Today's Topics

Monday, October 19, 2020 (Week 9, lecture 25) – Chapters 14.4, 30.1-3.

- A. 2019 Nobel Prize in Physics
 - → Discovery of Exoplanets
- **B.** Exoplanets
- C. Life in the Solar System?



James Peebles (Princeton U.)



Michel Mayor (U. of Geneva)



Didier Queloz (U. of Geneva)



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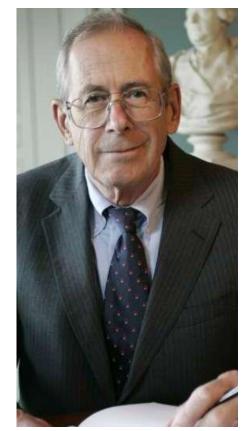
Didier Queloz (U. of Geneva)

Big Bang theory, dark matter & energy

Discovery of first star-based **exo-planet** (in 1995)

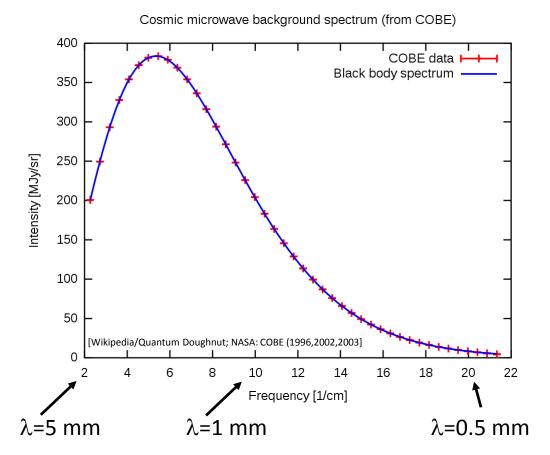
[Image credits: Royal Swedish Academy of Sciences; University of Geneva]

- Cosmic Microwave Background (CMB) is the remnant <u>blackbody</u> radiation from the Big Bang.
- This radiation used to much more energetic, but it has been "Doppler-shifted" by the expansion of the universe. Apparent temperature is **T** = **2.726** K (peak at $\lambda = 1$ mm).



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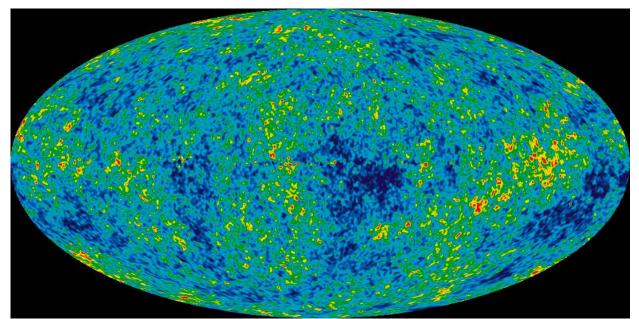




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- Spatial/angular variations of CMB ($\sim 0.01\%$ level) are thought to have given rise to galactic superclusters, galaxies, and stars.



[https://map.gsfc.nasa.gov/media/121238/index.html]

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Spatial variations in the Cosmic Microwave Background (WMAP).

[Image credits: Royal Swedish Academy of Sciences: University of Geneval

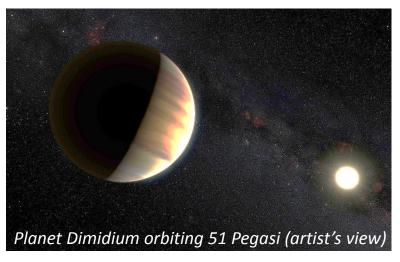
- 1995: First discovery of a planet orbiting another star (51 Pegasi) at about 50 ly.
 - → Pegasus constellation.
- Discovered by seeing the **Doppler shift** in the star's light caused by orbiting the center-of-mass of the star-planet system.
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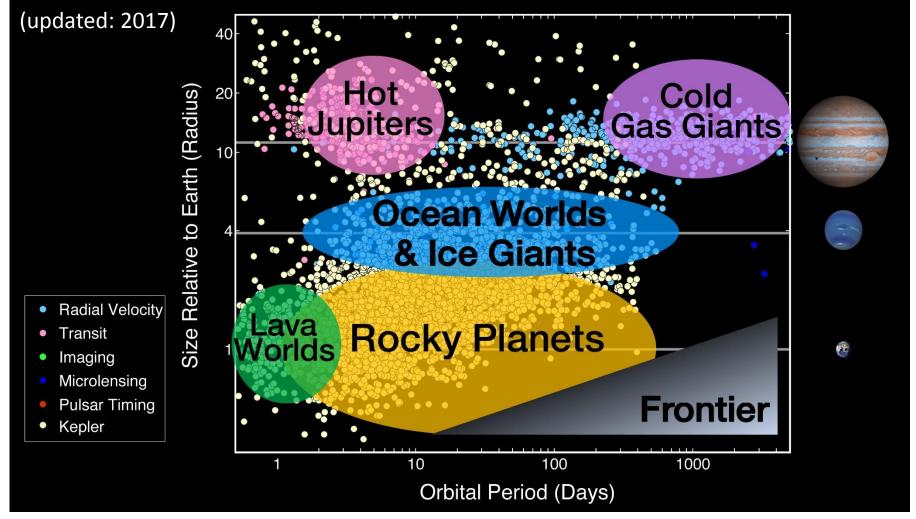


Discovery launched exo-planet astronomy.

→ ~ 4000 exo-planets discovered so far.

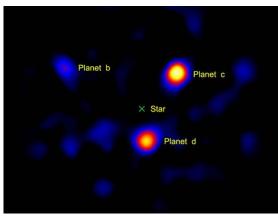
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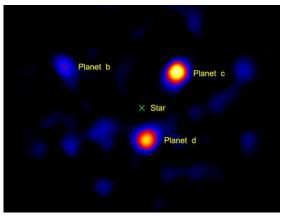


3 planets around star HR8799 (120 ly) Orbits: 24 AU, 38 AU, 68 AU. [Hale telescope, 2010]

ExoplanetsWhat we know so far

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- We see many gas giants inside the frost line.

Models of evolution for solar systems show that planets often perturb the orbits of other planets and **move them towards the star** (or shoot them out).



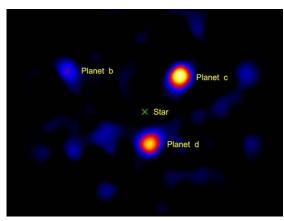
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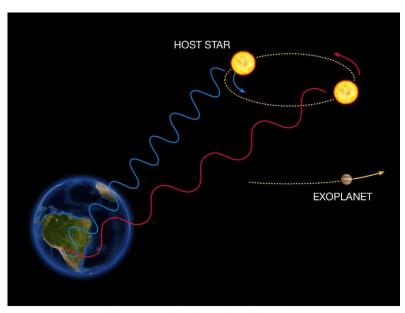
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- Roughly 40% of Sun-like stars have terrestrial planets in the "goldilocks" region.
 - → Above freezing and below boiling for water.
- Earth-like planets are very common

They are harder to detect than larger ones, so we have not seen very many yet.



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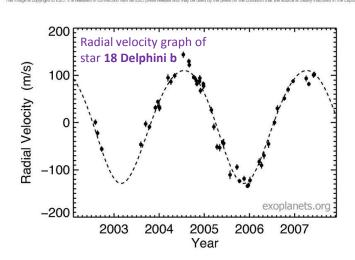
Main Detection Methods



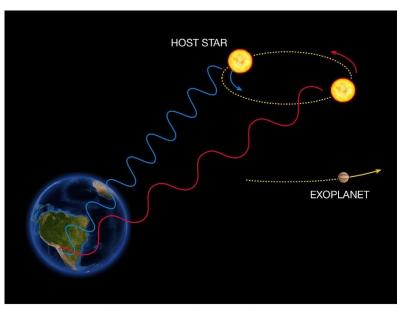
The Radial Velocity Method

ESO Press Photo 22e/07 (25 April 2007)





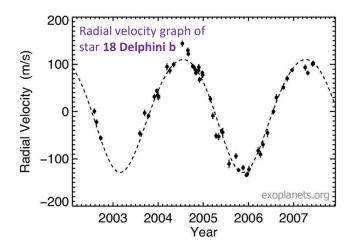
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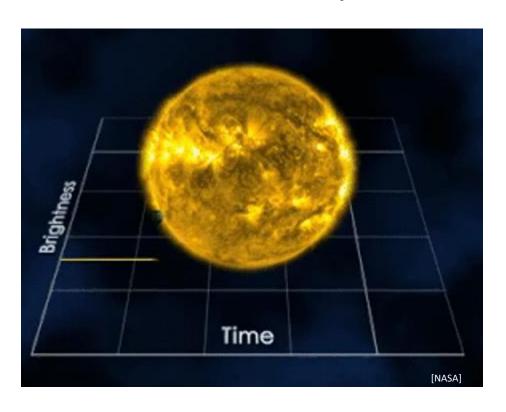
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Transit Photometry



Signal is typically 1 part per 10,000 dimming.

Definitions of Life (biology) – there are many definitions (no consensus)

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Life is considered a characteristic of something that preserves, furthers or reinforces its existence in the given environment [Wikipedia].

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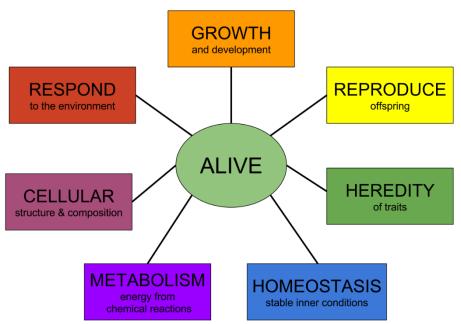
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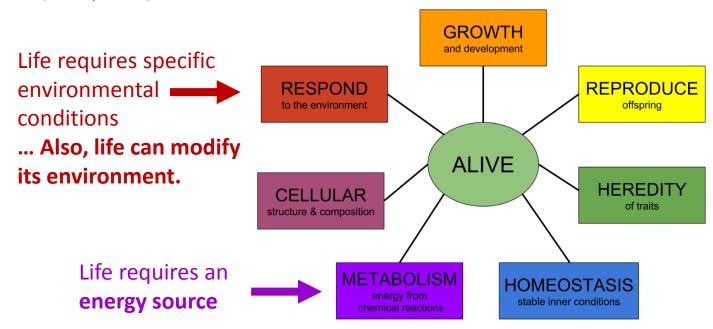
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What are the environmental requirements for Earth-style Life?

Ingredients for Earth-style life

- Liquid water (solvent).
- Carbon, hydrogen, nitrogen, oxygen, phosphorus, and sulfur.
- Temperature range: -25° C to 120° C.
- Not too much ionizing radiation.
- An energy source: solar energy, chemical energy, heat.

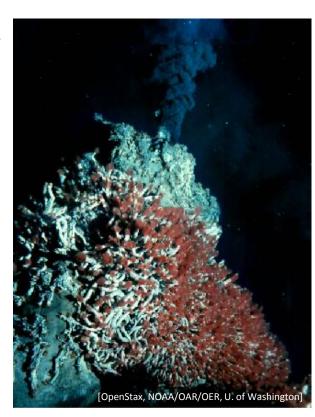
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Hot spring (Yellowstone Nat. Park): multi-colored photosynthesizing bacteria (solar energy). Central pool (92 $^\circ$ C) bacteria use chemical energy.



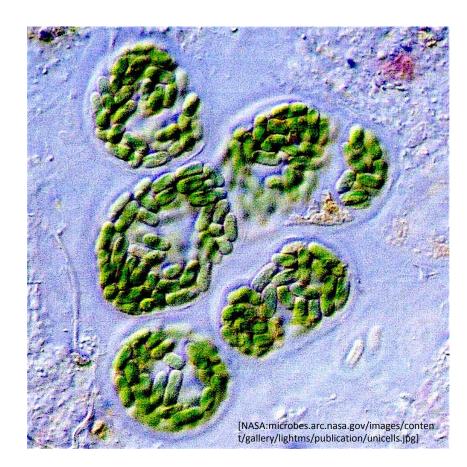
Hydrothermal vent: chemical energy (metal sulfides).

Examples of Hardy Life



The Tardigrade (length: ~0.5 mm) can survive

- -270° C to 150° C for a few minutes.
- Vacuum to 1000 bars.
- Years of dehydration.
- 1000x lethal radiation dose of animals.



Cyanobacteria are photosynthesizing bacteria that have existed on Earth for 2.1 billion years ago (maybe longer)

Where to look for life In the Solar System?



Mars may have suitable conditions for life just below its surface:

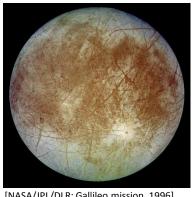
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[NASA/JPL/DLR: Gallileo mission, 1996]

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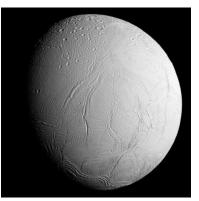
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(Saturn **Enceladus**

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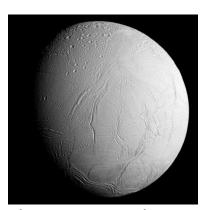
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More speculative: Ganymede sub-surface water, upper atmosphere of Venus.