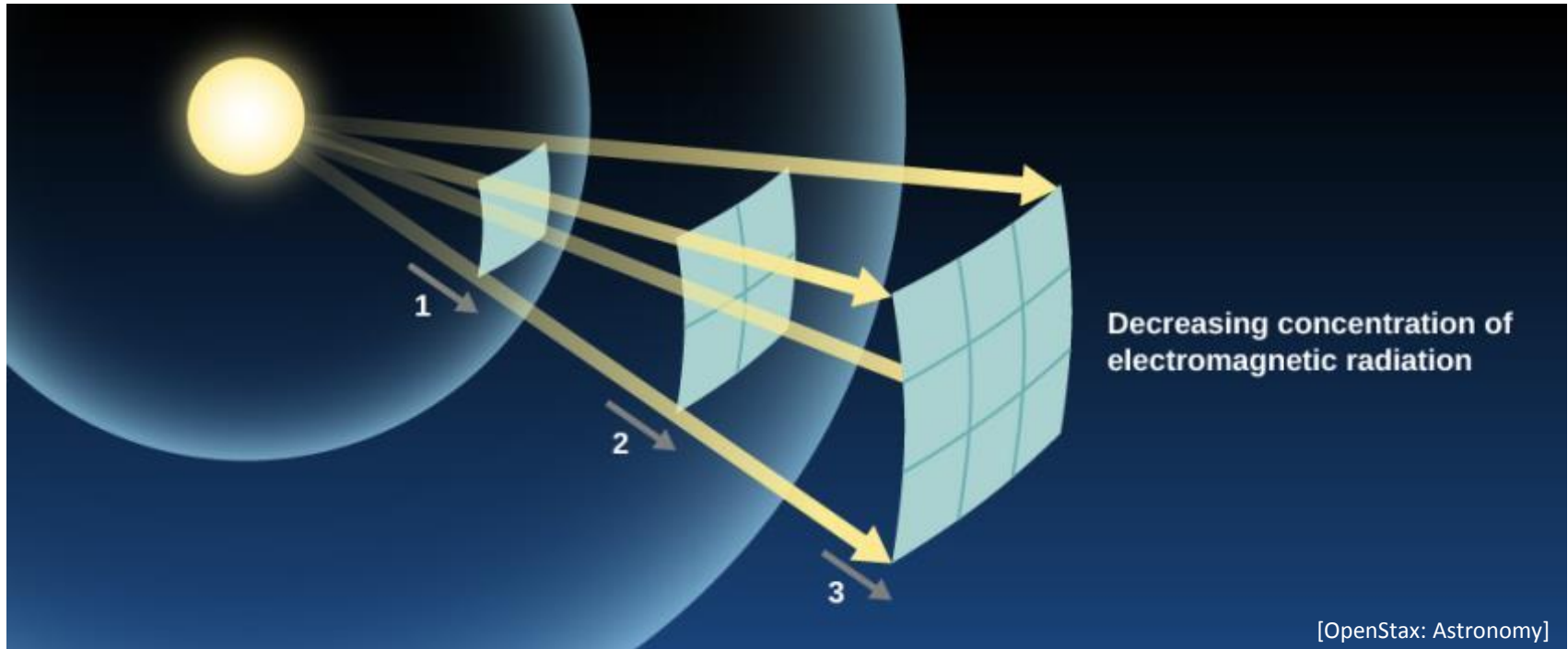
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$$Intensity \propto f^4 \propto \frac{1}{\lambda^4} \Rightarrow \left. \begin{array}{l} \lambda_{\text{blue}} = 450 \text{ nm} \\ \lambda_{\text{red}} = 650 \text{ nm} \end{array} \right\} \frac{I_{\text{blue}}}{I_{\text{red}}} = \left( \frac{650}{450} \right)^4 \approx 4.3$$



# Inverse Square Law for Light



- As light radiates away from its source, it spreads out such that its intensity decreases as the **square** of the **distance  $d$**  from its source.
- *Intensity*  $\propto 1/d^2$ .

# Photon Basics

$$\text{Photon energy} = E_{\gamma} = hf$$

Important: Photon is massless

$$M_{\gamma} = 0$$



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Note:  $p_{\gamma} \neq M_{\gamma}c$  ( $= 0$ )

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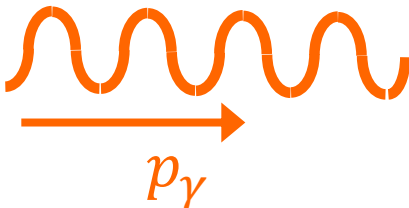
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Shine light on an atom



$$p_{atom} = 0$$

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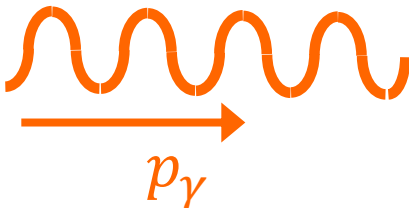
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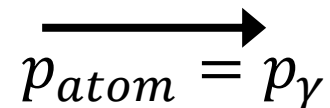
Shine light on an atom



$$p_{atom} = 0$$

Atom after absorption of 1 photon

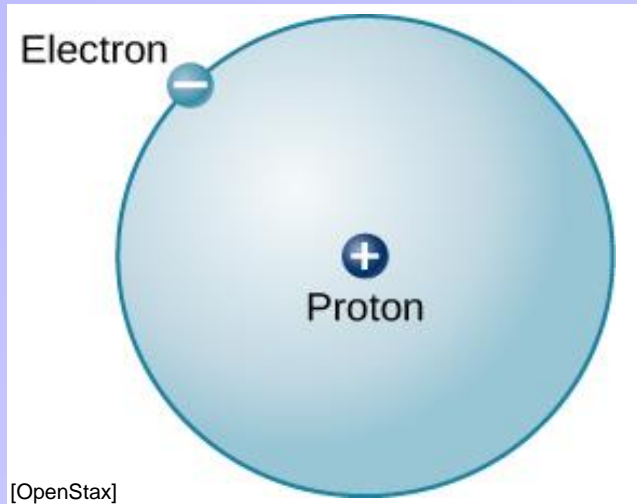


A diagram showing the atom from the previous step with a black arrow pointing to the right below it, indicating that the atom now has momentum. The arrow is labeled with the equation  $p_{atom} = p_\gamma$ .

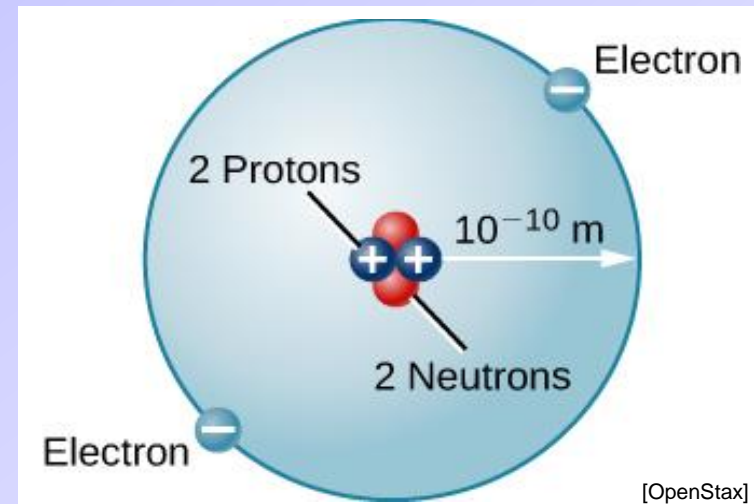
(conservation of momentum)

# Basic Structure of Atoms

Hydrogen:  ${}^1\text{H}$



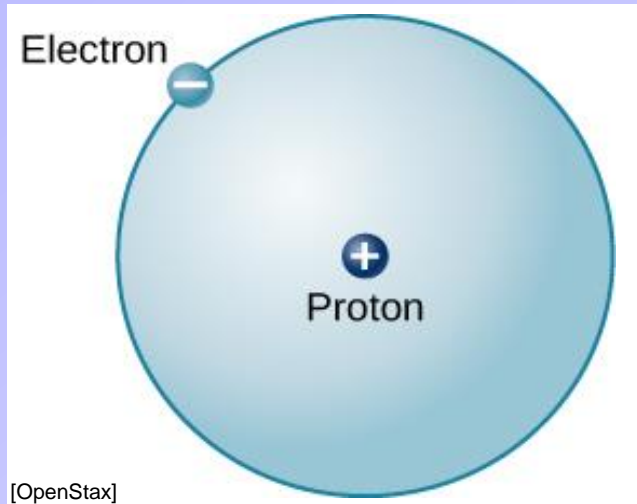
Helium:  ${}^4\text{He}$   $4 = 2 \text{ protons} + 2 \text{ neutrons}$



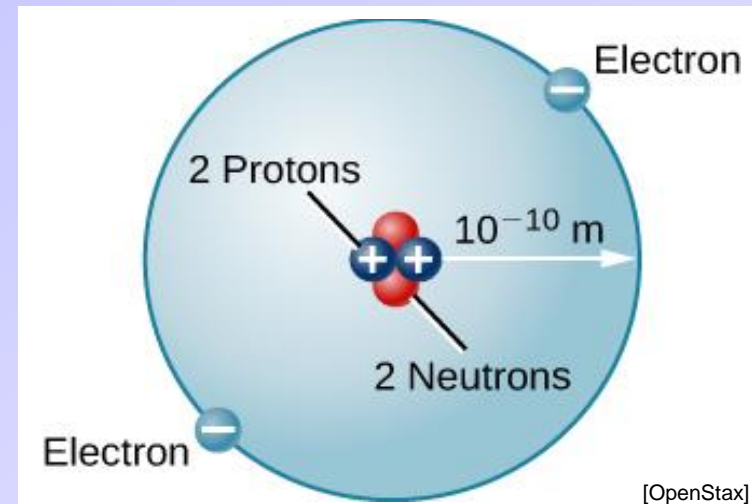
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- Electron number, orbits, and properties determine the **chemistry** of the atom.

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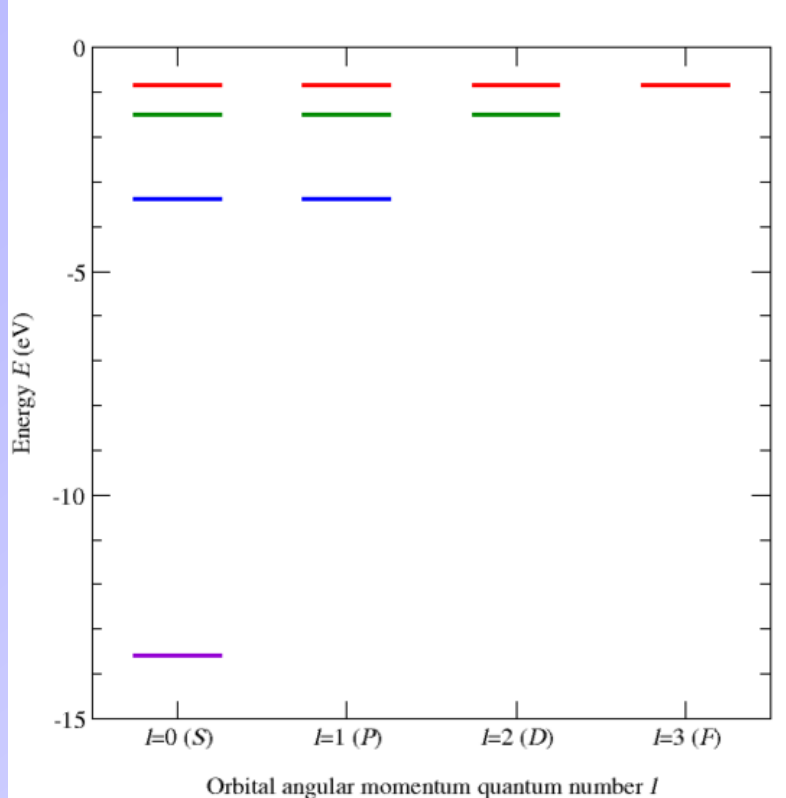
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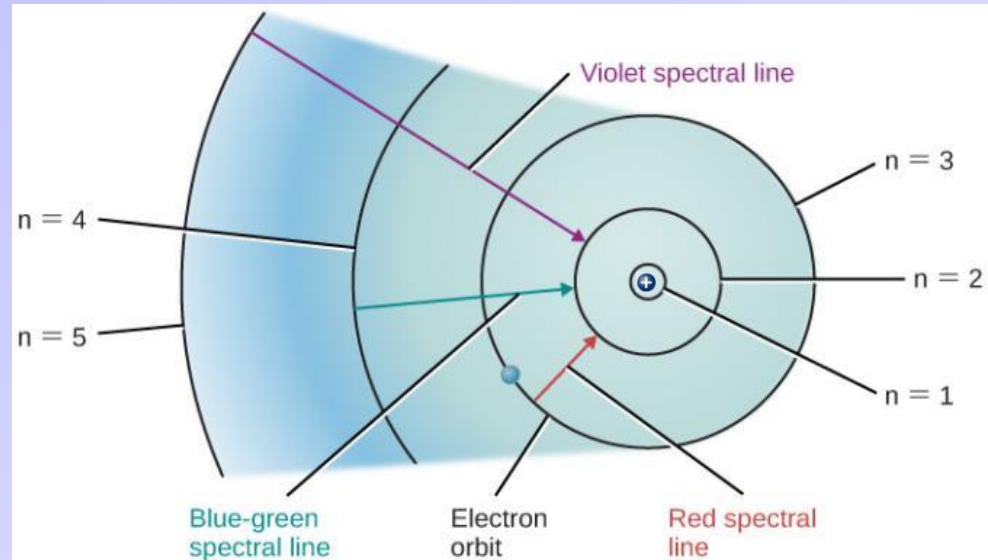
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- Electron number, orbits, and properties determine the **chemistry** of the atom.
- **Nucleus** consists of positively charged **protons** and **neutral neutrons**.
- For neutral atoms: **Number of protons = number of electrons**.
- Neutrons help bind protons together. **Number of neutrons  $\geq$  number of protons**.

# Electronic Structure of Atoms

Energy Levels of Hydrogen ( $n=1-4$ )



[Figure from wikimedia.org]



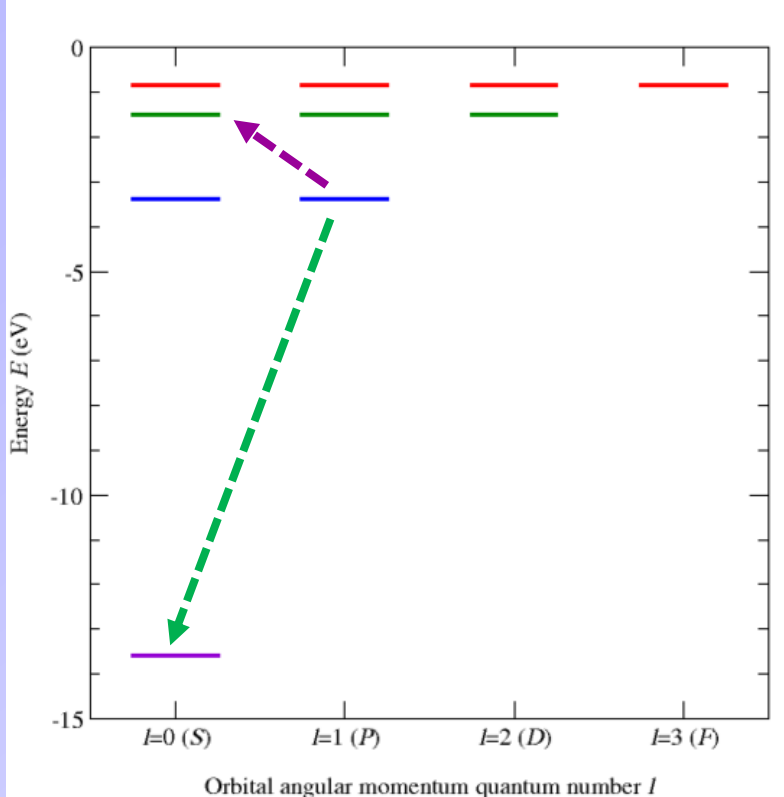
[OpenStax]

- Electrons have **discrete allowed energies and orbits**.

Note:  $1 \text{ eV} = 1.602 \times 10^{-19} \text{ J}$

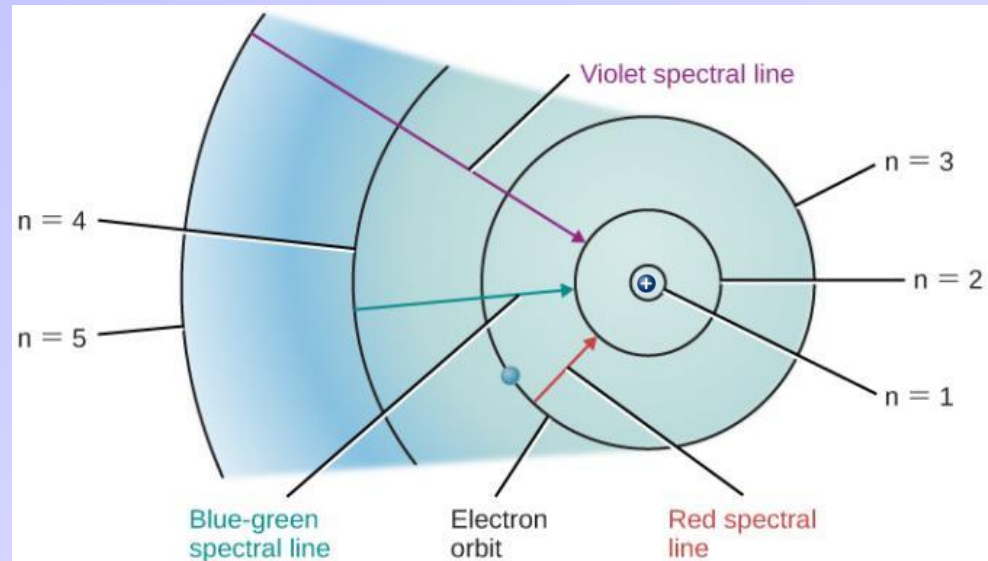
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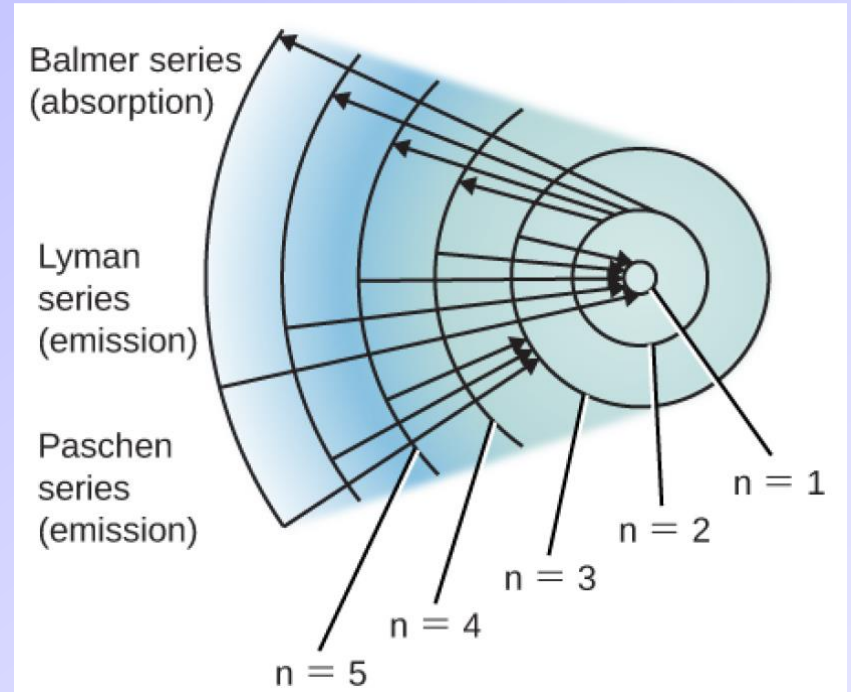
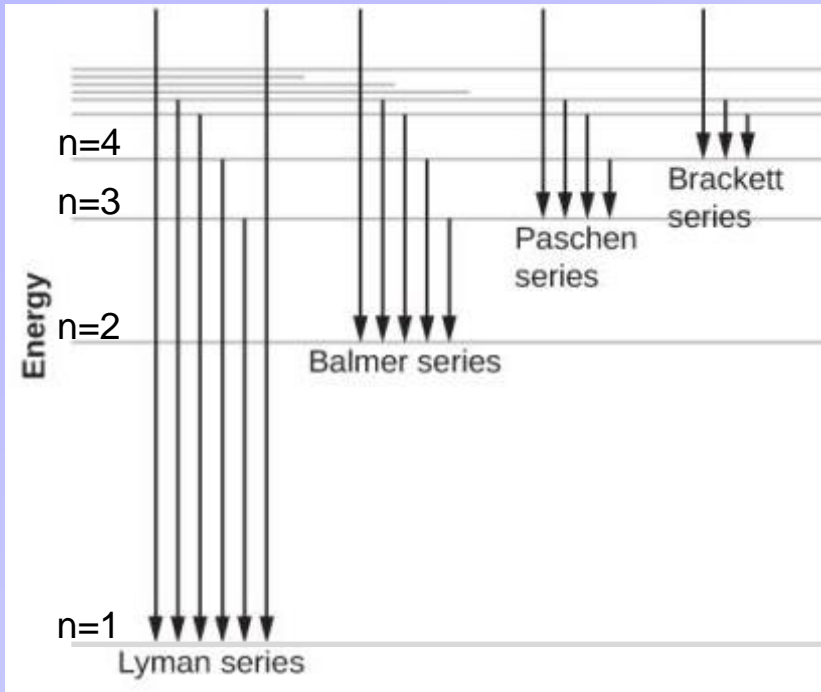
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[OpenStax]

- Electrons have **discrete allowed energies and orbits**.
- Transitions between two energy levels requires **emission** or **absorption** of a photon that bridges the energy gap.
- **Discrete** emission and absorption **spectra**.

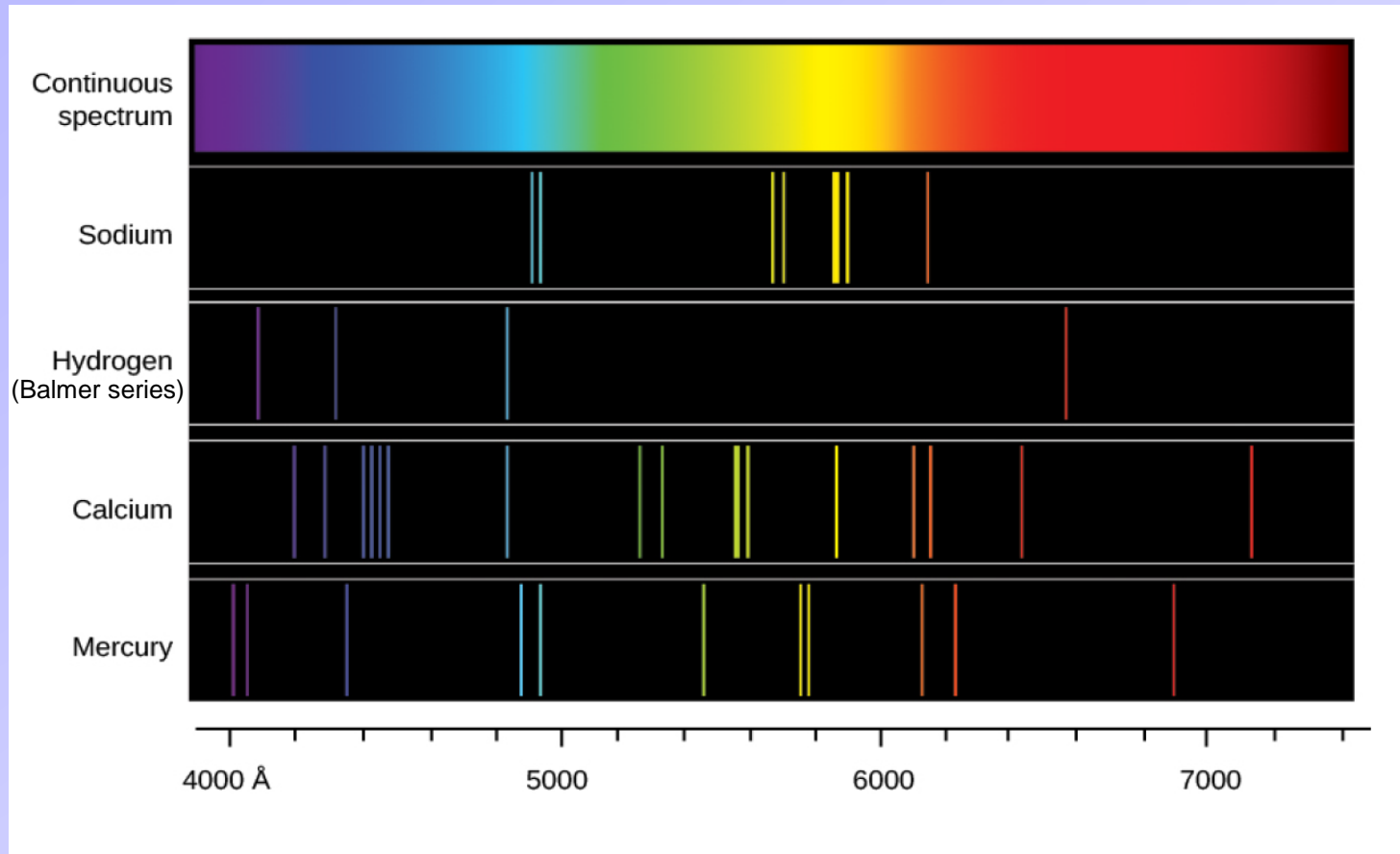
# Emission Spectrum of Hydrogen



- Hydrogen has a number of emission and absorption spectral series that depend on the start/end point of the transition.
- Other elements are qualitatively similar.
- Also true for molecules, but their spectra are more complicated.

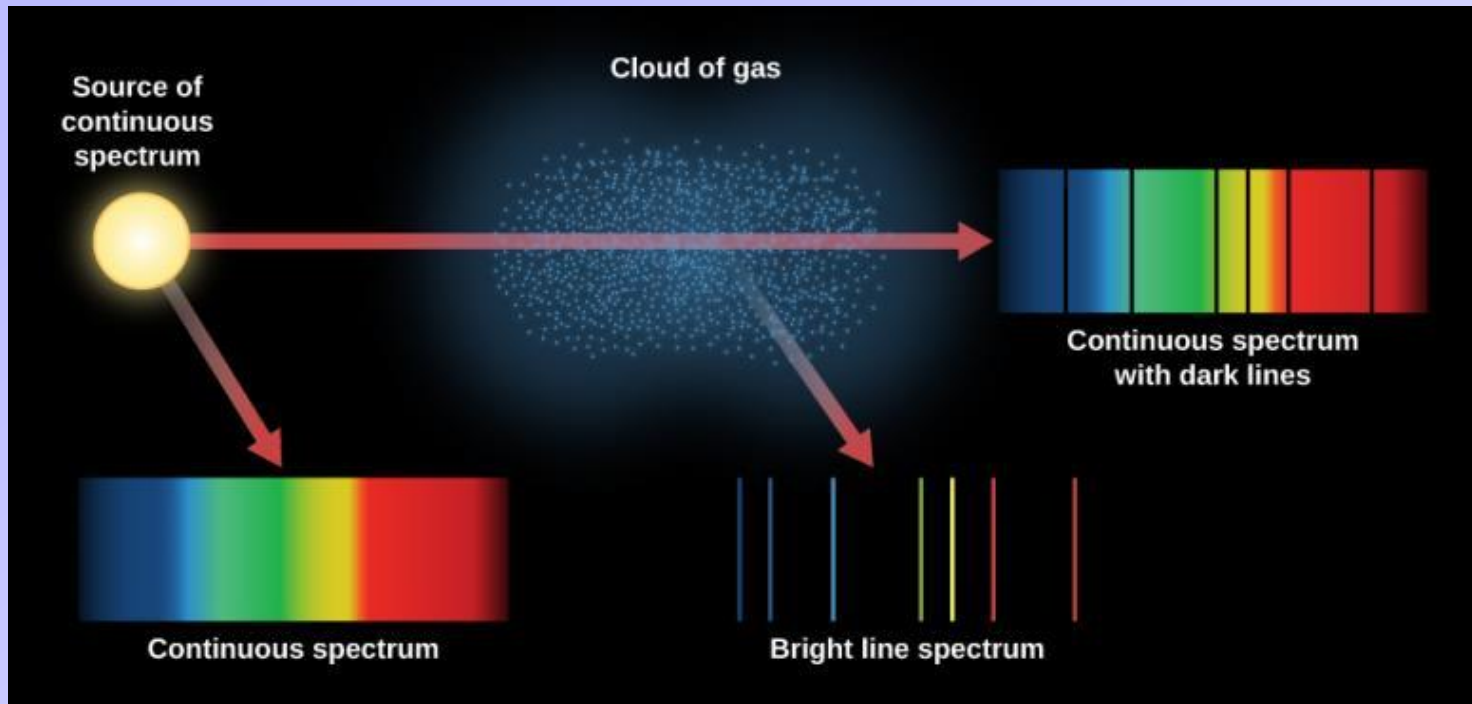


# Emission Spectra “Fingerprints”



If you build a **catalog of spectral lines**, then you can **determine the elements** that are present from the spectrum.

# Emission & Absorption Processes



## Three types of spectra

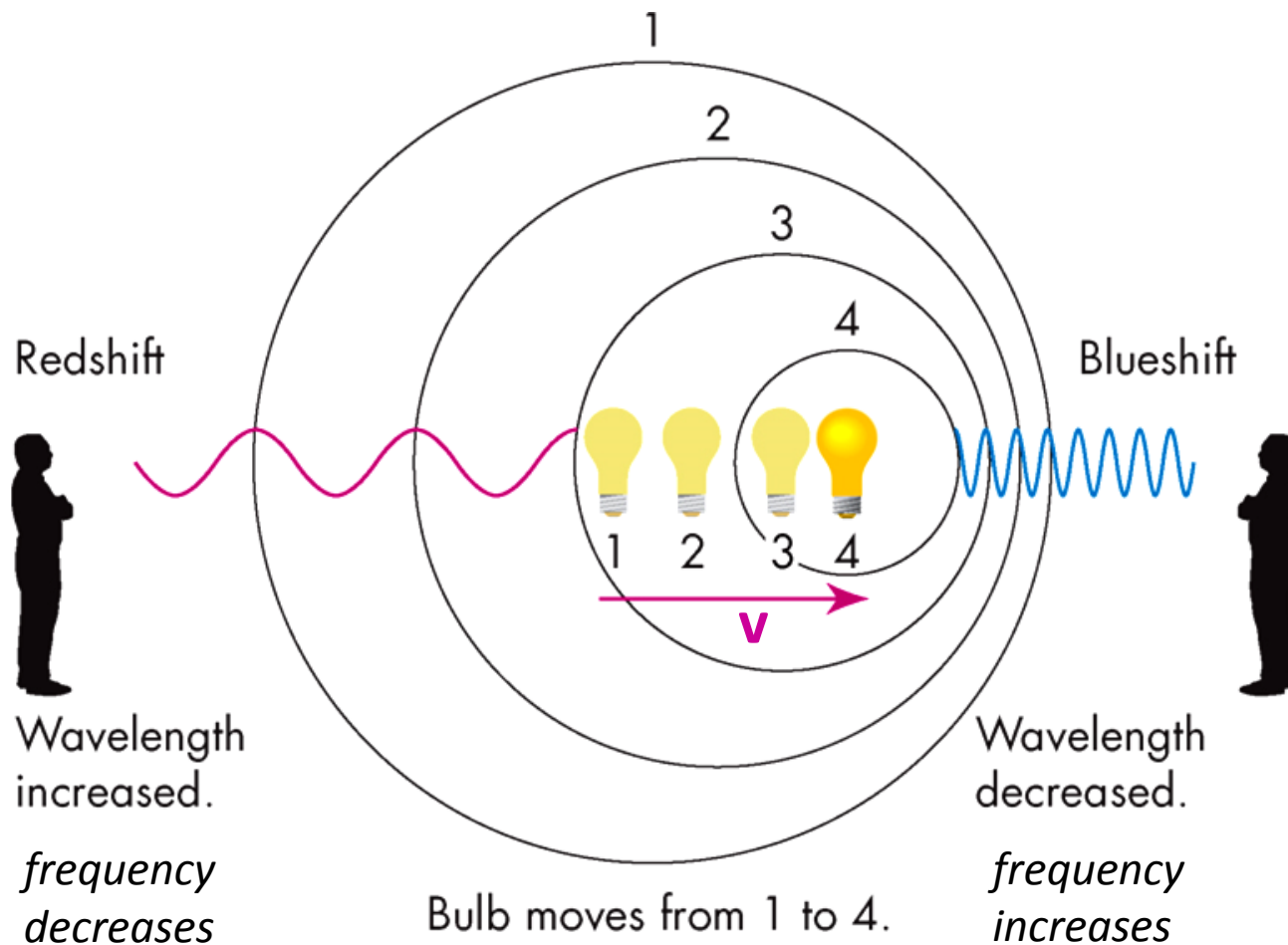
1. **Continuous** spectrum, e.g. a thermal **blackbody** source.
2. **Emission** spectrum (discrete): if light excites atoms, then the atomic emission will be at discrete frequencies.
3. **Absorption** spectrum (discrete): if a continuous spectrum excites atoms, then the absorption of photons will remove light at discrete frequencies (“shadow lines”).

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A **moving source** cannot change the speed of its emitted light, but it does change its **frequency & wavelength**.

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Works for sound too !!!

# Doppler Shift Calculation

Doppler frequency shift:  $\frac{\Delta f}{f} = -\frac{\Delta \lambda}{\lambda} = \frac{v}{c}$  with  $f' = f + \Delta f$

frequency of stationary source 

perceived frequency of moving source 

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frequency of stationary source  $\nearrow$   $f$

$\nearrow$  perceived frequency of moving source  $f'$

If source is moving towards you, then light is blue shifted.

$$v > 0$$

$$\Delta f > 0, f' \text{ goes up}$$

$$\Delta \lambda < 0$$

If source is moving away from you, then light is red shifted.

$$v < 0$$

$$\Delta f < 0, f' \text{ goes down}$$

$$\Delta \lambda > 0$$