

Today's Topics

Wednesday, October 14, 2020 (Week 8, lecture 23) – Chapters 11, 12.

A. Saturn

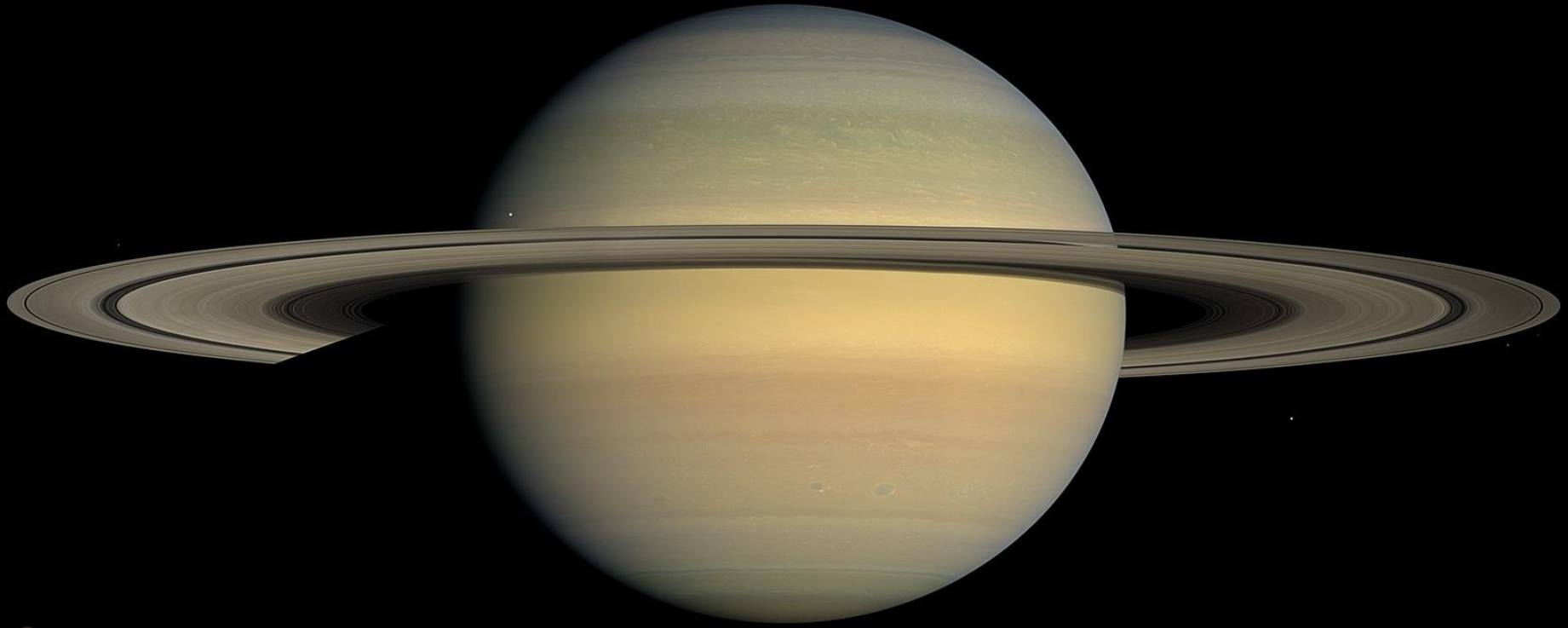
B. The Roche Limit

C. Titan

D. Uranus

E. Neptune

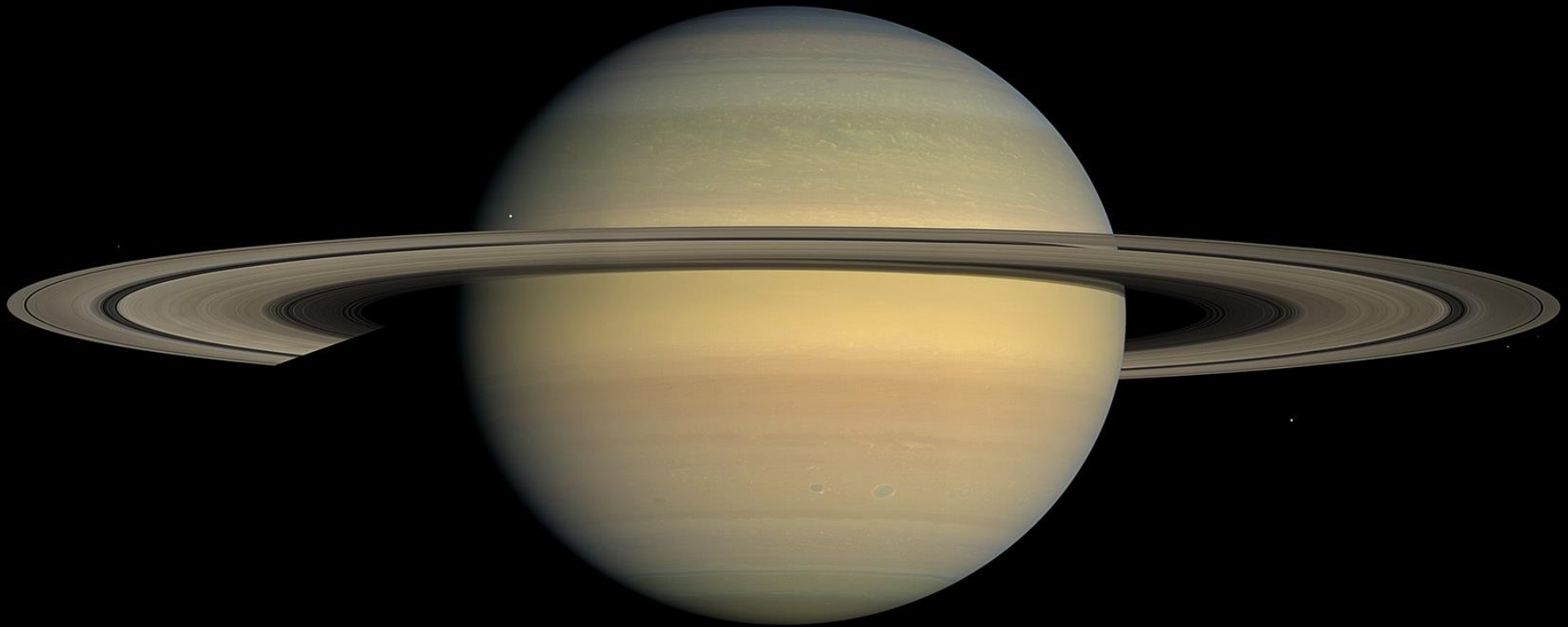
Saturn: “Lord of the Rings”



[NASA: Cassini mission, 2008]

- “Lightweight Jupiter” with rings.
 - $M_{\text{Saturn}} \approx 0.3 M_{\text{Jupiter}}$
 - $R_{\text{Saturn}} = 58,500 \text{ km} \approx 0.84 R_{\text{Jupiter}}$
- Almost twice as far as Jupiter.
 - Semimajor axis: 9.5 AU (J: 5.2 AU).
 - Orbital period: $T = 29.5 \text{ yrs}$ (J: 11.9 yrs).

Saturn: “Lord of the Rings”



[NASA: Cassini mission, 2008]

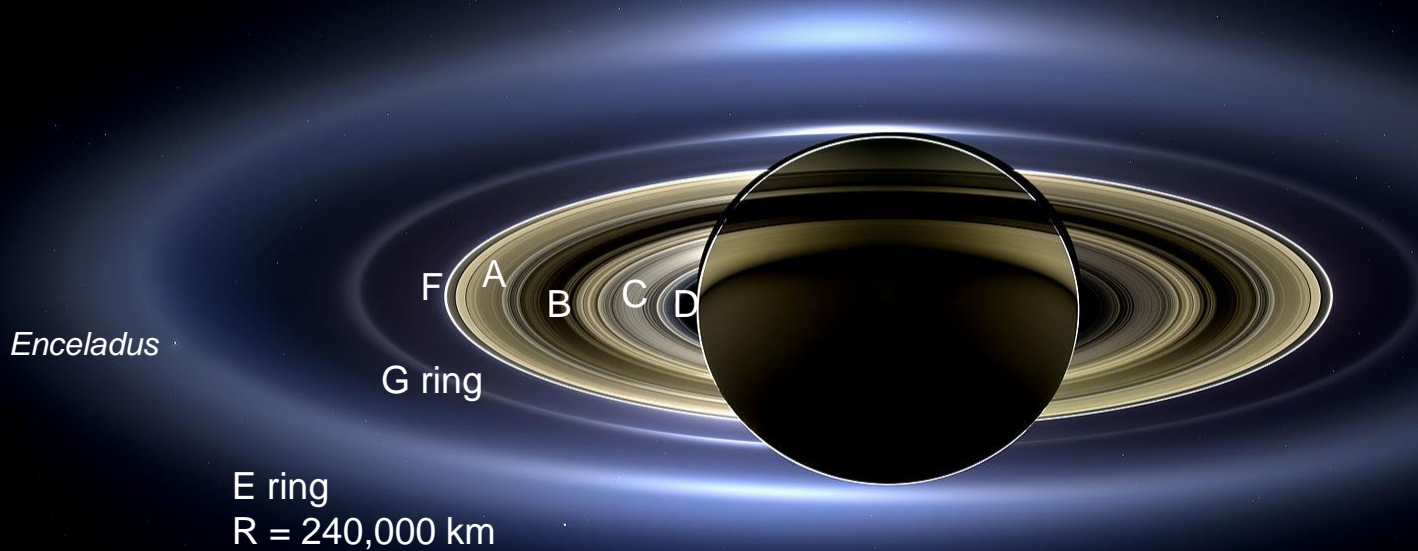
- “Lightweight Jupiter” with rings.
 - $M_{\text{Saturn}} \approx 0.3 M_{\text{Jupiter}}$
 - $R_{\text{Saturn}} = 58,500 \text{ km} \approx 0.84 R_{\text{Jupiter}}$
- Almost twice as far as Jupiter.
 - Semimajor axis: 9.5 AU (J: 5.2 AU).
 - Orbital period: $T = 29.5 \text{ yrs}$ (J: 11.9 yrs).
- Rotates quickly: $T_{\text{rotation}} = 10.6 \text{ hrs.}$
- At least 82 moons.
 - Only moon in Solar System with an atmosphere: **Titan**.
- Giant rings.
- Large magnetosphere.

The Rings I

7 distinct ring families

- A, B, C are the most visible.
- B is the brightest (not in photo).
- D, F, G, and E are very faint.

Outer radius of A = 137,000 km
Inner radius of C = 74,500 km



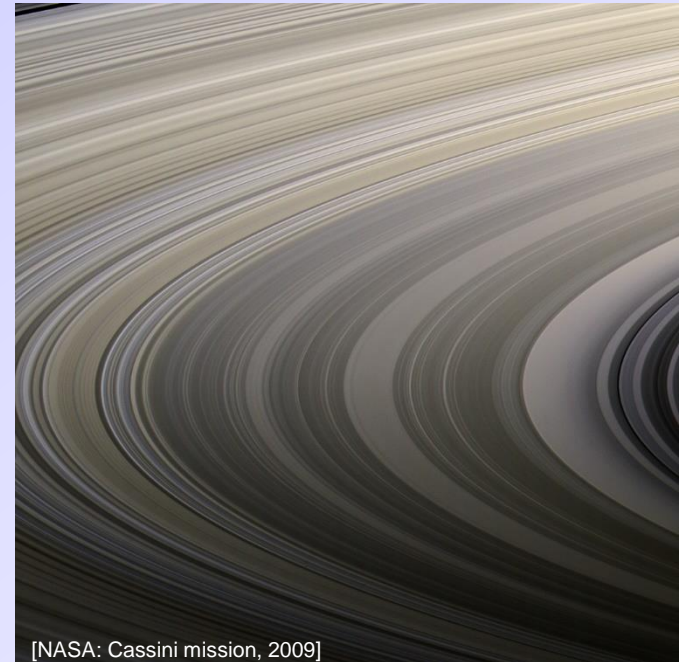
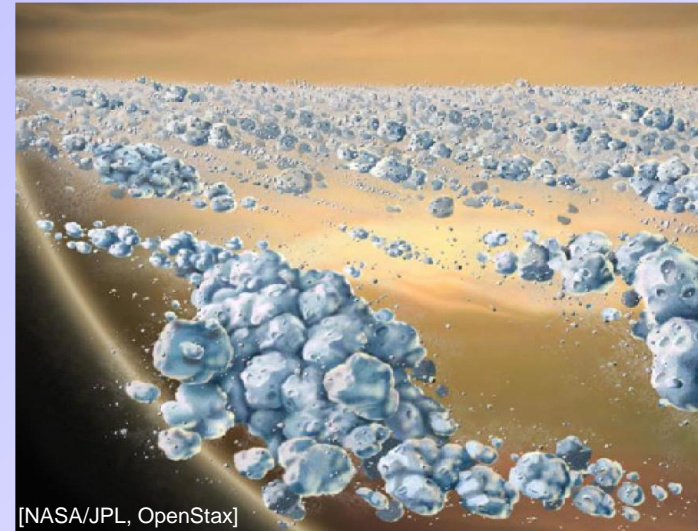
[NASA: Cassini mission, 2013]

Enhanced image of Saturn's rings (Sun is eclipsed by Saturn)

The Rings II

Ring Properties

- Mostly **water ice** (99.9%).
- Some silicates
- “Particle” size range: 1 cm to 10 m.
- Ring particles **clump together**.
(mostly multi-meter scale, but up to 10 km in size)
- Thickness of rings: 10-100 m.
- Rings slowly falling into Saturn as “**ring rain.**”
→ *Rings may be gone in 100s of millions of years.*



Saturn's B ring

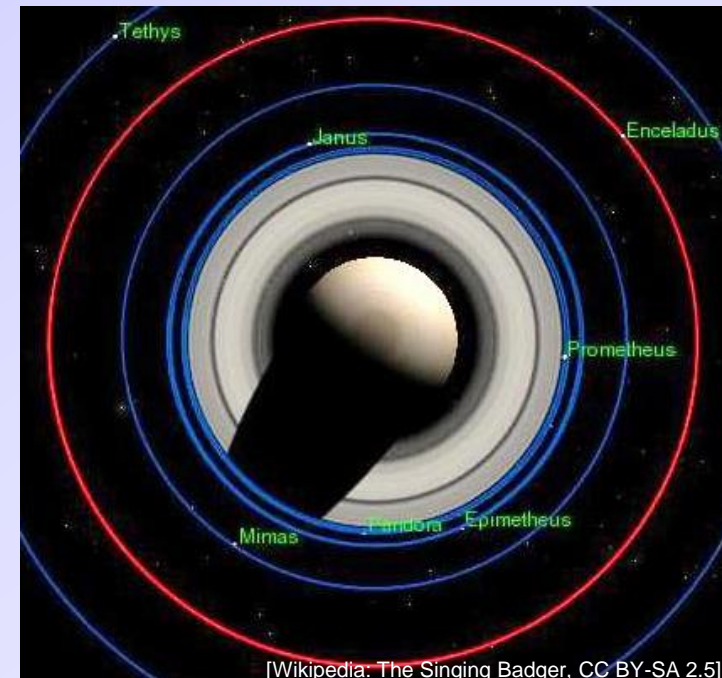
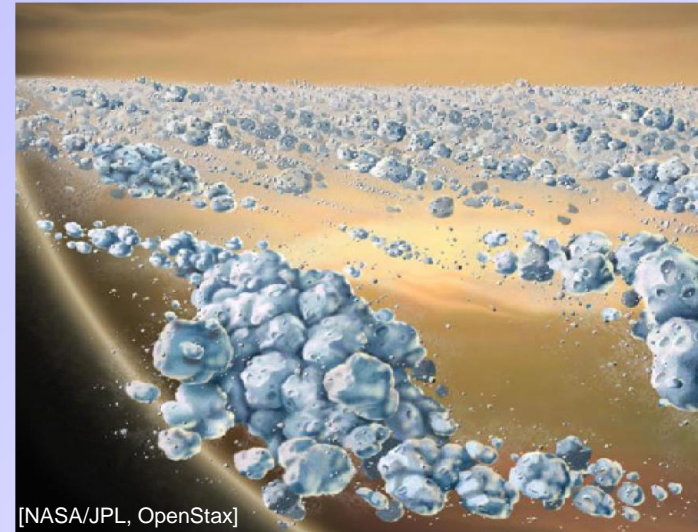
The Rings II

Ring Properties

- Mostly **water ice** (99.9%).
- Some silicates
- “Particle” size range: 1 cm to 10 m.
- Ring particles **clump together**.
(mostly multi-meter scale, but up to 10 km in size)
- Thickness of rings: 10-100 m.
- Rings slowly falling into Saturn as “**ring rain.**”
→ Rings may be gone in 100s of millions of years.

Shepherd moons

- Moons help to keep the rings from spreading out.
→ *Prometheus*, *Daphnis*, *Pan*, *Janus*, and *Epimetheus*.
- Moonlets within the rings also act as **shepherds**.
- **Cryogeisers on Enceladus** feed the E ring.



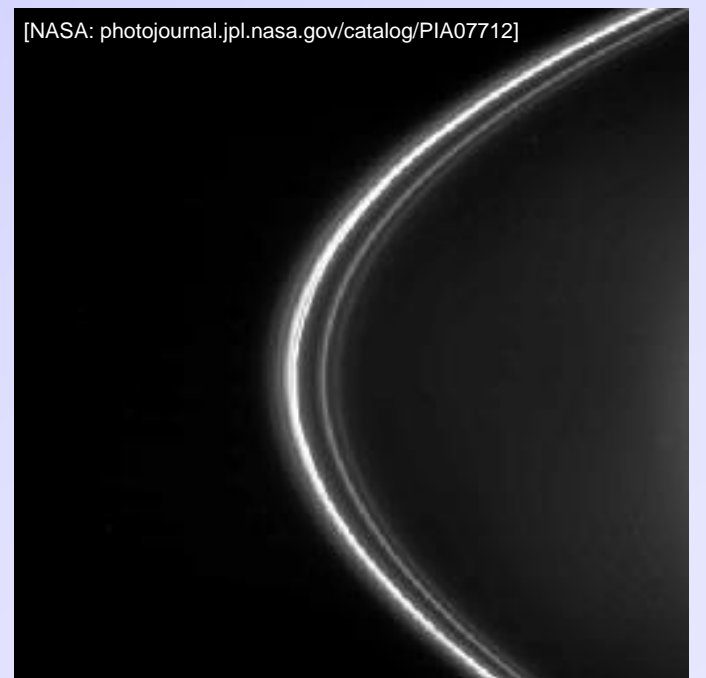
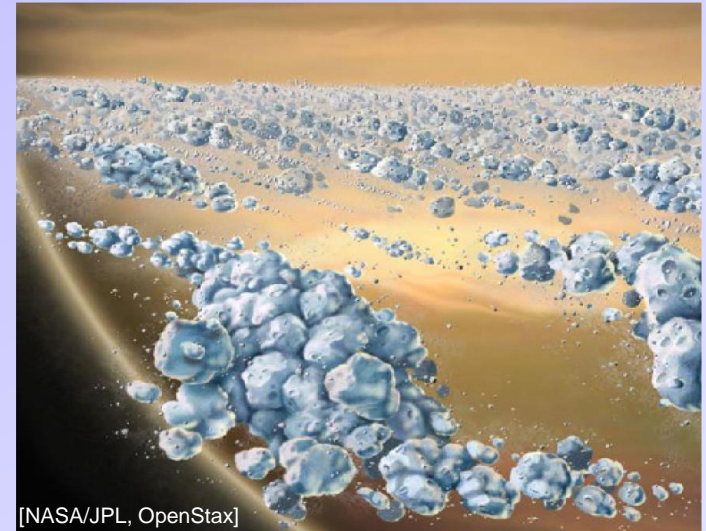
The Rings II

Ring Properties

- Mostly **water ice** (99.9%).
- Some silicates
- “Particle” size range: 1 cm to 10 m.
- Ring particles **clump together**.
(mostly multi-meter scale, but up to 10 km in size)
- Thickness of rings: 10-100 m.
- Rings slowly falling into Saturn as “**ring rain.**”
→ *Rings may be gone in 100s of millions of years.*

Shepherd moons

- Moons help to keep the rings from spreading out.
→ *Prometheus, Daphnis, Pan, Janus, and Epimetheus.*
- Moonlets within the rings also act as **shepherds**.
- **Cryogeisers on Enceladus** feed the E ring.

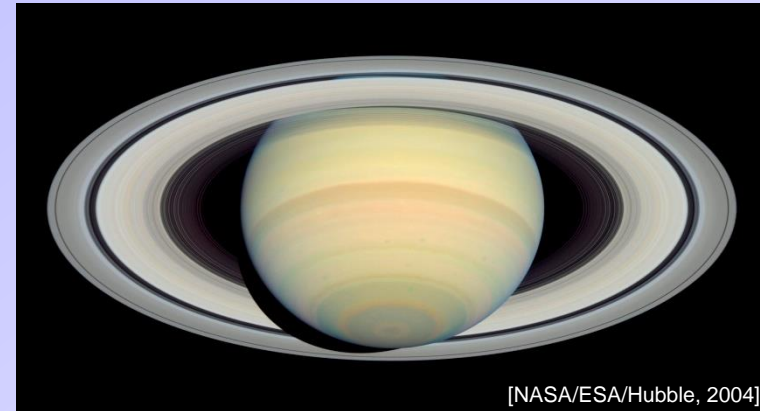


Pandora

Prometheus

Origin of the Rings

The origin and age of the rings is not well understood (i.e. open question).



Hypothesis A: *Old rings*

The rings are the remnant of the **formation** of the mini “solar system” of Saturn.

→ Rings are very old: about as old as the Solar System, i.e. ~ 4.4 billion years old.

Hypothesis B: *Young rings*

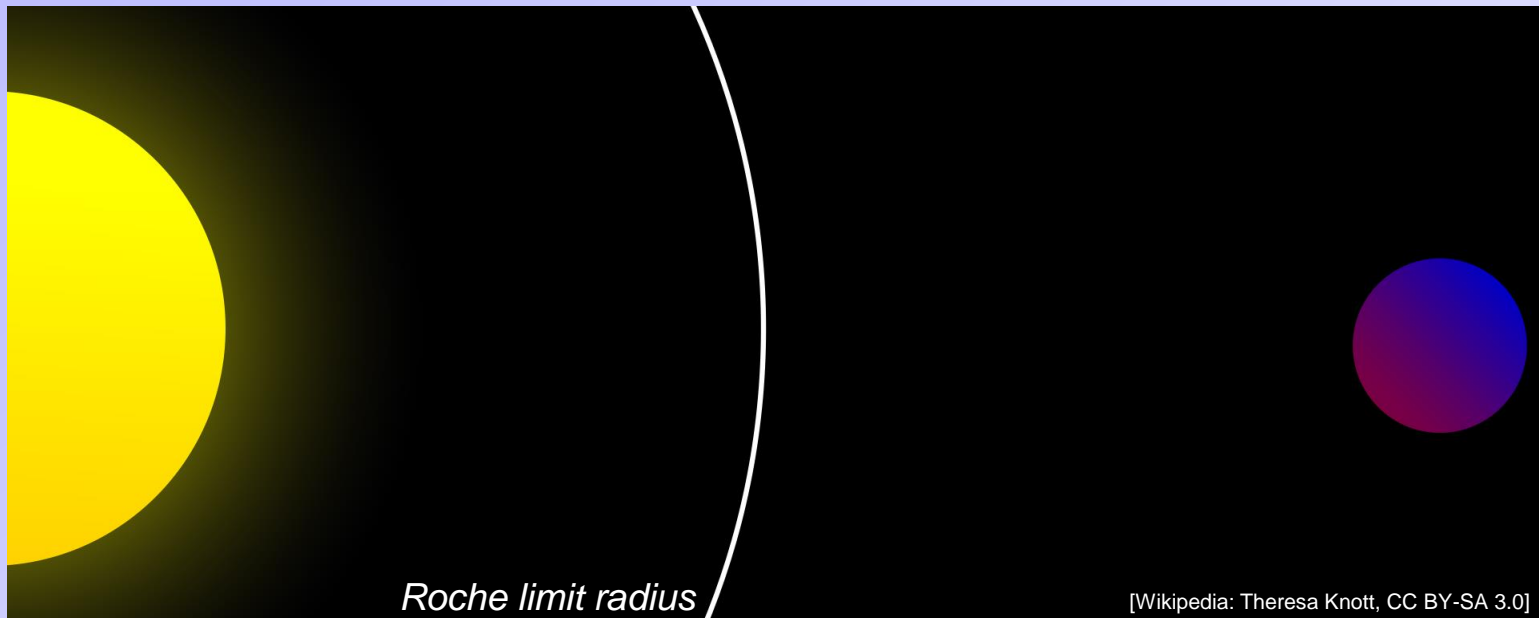
The rings were formed when two icy moons **collided** or when a moon got too close to Saturn (**Roche limit**) and was pulled apart by **tidal forces**.

→ Rings are relatively young: perhaps 100s of millions of years old.

The Roche Limit

The Roche limit is the orbital radius at which a gravitational bound object will be **pulled apart** by the **tidal force** from the central mass (i.e. Sun, Saturn, etc).
(gravity gradient)

- The Roche limit depends on nature of body (solid, fluid, density).
- Proposed by Eduard Roche in 1848 (French astronomer).

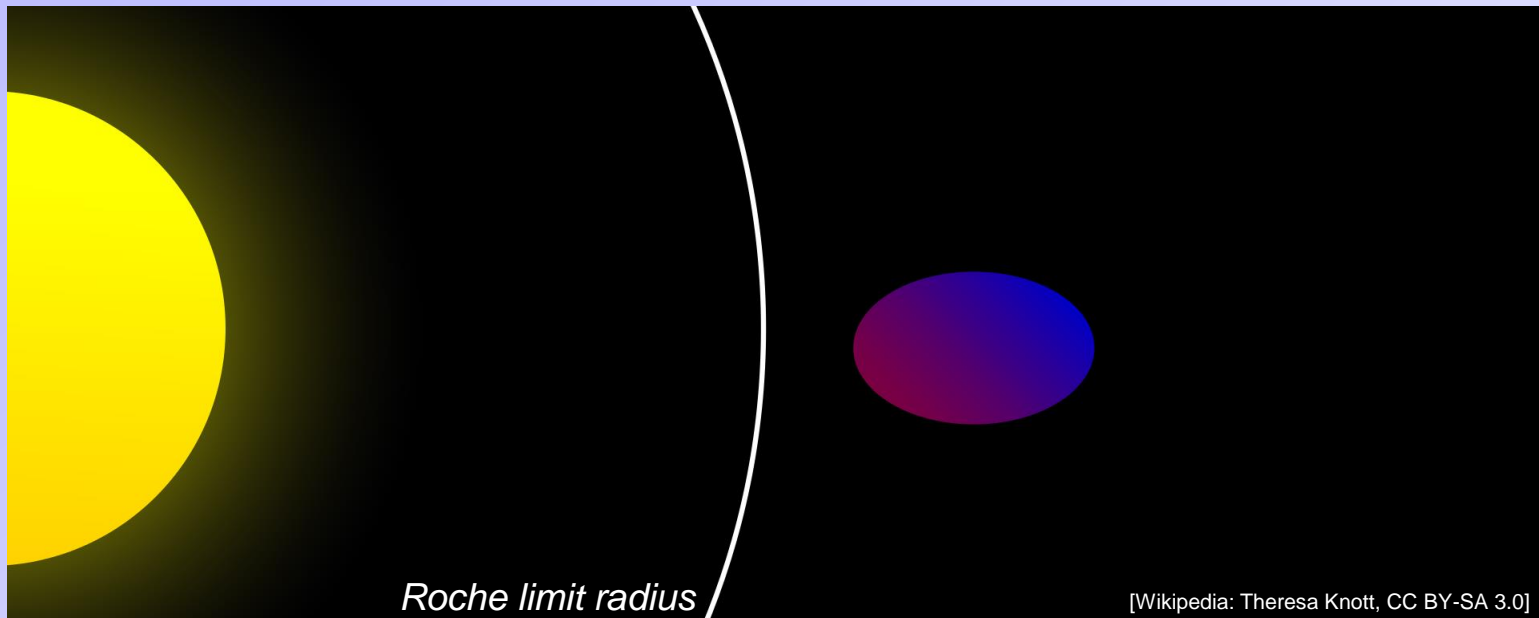


Far outside the Roche limit radius, the tidal force and deformation are **weak**.

The Roche Limit

The Roche limit is the orbital radius at which a gravitational bound object will be **pulled apart** by the **tidal force** from the central mass (i.e. Sun, Saturn, etc).
(gravity gradient)

- The Roche limit depends on nature of body (solid, fluid, density).
- Proposed by Eduard Roche in 1848 (French astronomer).

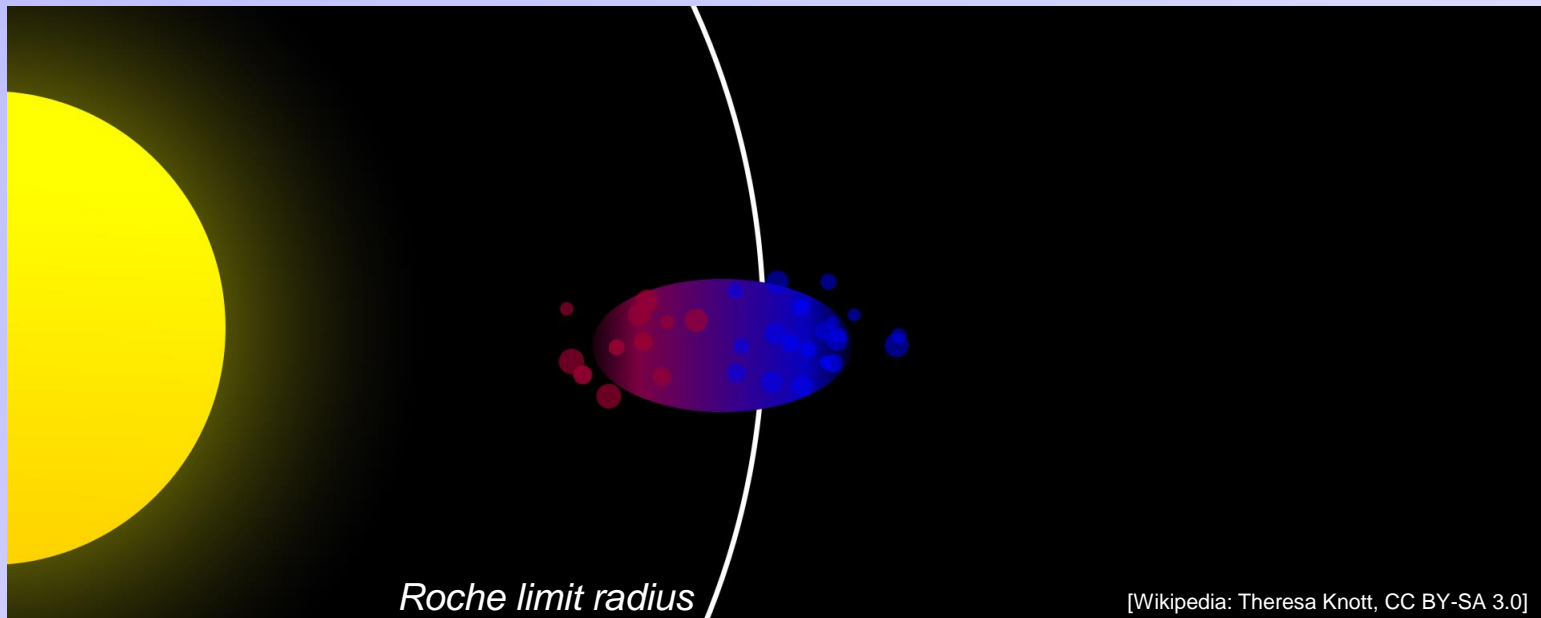


Close to the Roche limit radius, the tidal force and deformation are **strong**.

The Roche Limit

The Roche limit is the orbital radius at which a gravitational bound object will be **pulled apart** by the **tidal force** from the central mass (i.e. Sun, Saturn, etc).
(*gravity gradient*)

- The Roche limit depends on nature of body (solid, fluid, density).
- Proposed by Eduard Roche in 1848 (French astronomer).

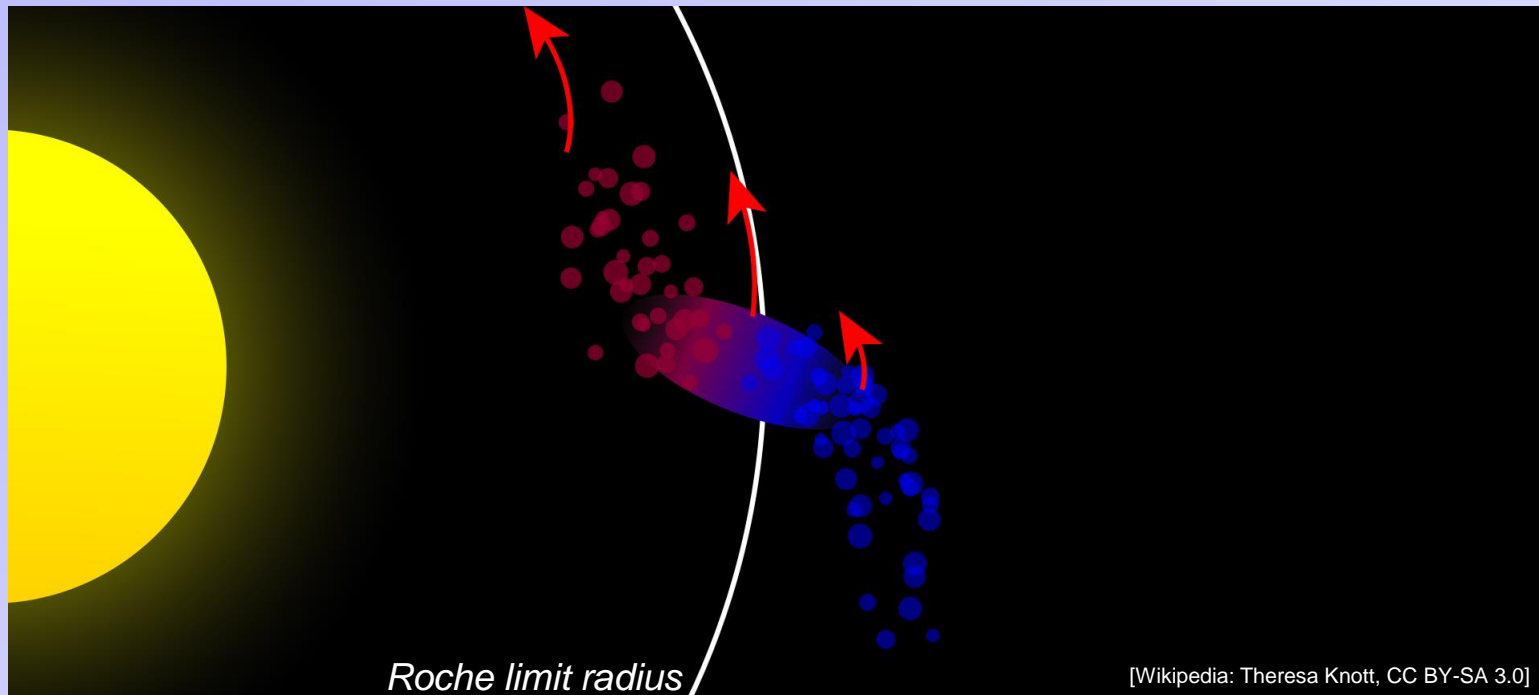


At the Roche limit radius and within it, the tidal force and deformation pull the planet/moon apart.

The Roche Limit

The Roche limit is the orbital radius at which a gravitational bound object will be **pulled apart** by the **tidal force** from the central mass (i.e. Sun, Saturn, etc).
(*gravity gradient*)

- The Roche limit depends on nature of body (solid, fluid, density).
- Proposed by Eduard Roche in 1848 (French astronomer).

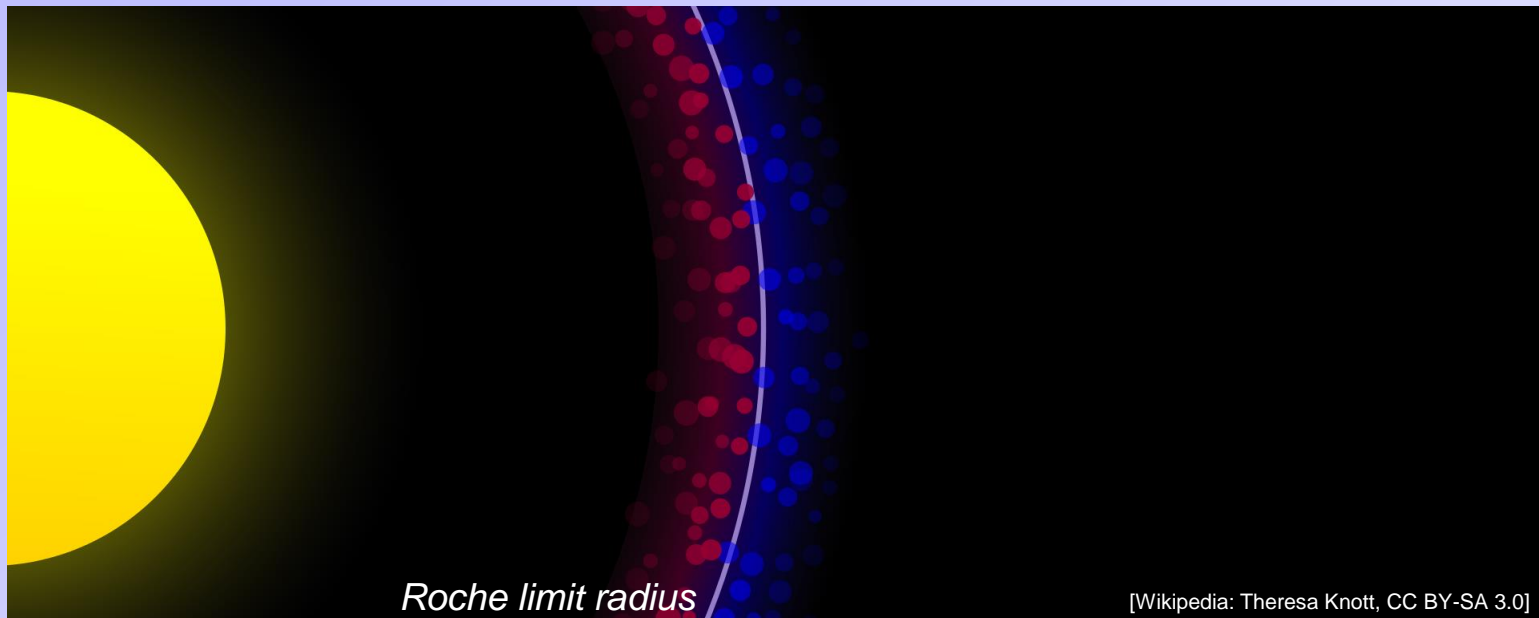


At the Roche limit radius and within it, the tidal force and deformation pull the planet/moon apart.

The Roche Limit

The Roche limit is the orbital radius at which a gravitational bound object will be **pulled apart** by the **tidal force** from the central mass (i.e. Sun, Saturn, etc).
(gravity gradient)

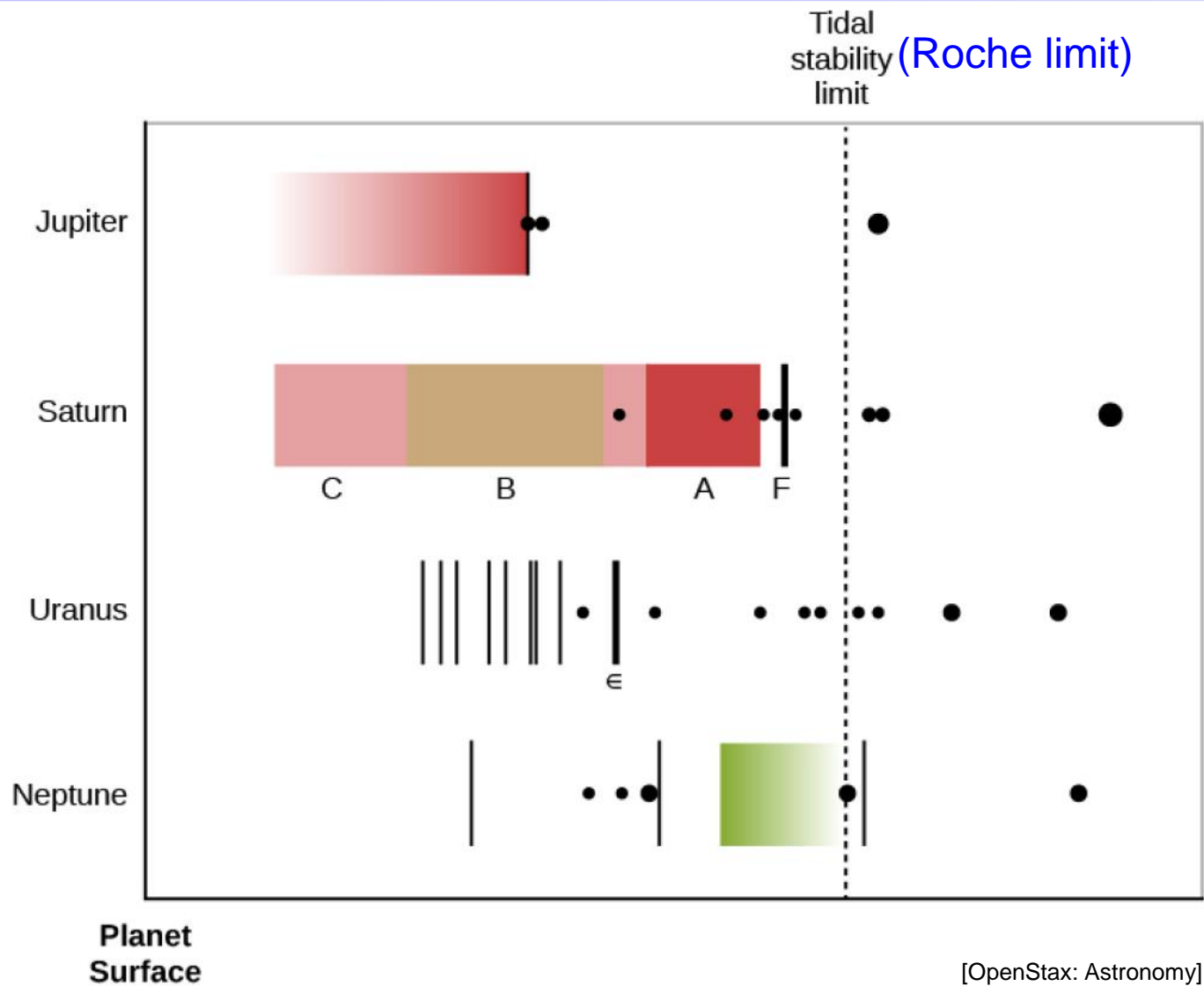
- The Roche limit depends on nature of body (solid, fluid, density).
- Proposed by Eduard Roche in 1848 (French astronomer).



Within the Roche limit radius, planets/moons are pulled apart, and they also cannot form.

- *The accretion process from planetesimals/rock piles cannot happen.*
- *Saturn's rings are unlikely to become moons (Roche limit \approx 75,000-150,000 km).*

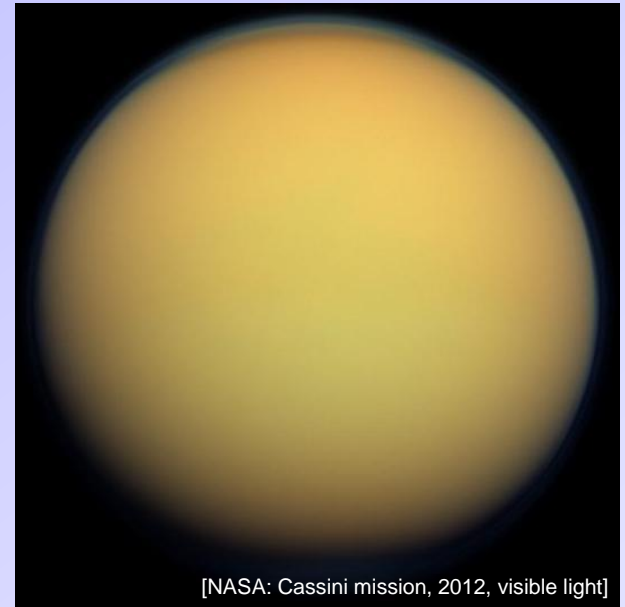
Roche Limit and the Icy Gas Giant



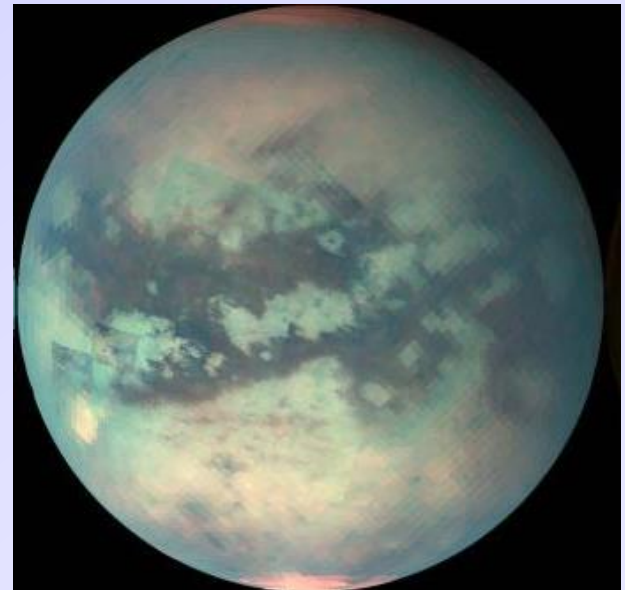
Note: Small moons can survive within the Roche limit.

Titan

- Second largest moon in Solar System.
- **Thick atmosphere:** surface pressure ~ 1.5 bar.
 - Nitrogen (N_2): 95-98%.
 - Methane (CH_4): 2-5%.
 - Other hydrocarbons: trace.
- Temperature: 94 K (surface), i.e. -179° C.
 - **Greenhouse effect** from methane.
 - Anti-greenhouse effect from reflecting smog.



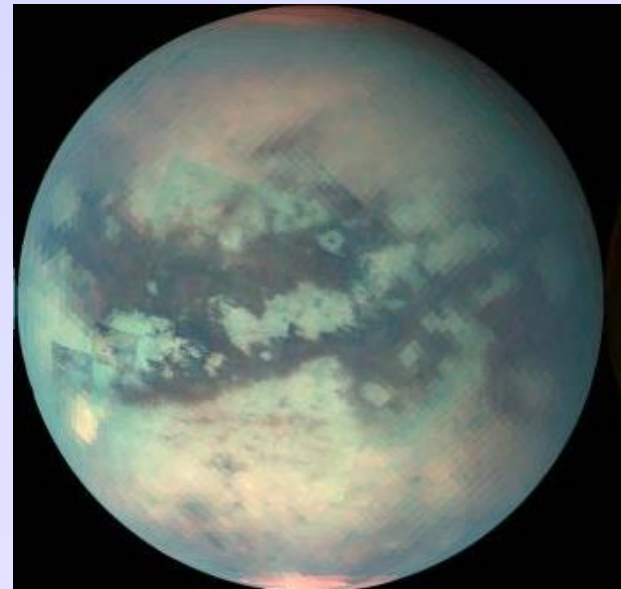
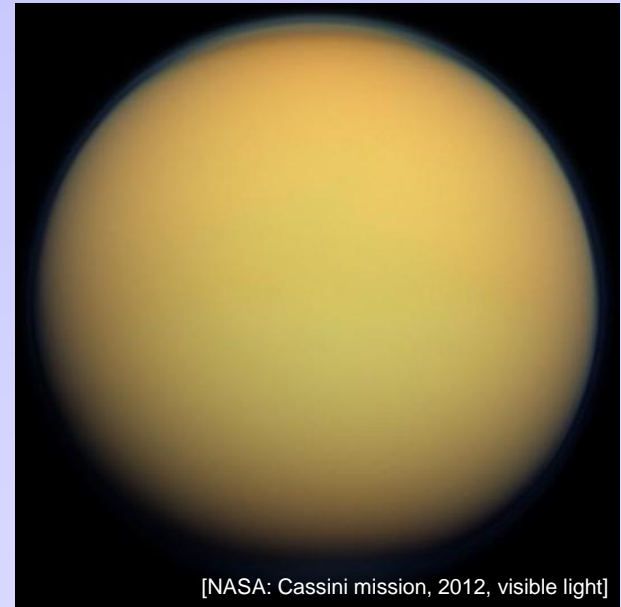
[NASA: Cassini mission, 2012, visible light]



[NASA: Cassini mission, 2005, infrared 1.6-5 μ m]

Titan

- Second largest moon in Solar System.
- **Thick atmosphere:** surface pressure ~ 1.5 bar.
 - Nitrogen (N_2): 95-98%.
 - Methane (CH_4): 2-5%.
 - Other hydrocarbons: trace.
- Temperature: 94 K (surface), i.e. -179° C.
 - **Greenhouse effect** from methane.
 - Anti-greenhouse effect from reflecting smog. } $\Delta T = 12$ K
- **Clouds:** Methane, ethane, other hydrocarbons.
 - Methane rain.
 - Cloud cover: 1-8%.
- Not many craters: Geologically young surface or weathered surface.
 - Surface: Water ice and hydrocarbon ice “rocks”.
 - Interior should be rocky.
- **Hydrocarbon lakes, rivers, seas.**



Titan

- Second largest moon in Solar System.
- **Thick atmosphere:** surface pressure ~ 1.5 bar.
 - Nitrogen (N₂): 95-98%.
 - Methane (CH₄): 2-5%.
 - Other hydrocarbons: trace.
- Temperature: 94 K (surface), i.e. -179° C.
 - **Greenhouse effect** from methane.
 - Anti-greenhouse effect from reflecting smog.
- **Clouds:** Methane, ethane, other hydrocarbons.
 - Methane rain.
 - Cloud cover: 1-8%.
- Not many craters: Geologically young surface or weathered surface.
 - Surface: Water ice and hydrocarbon ice “rocks”.
 - Interior should be rocky.
- **Hydrocarbon lakes, rivers, seas.**

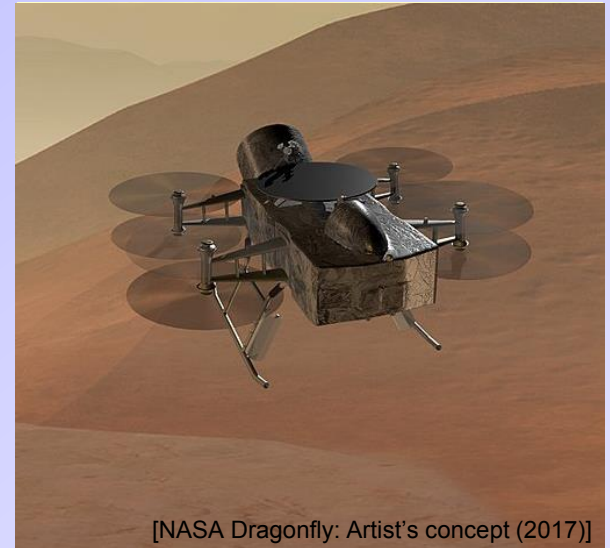


[NASA Cassini mission: ESA Huygens lander, 2005]

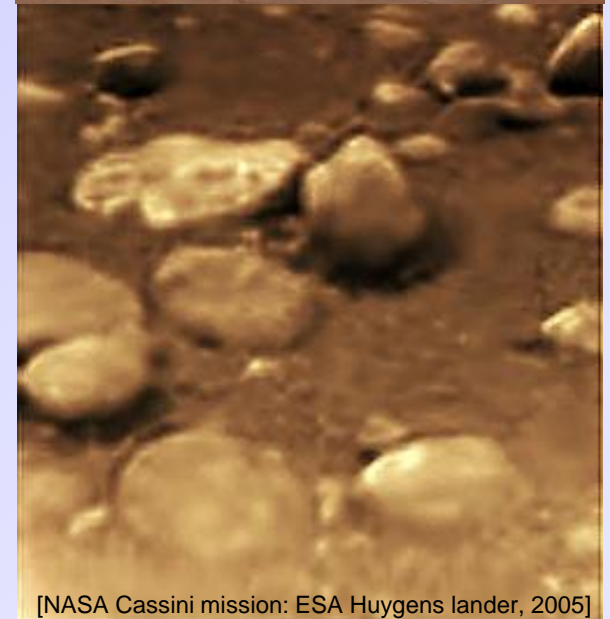
Surface of Titan, 2005 (Huygens lander).

Titan

- Second largest moon in Solar System.
- **Thick atmosphere:** surface pressure ~ 1.5 bar.
 - Nitrogen (N₂): 95-98%.
 - Methane (CH₄): 2-5%.
 - Other hydrocarbons: trace.
- Temperature: 94 K (surface), i.e. -179° C.
 - **Greenhouse effect** from methane.
 - Anti-greenhouse effect from reflecting smog.
- **Clouds:** Methane, ethane, other hydrocarbons.
 - Methane rain.
 - Cloud cover: 1-8%.
- Not many craters: Geologically young surface or weathered surface.
 - Surface: Water ice and hydrocarbon ice “rocks”.
 - Interior should be rocky.
- **Hydrocarbon lakes, rivers, seas.**
- Target for **astrobiology** mission.
 - Dragonfly (NASA/JHU), launch: 2026, arrive: 2034.



[NASA Dragonfly: Artist's concept (2017)]



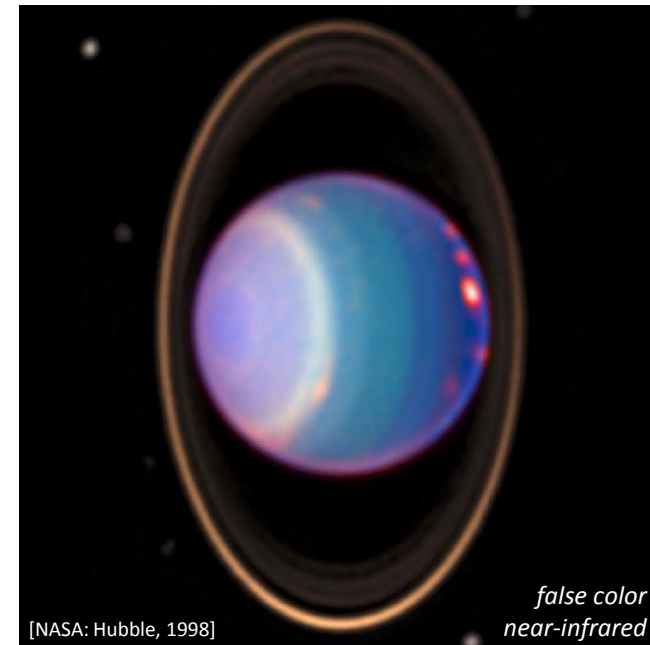
[NASA Cassini mission: ESA Huygens lander, 2005]

Surface of Titan, 2005 (Huygens lander).

Uranus

(“discovered” by William Herschel, 1781)

- **Orbit properties**
 - Semimajor axis = 19.2 AU.
 - Orbital period = 84 years.
- Several times larger than Earth.
 - $M_{\text{Uranus}} = 14.5 M_{\text{Earth}}$
 - $R_{\text{Uranus}} \approx 4 R_{\text{Earth}}$
- **Coldest atmosphere** in Solar System: ~ 49 K.
- Atmosphere: H_2 (83%), He (15%), CH_4 (2%).
- Rotation period $T_{\text{rotation}} = 17$ hrs.



Uranus

(“discovered” by William Herschel, 1781)

- **Orbit properties**

 - Semimajor axis = 19.2 AU.

 - Orbital period = 84 years.

- Several times larger than Earth.

 - $M_{\text{Uranus}} = 14.5 M_{\text{Earth}}$

 - $R_{\text{Uranus}} \approx 4 R_{\text{Earth}}$

- **Coldest atmosphere** in Solar System: ~ 49 K.

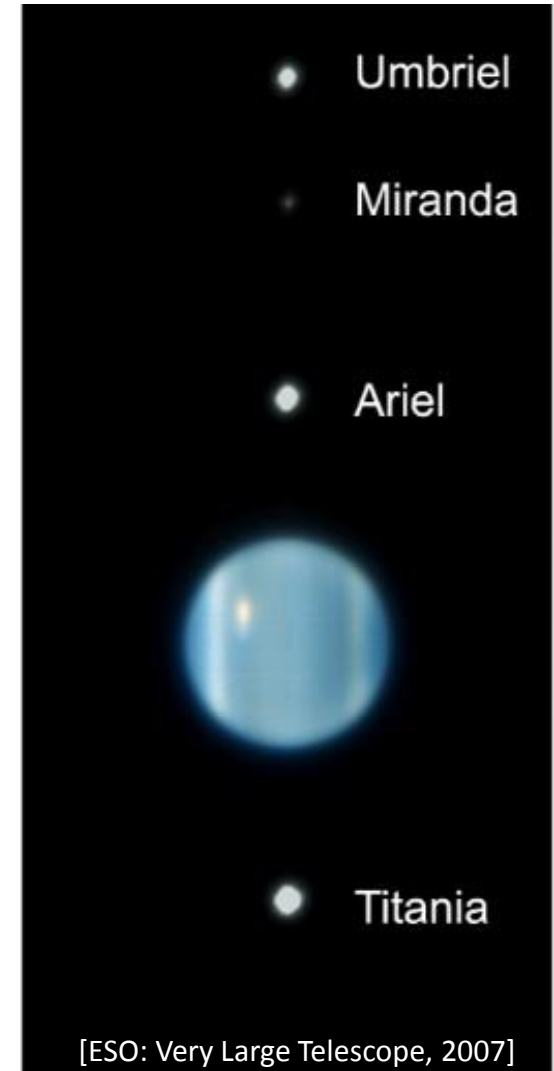
- Atmosphere: H₂ (83%), He (15%), CH₄ (2%).

- Rotation period $T_{\text{rotation}} = 17$ hrs.

- **Rotation axis is almost in orbital plane !**

 - “Planet is on its side.”

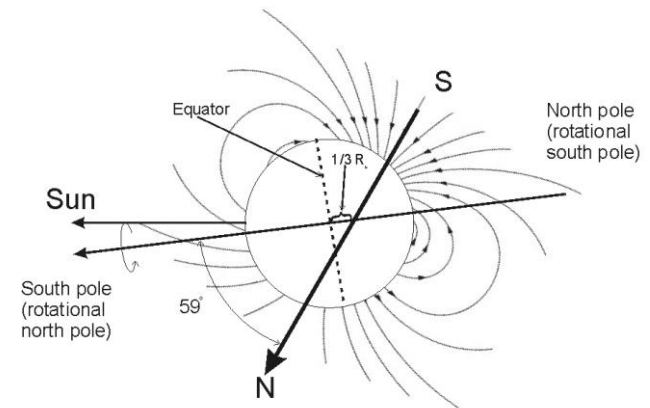
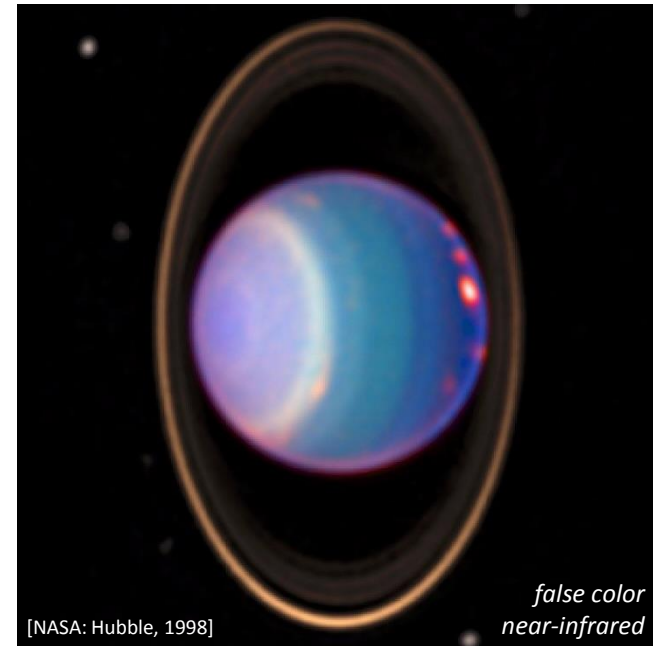
- 27 moons: **Major moons orbit in plane of rotation!**



Uranus

(“discovered” by William Herschel, 1781)

- **Orbit properties**
 - Semimajor axis = 19.2 AU.
 - Orbital period = 84 years.
- Several times larger than Earth.
 - $M_{\text{Uranus}} = 14.5 M_{\text{Earth}}$
 - $R_{\text{Uranus}} \approx 4 R_{\text{Earth}}$
- **Coldest atmosphere** in Solar System: ~ 49 K.
- Atmosphere: H₂ (83%), He (15%), CH₄ (2%).
- Rotation period $T_{\text{rotation}} = 17$ hrs.
- **Rotation axis is almost in orbital plane !**
 - “Planet is on its side.”
- 27 moons: **Major moons orbit in plane of rotation!**
- **Magnetosphere** is **tilted 59°** with respect to rotation axis and is **off-center !!!**



[Wikipedia, based on Voyager 2 data]

Uranus

(“discovered” by William Herschel, 1781)

- **Orbit properties**

Semimajor axis = 19.2 AU.

Orbital period = 84 years.

- Several times larger than Earth.

→ $M_{\text{Uranus}} = 14.5 M_{\text{Earth}}$

→ $R_{\text{Uranus}} \approx 4 R_{\text{Earth}}$

- **Coldest atmosphere** in Solar System: ~ 49 K.

- Atmosphere: H₂ (83%), He (15%), CH₄ (2%).

- Rotation period $T_{\text{rotation}} = 17$ hrs.

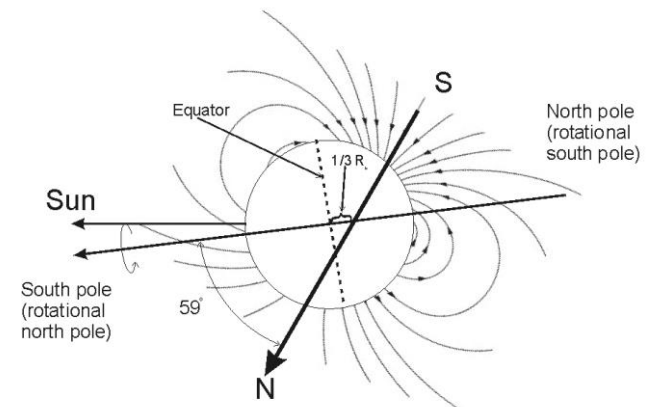
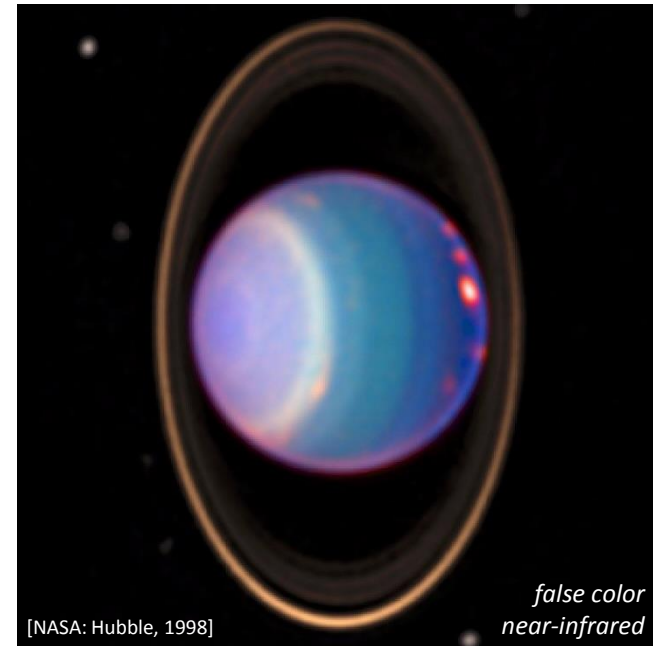
- **Rotation axis is almost in orbital plane !**

→ “Planet is on its side.”

- 27 moons: **Major moons orbit in plane of rotation!**

- **Magnetosphere** is **tilted 59°** with respect to rotation axis and is **off-center !!!**

Uranus’s strange orientation is thought to be due to a collision with an Earth-sized protoplanet during formation of Solar System.

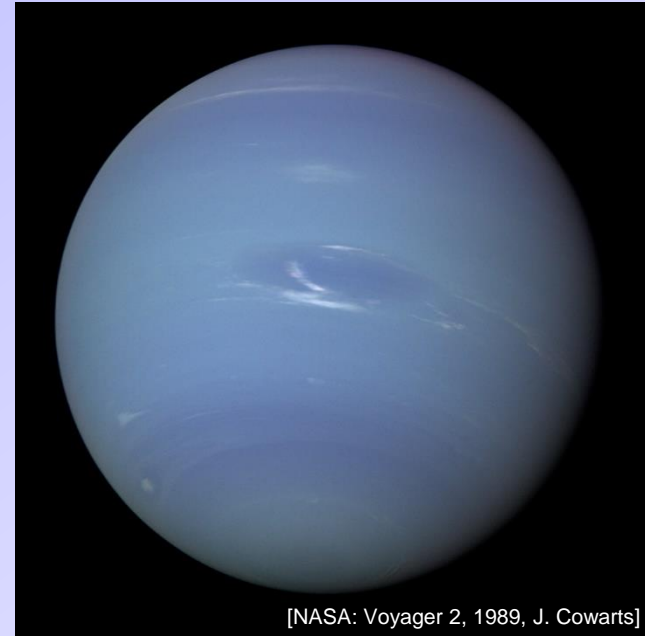


[Wikipedia, based on Voyager 2 data]

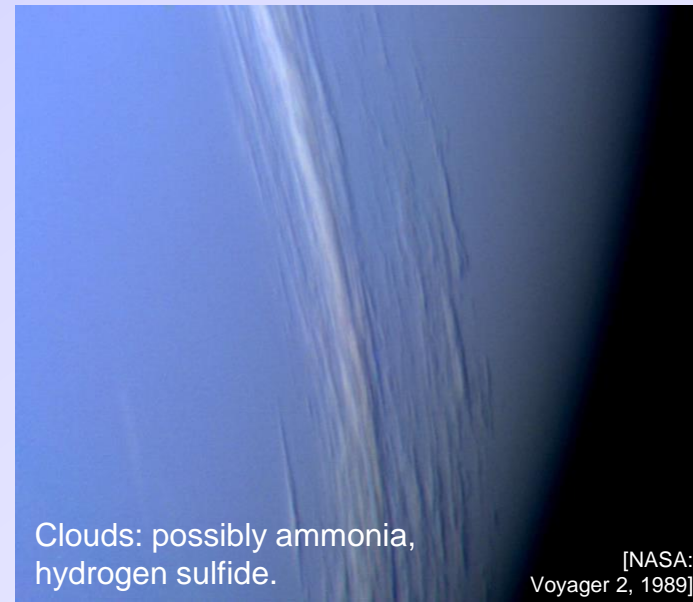
Neptune

Neptune was “discovered” theoretically from its gravitational effect on Uranus’s orbit.

- Probably observed by **Galileo** 1612-13 (retrograde).
- **Alexis Bouvard** noticed irregularities in Uranus’s motion (1820’s).
- **Urbain Le Verrier** and **John Couch Adams** independently predict location of Neptune (1845-46).
- **Johann Gottfried Galle** observed Neptune within 1° of **Le Verrier’s** prediction (1846).



[NASA: Voyager 2, 1989, J. Cowarts]



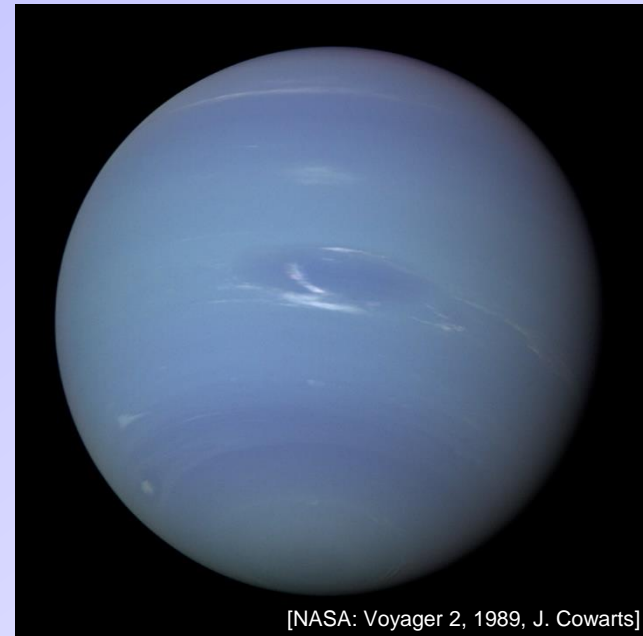
Clouds: possibly ammonia, hydrogen sulfide.

[NASA: Voyager 2, 1989]

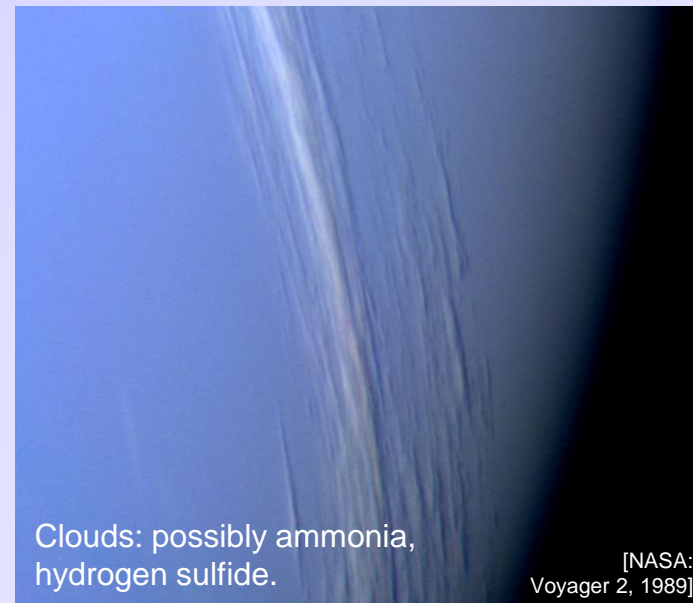
Neptune

Neptune was “discovered” theoretically from its gravitational effect on Uranus’s orbit.

- Probably observed by **Galileo** 1612-13 (retrograde).
- **Alexis Bouvard** noticed irregularities in Uranus’s motion (1820’s).
- **Urbain Le Verrier** and **John Couch Adams** independently predict location of Neptune (1845-46).
- **Johann Gottfried Galle** observed Neptune within 1° of **Le Verrier’s** prediction (1846).
- Semimajor axis = 30 AU, orbital period = 165 yrs.
- 14 moons. Largest is Triton (retrograde orbit).
- **Similar** in size and composition **to Uranus**.
 - $M_{\text{Neptune}} = 1.2 M_{\text{Uranus}}$
 - $R_{\text{Neptune}} = 0.97 R_{\text{Uranus}}$
 - Rotation: $T_{\text{rotation}} = 16 \text{ hrs} = 0.93 T_{\text{Uranus}}$
 - Atmosphere: H_2 (80%), He (19%), CH_4 (1.5%).
 - Temperature: 55 K (a tad warmer than Uranus).



[NASA: Voyager 2, 1989, J. Cowarts]



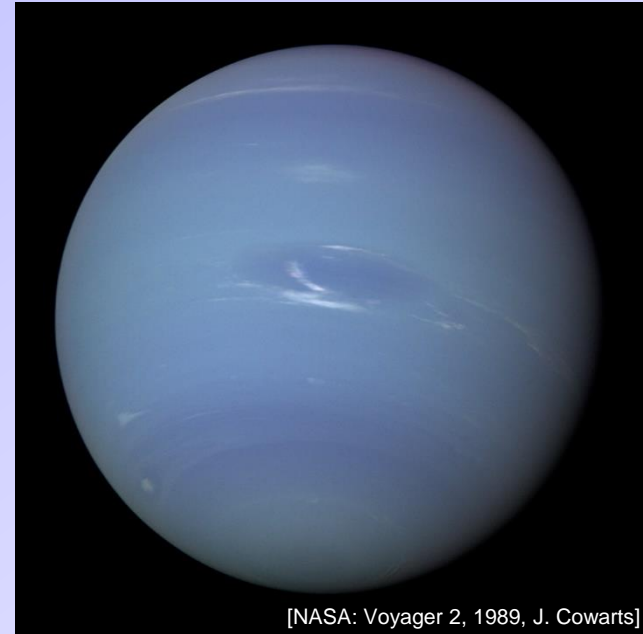
Clouds: possibly ammonia, hydrogen sulfide.

[NASA: Voyager 2, 1989]

Neptune

Neptune was “discovered” theoretically from its gravitational effect on Uranus’s orbit.

- Probably observed by **Galileo** 1612-13 (retrograde).
- **Alexis Bouvard** noticed irregularities in Uranus’s motion (1820’s).
- **Urbain Le Verrier** and **John Couch Adams** independently predict location of Neptune (1845-46).
- **Johann Gottfried Galle** observed Neptune within 1° of **Le Verrier’s** prediction (1846).
- Semimajor axis = 30 AU, orbital period = 165 yrs.
- 14 moons. Largest is Triton (retrograde orbit).
- **Similar** in size and composition **to Uranus**.
 - $M_{\text{Neptune}} = 1.2 M_{\text{Uranus}}$
 - $R_{\text{Neptune}} = 0.97 R_{\text{Uranus}}$
 - Rotation: $T_{\text{rotation}} = 16 \text{ hrs} = 0.93 T_{\text{Uranus}}$
 - Atmosphere: H_2 (80%), He (19%), CH_4 (1.5%).
 - Temperature: 55 K (a tad warmer than Uranus).
- Planet’s axis is tilted 28° , but **magnetosphere is tilted 47°** to this axis and also off center.



[NASA: Voyager 2, 1989, J. Cowarts]



Clouds: possibly ammonia, hydrogen sulfide.

[NASA: Voyager 2, 1989]