

Today's Topics

Friday, February 14, 2025 (Week 3, lecture 10) – Chapter 5.

A. Light particles: photons

B. Electronic structure of atoms & spectroscopy

C. Doppler effect

D. Nuclear particles

REMINDER #1: **Problem Set #4** is due on ExpertTA by **Wednesday**, February 19, 9:00 AM.

REMINDER #2: **Midterm #1** is on Friday, February 21.

Light is also a particle: the Photon

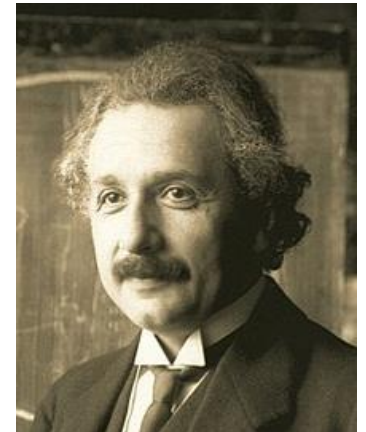
- **Max Planck** (1858-1947) figured out that light also behaves as a **particle** using **blackbody radiation**.
- Albert Einstein (1879-1955) also figured out that light behaves as a **particle** based on the **photo-electric effect**.
- **Light particle = photon = packet of EM energy**

- Energy = hf (f is the frequency)
 h = Planck's constant = 6.626×10^{-34} J·s

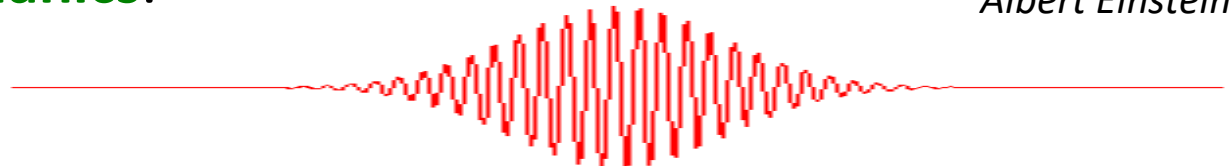
- Discovery of the photon helped initiate **Quantum Mechanics**.



Max Planck

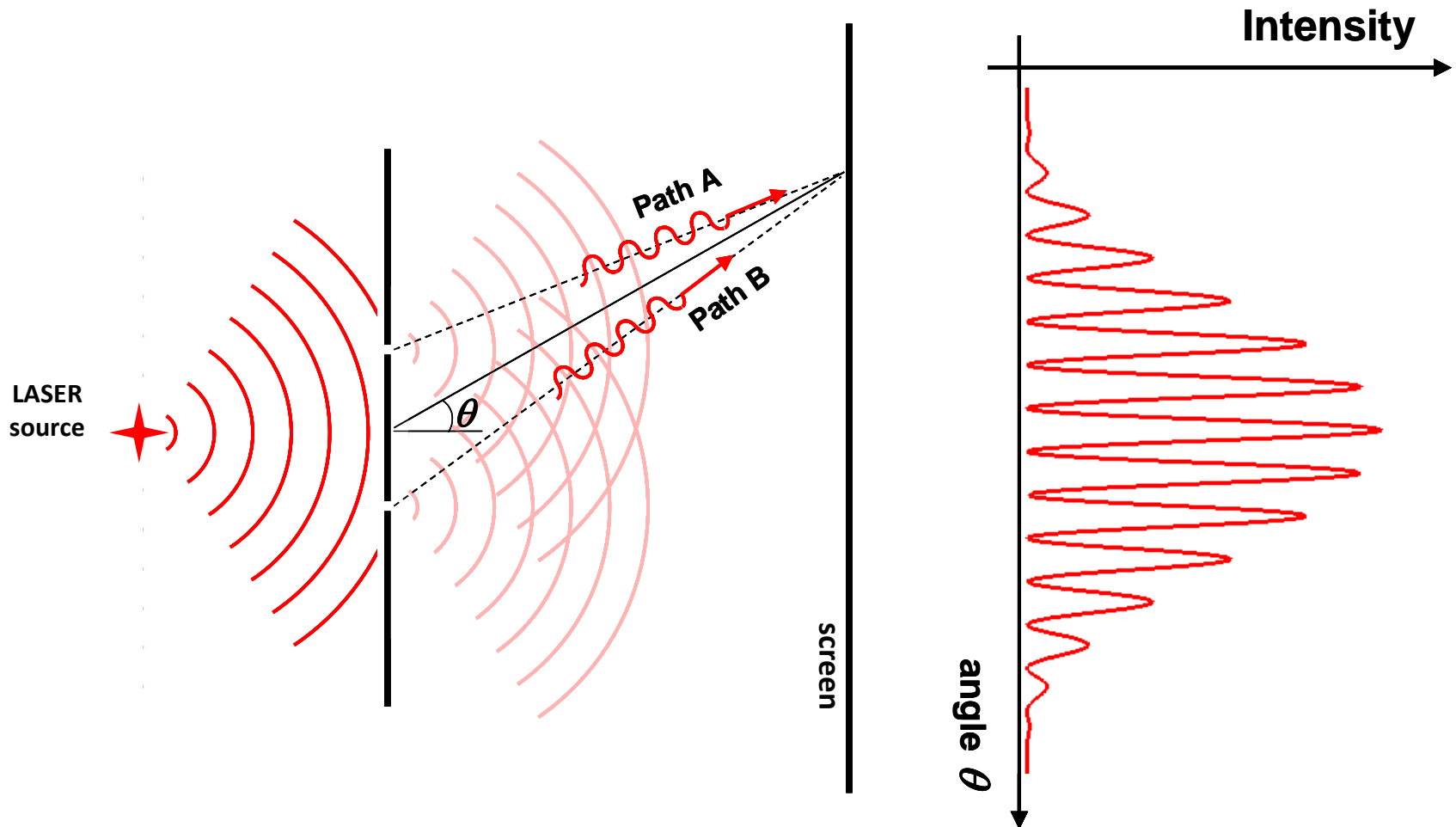


Albert Einstein

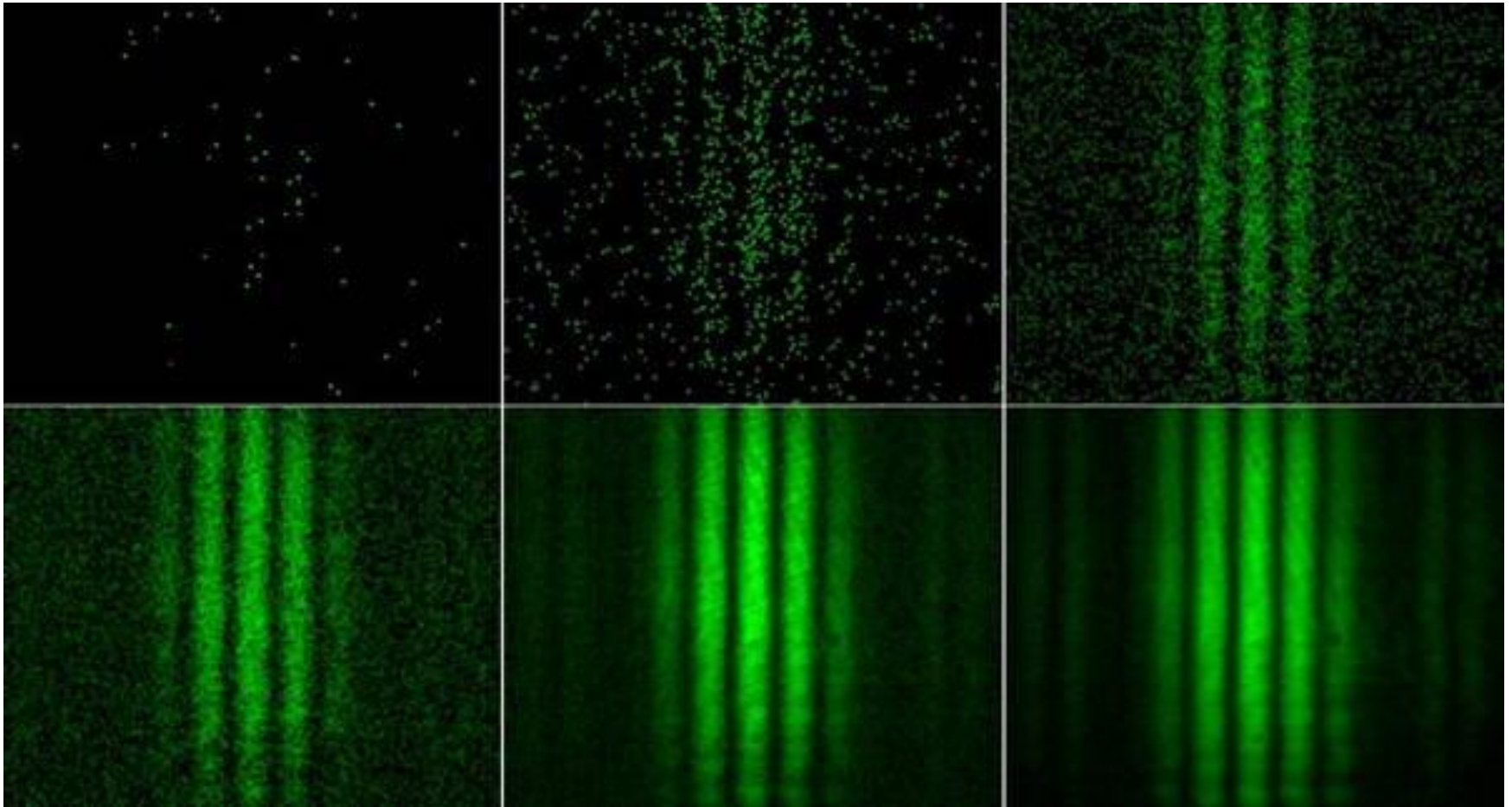


2-slit Interference Experiment

(proof that light is a wave)



Repeat Interference Experiment one photon at a time



[A. L. Weiss and T. L. Dimitrova, Swiss Physics Society, 2009.]

LIGHT IS A

WAVE!

Photon Basics

$$\text{Photon energy} = E_{\gamma} = hf$$

Important: Photon is massless

$$M_{\gamma} = 0$$

Photon Basics

$$\text{Photon energy} = E_{\gamma} = hf$$

$$\text{Photon momentum} = p_{\gamma} = \frac{h}{\lambda} = \frac{E_{\gamma}}{c}$$

Important: Photon is massless

$$M_{\gamma} = 0$$

Note: $p_{\gamma} \neq M_{\gamma}c$ ($= 0$)

Photon Basics

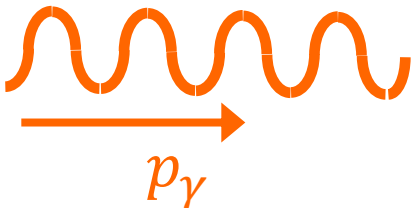
$$\text{Photon energy} = E_\gamma = hf$$

Important: Photon is massless
 $M_\gamma = 0$

$$\text{Photon momentum} = p_\gamma = \frac{h}{\lambda} = \frac{E_\gamma}{c}$$

Note: $p_\gamma \neq M_\gamma c (= 0)$

Shine light on an atom



$$p_{atom} = 0$$

Photon Basics

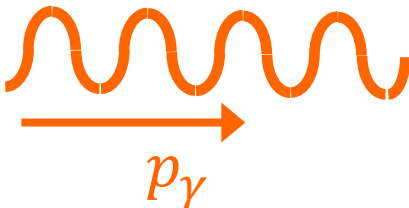
$$\text{Photon energy} = E_\gamma = hf$$

Important: Photon is massless
 $M_\gamma = 0$

$$\text{Photon momentum} = p_\gamma = \frac{h}{\lambda} = \frac{E_\gamma}{c}$$

Note: $p_\gamma \neq M_\gamma c (= 0)$


Shine light on an atom



$$p_{atom} = 0$$

Atom after absorption of 1 photon



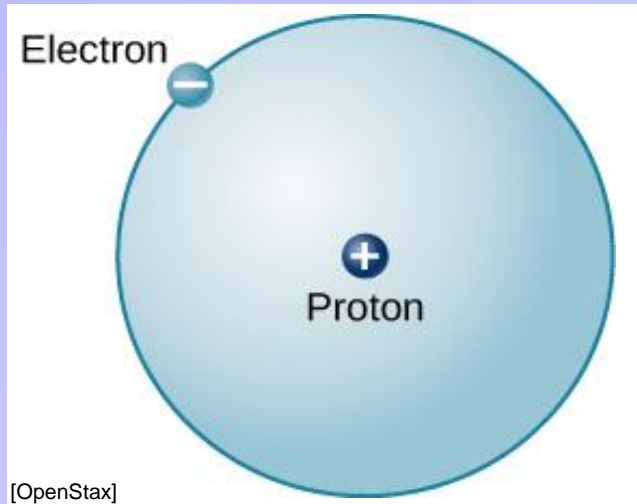


A solid black arrow pointing to the right, labeled with the symbol $p_{atom} = p_\gamma$ in black.

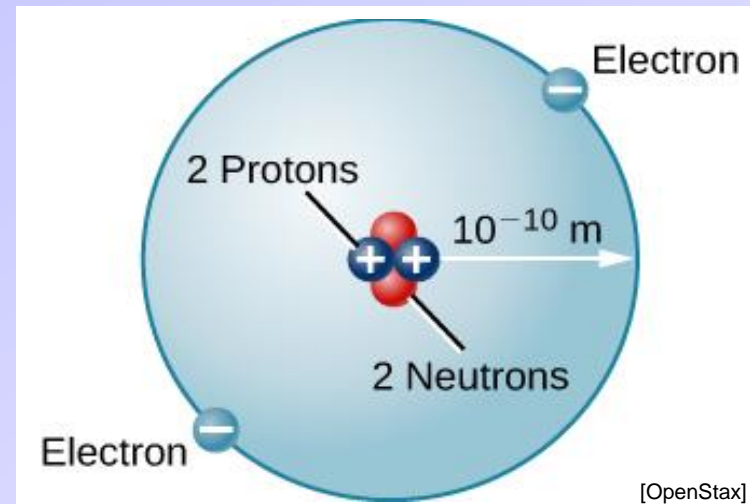
(conservation of momentum)

Basic Structure of Atoms

Hydrogen: ${}^1\text{H}$



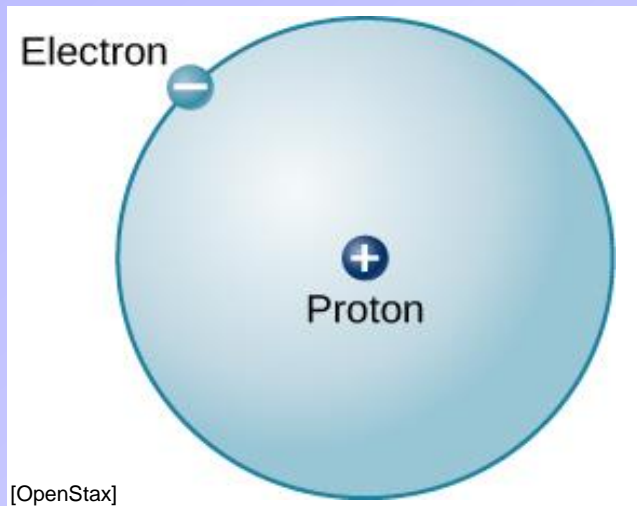
Helium: ${}^4\text{He}$ $4 = 2 \text{ protons} + 2 \text{ neutrons}$



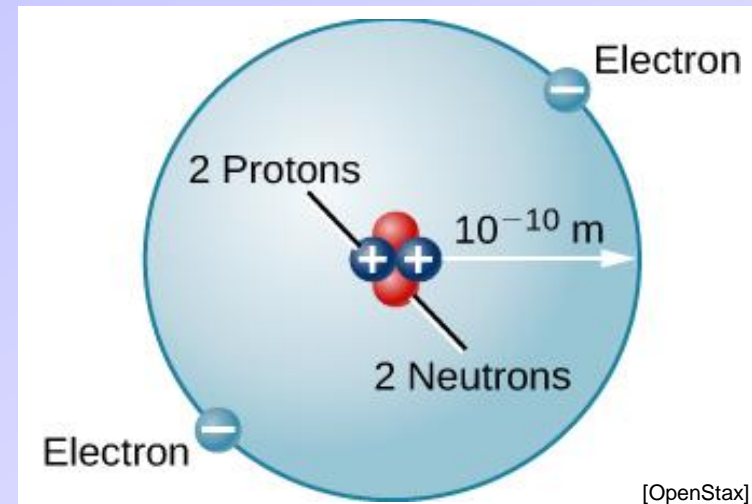
- Atom consist of **positively charged nucleus** orbited by **negatively charged electrons** (for neutral atoms: total charge = zero).
- Electron number, orbits, and properties determine the **chemistry** of the atom.

Basic Structure of Atoms

Hydrogen: ${}^1\text{H}$



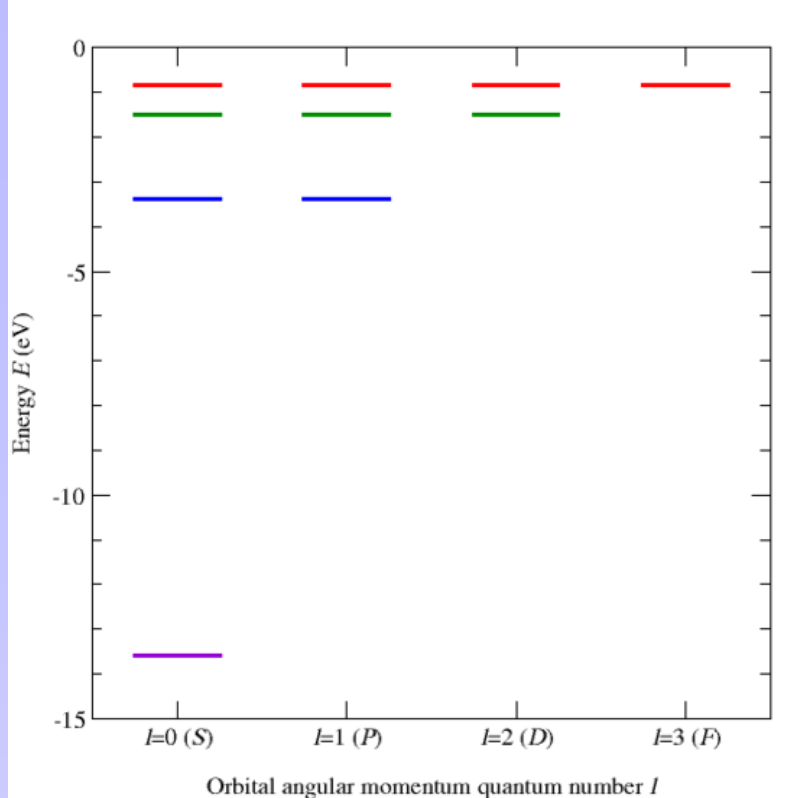
Helium: ${}^4\text{He}$ $4 = 2 \text{ protons} + 2 \text{ neutrons}$



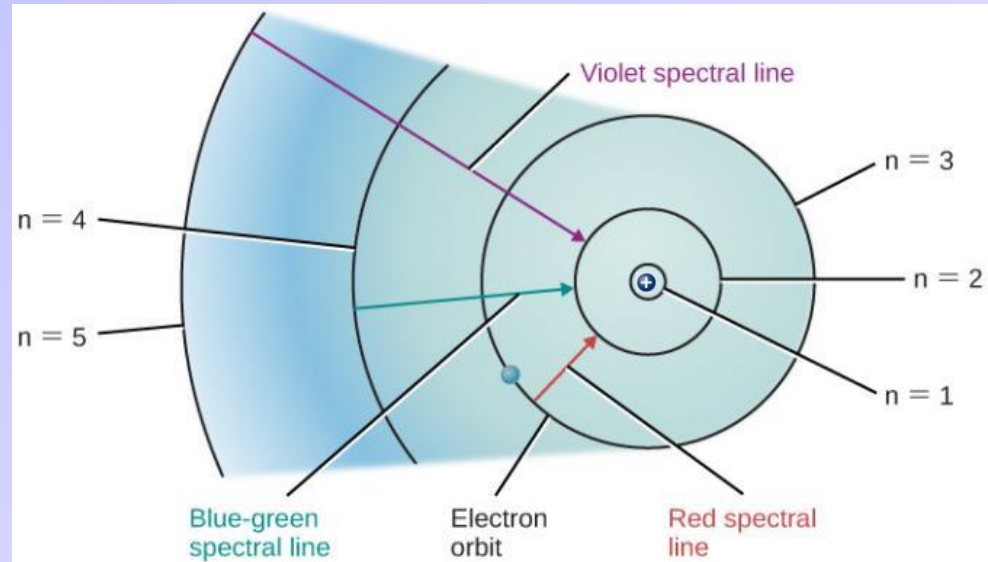
- Atom consist of **positively charged nucleus** orbited by **negatively charged electrons** (for neutral atoms: total charge = zero).
- Electron number, orbits, and properties determine the **chemistry** of the atom.
- **Nucleus** consists of positively charged **protons** and **neutral neutrons**.
- For neutral atoms: **Number of protons = number of electrons**.
- Neutrons help bind protons together. **Number of neutrons \geq number of protons**.

Electronic Structure of Atoms

Energy Levels of Hydrogen ($n=1-4$)



[Figure from wikimedia.org]



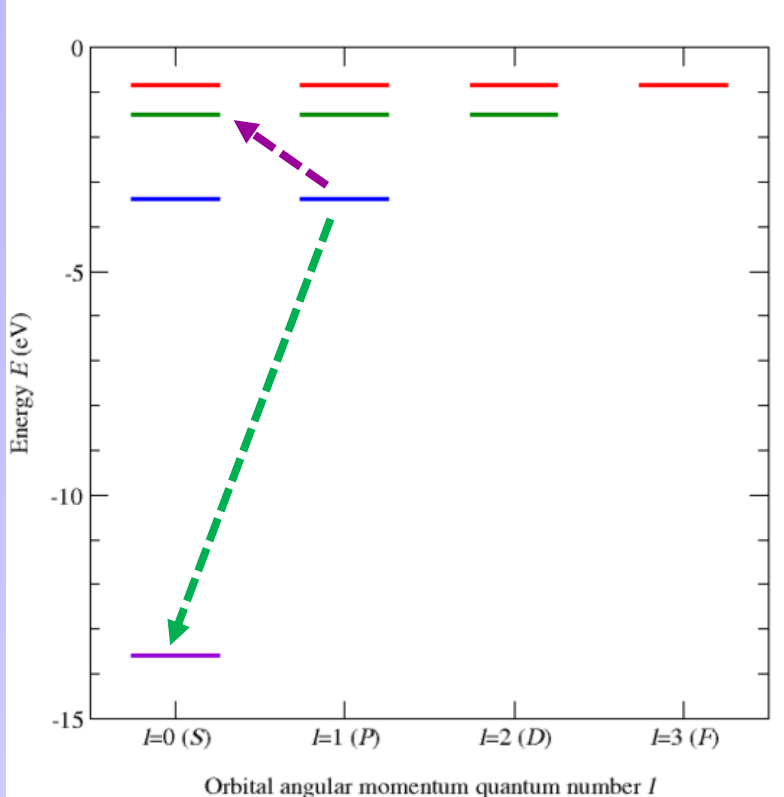
[OpenStax]

- Electrons have **discrete allowed energies and orbits**.

Note: $1 \text{ eV} = 1.602 \times 10^{-19} \text{ J}$

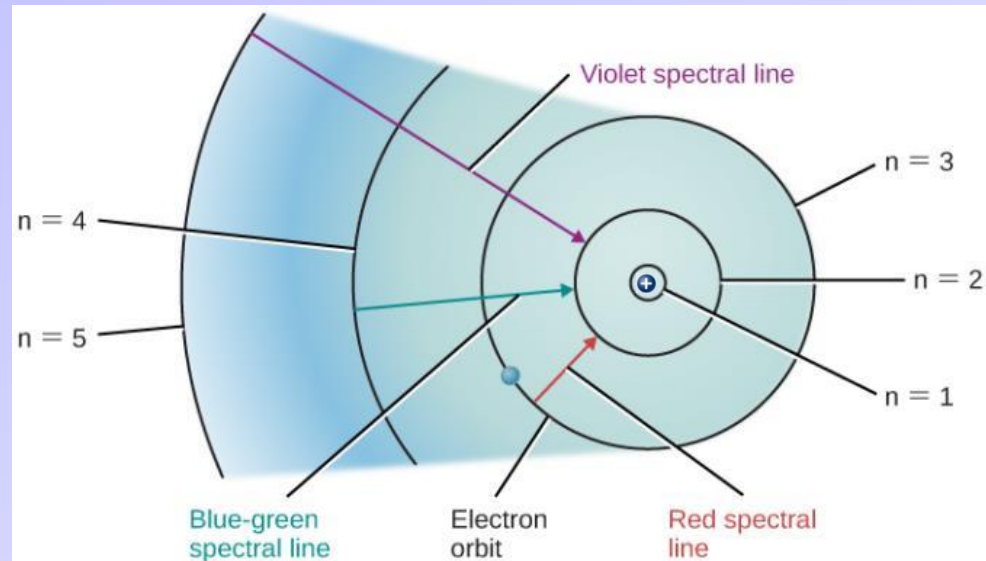
Electronic Structure of Atoms

Energy Levels of Hydrogen ($n=1-4$)



[Figure from wikimedia.org]

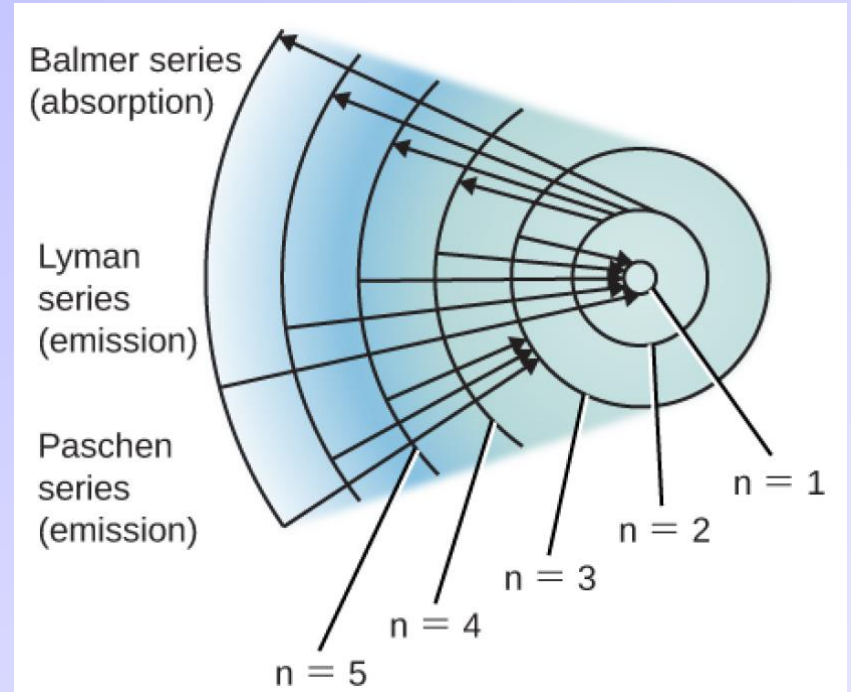
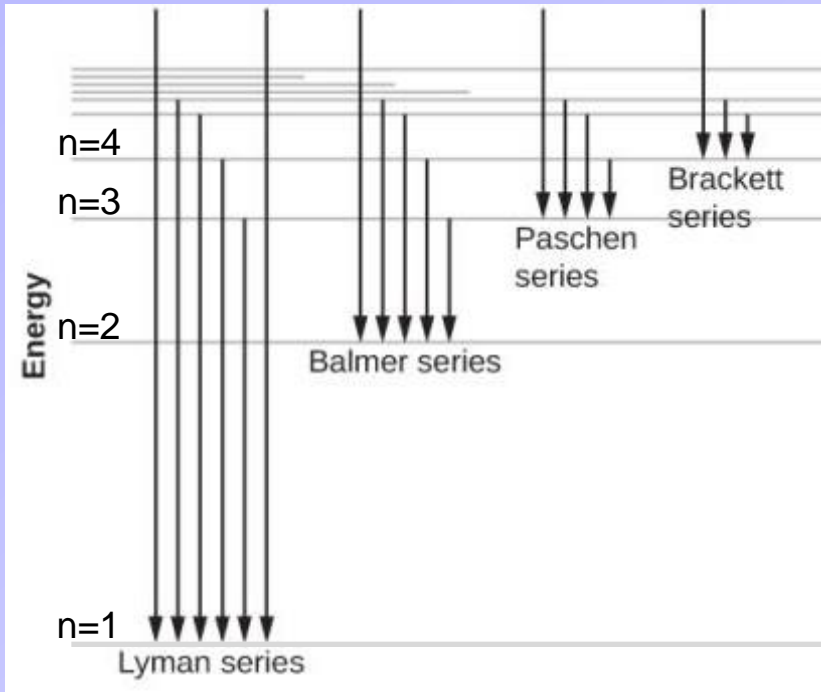
Note: $1 \text{ eV} = 1.602 \times 10^{-19} \text{ J}$



[OpenStax]

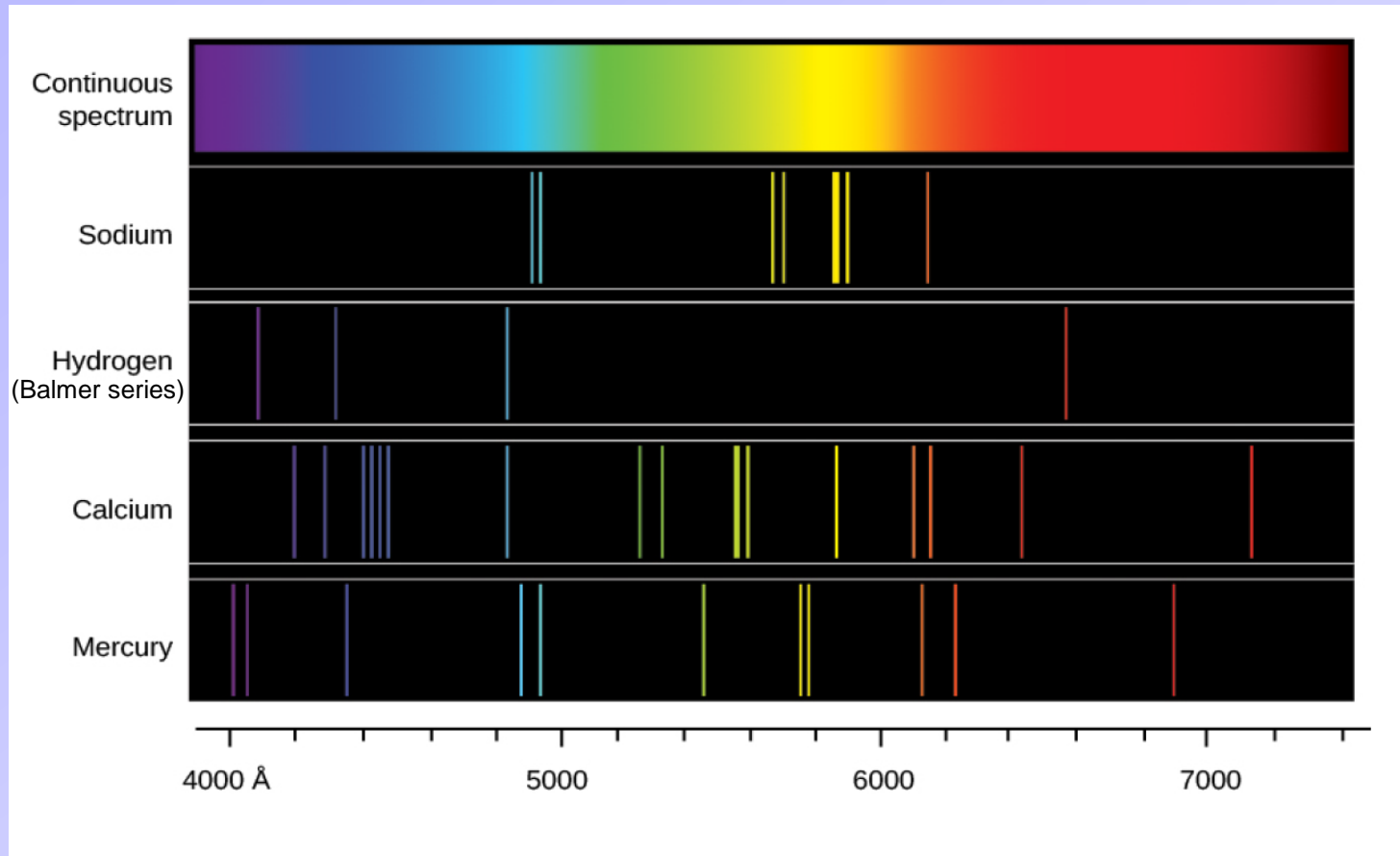
- Electrons have **discrete allowed energies and orbits**.
- Transitions between two energy levels requires **emission** or **absorption** of a photon that bridges the energy gap.
- **Discrete** emission and absorption **spectra**.

Emission Spectrum of Hydrogen



- Hydrogen has a number of emission and absorption spectral series that depend on the start/end point of the transition.
- Other elements are qualitatively similar.
- Also true for molecules, but their spectra are more complicated.

Emission Spectra “Fingerprints”



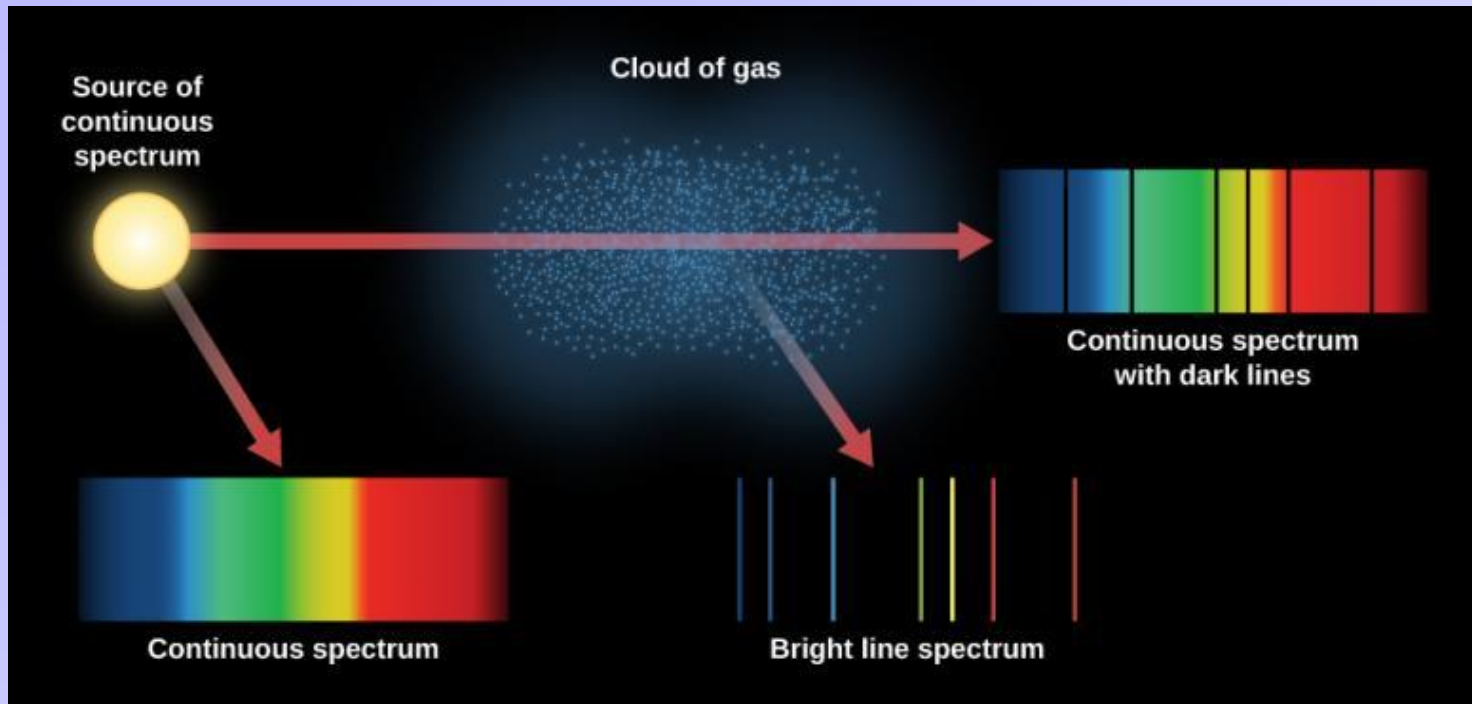
If you build a **catalog of spectral lines**, then you can **determine the elements** that are present from the spectrum.

PollEv Quiz: PollEv.com/sethaubin

Identify the element in the lamp:

- A.** Sodium.
- B.** Hydrogen.
- C.** Calcium.
- D.** Mercury

Emission & Absorption Processes



Three types of spectra

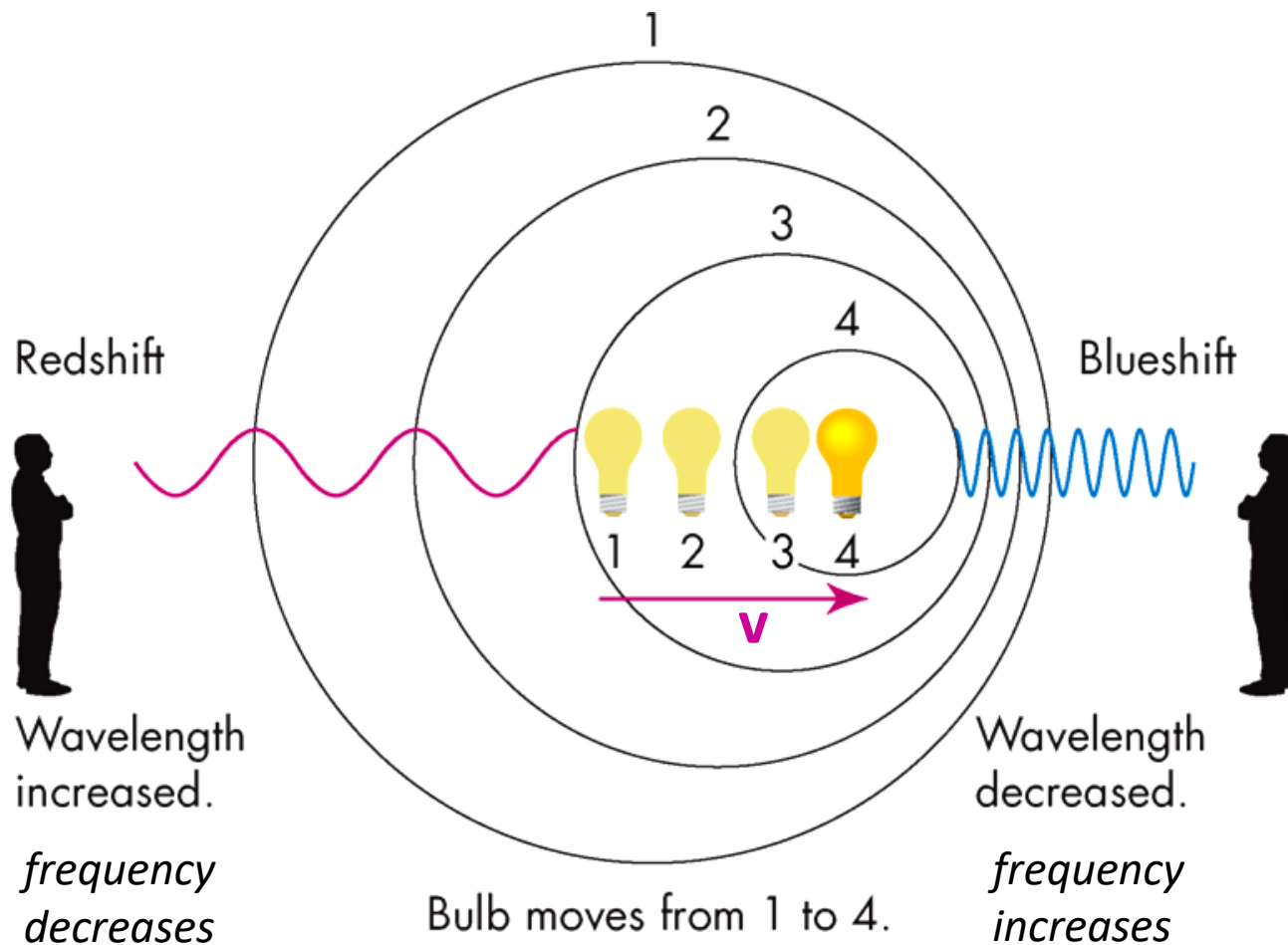
1. **Continuous** spectrum, e.g. a thermal **blackbody** source.
2. **Emission** spectrum (discrete): if light excites atoms, then the atomic emission will be at discrete frequencies.
3. **Absorption** spectrum (discrete): if a continuous spectrum excites atoms, then the absorption of photons will remove light at discrete frequencies (“shadow lines”).

Doppler Effect

A **moving source** cannot change the speed of its emitted light, but it does change its **frequency & wavelength**.

Doppler Effect

A **moving source** cannot change the speed of its emitted light, but it does change its **frequency & wavelength**.




Doppler Effect


A **moving source** cannot change the speed of its emitted light, but it does change its **frequency & wavelength**.

Works for sound too !!!

Doppler Shift Calculation

Doppler frequency shift: $\frac{\Delta f}{f} = -\frac{\Delta \lambda}{\lambda} = \frac{v}{c}$ with $f' = f + \Delta f$

frequency of stationary source 

perceived frequency of moving source 

Doppler Shift Calculation

Doppler frequency shift: $\frac{\Delta f}{f} = -\frac{\Delta \lambda}{\lambda} = \frac{v}{c}$ with $f' = f + \Delta f$

frequency of stationary source \nearrow f

\nearrow perceived frequency of moving source f'

If source is moving towards you, then light is blue shifted.

$$v > 0$$

$$\Delta f > 0, f' \text{ goes up}$$

$$\Delta \lambda < 0$$

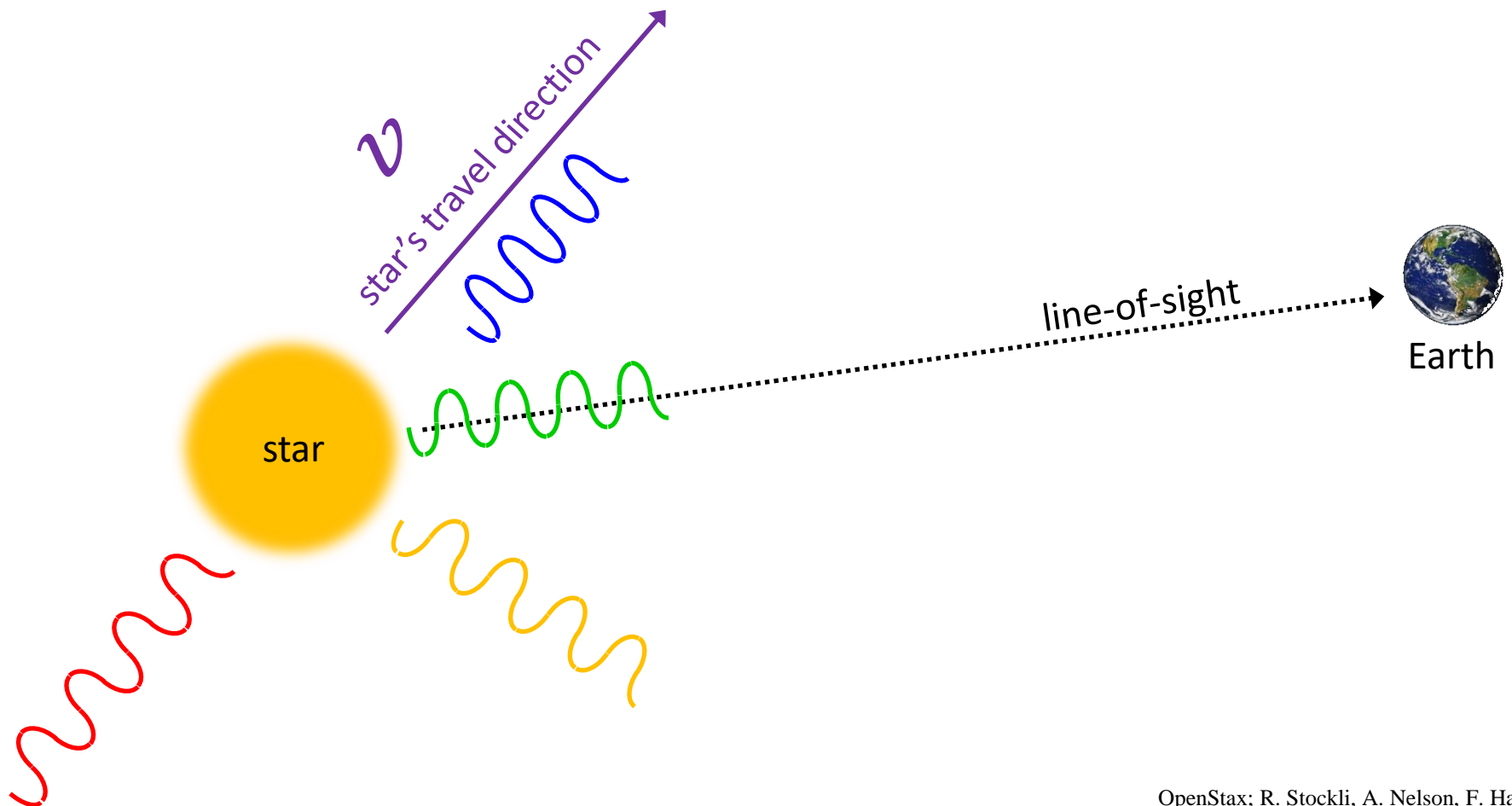
If source is moving away from you, then light is red shifted.

$$v < 0$$

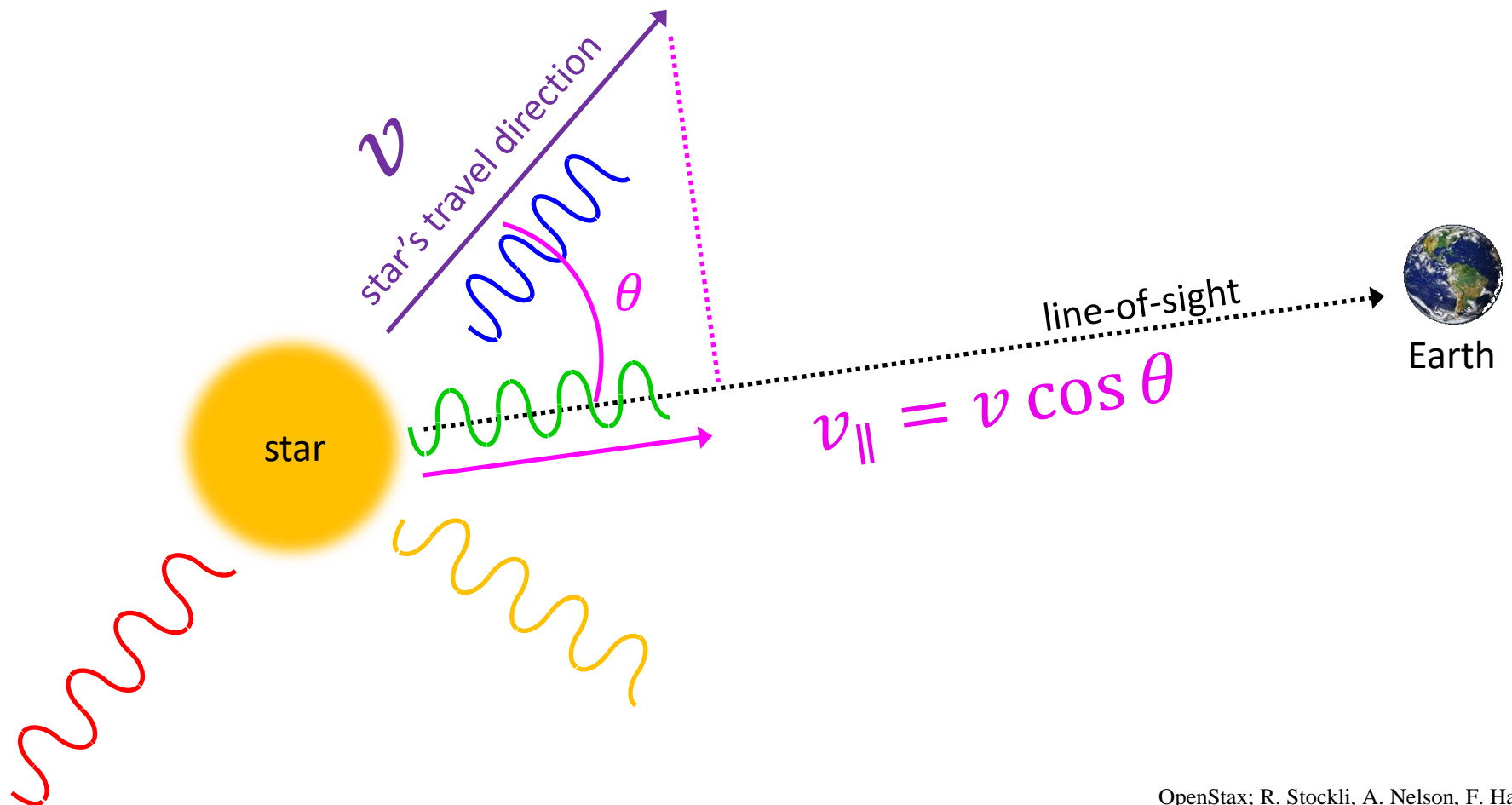
$$\Delta f < 0, f' \text{ goes down}$$

$$\Delta \lambda > 0$$

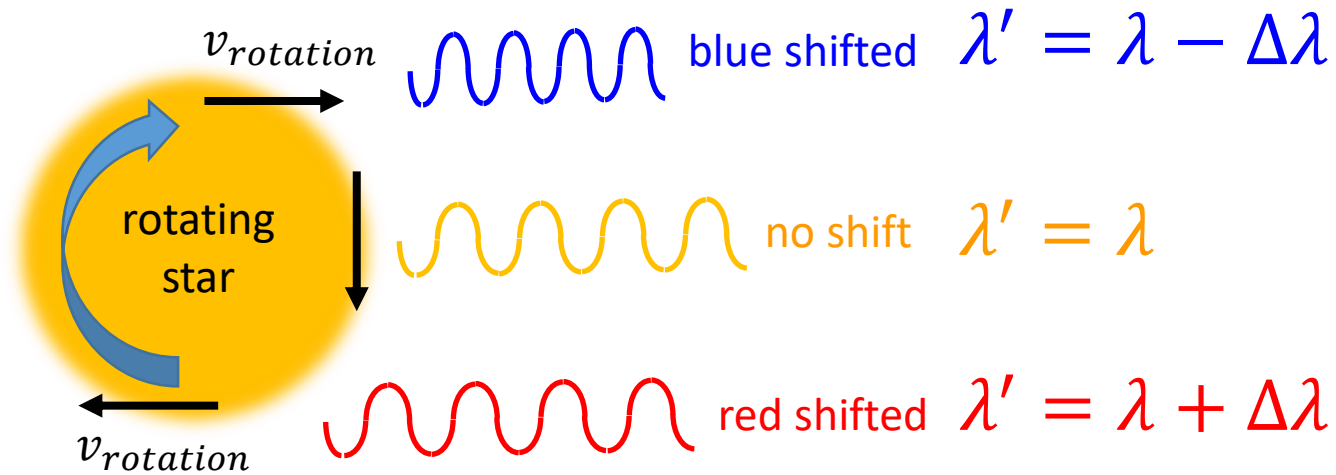
Doppler Shift is for Line-of Sight Velocity Component



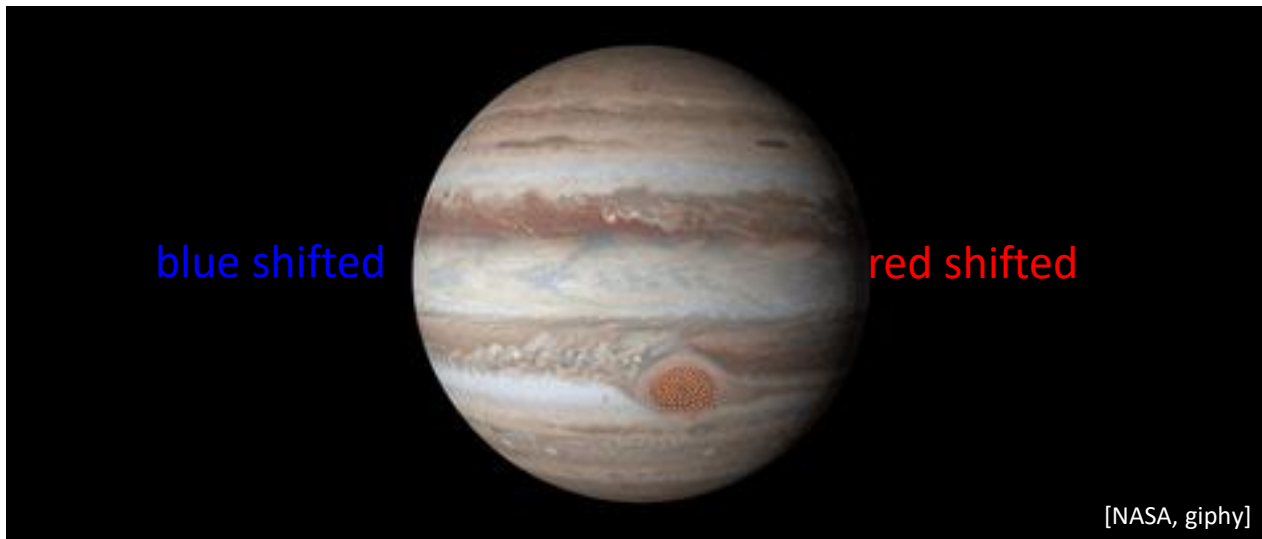
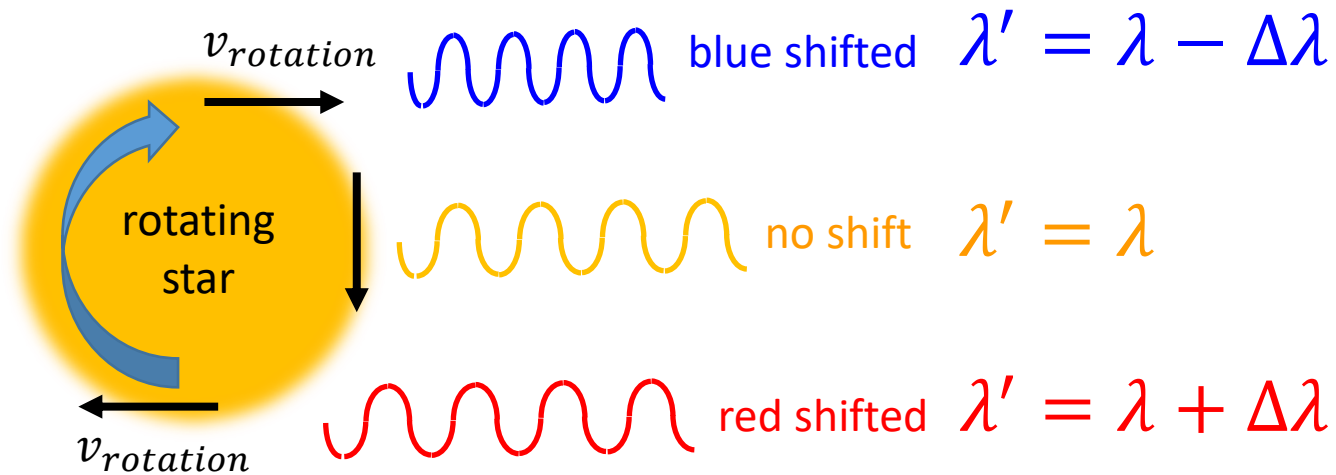
Doppler Shift is for Line-of Sight Velocity Component



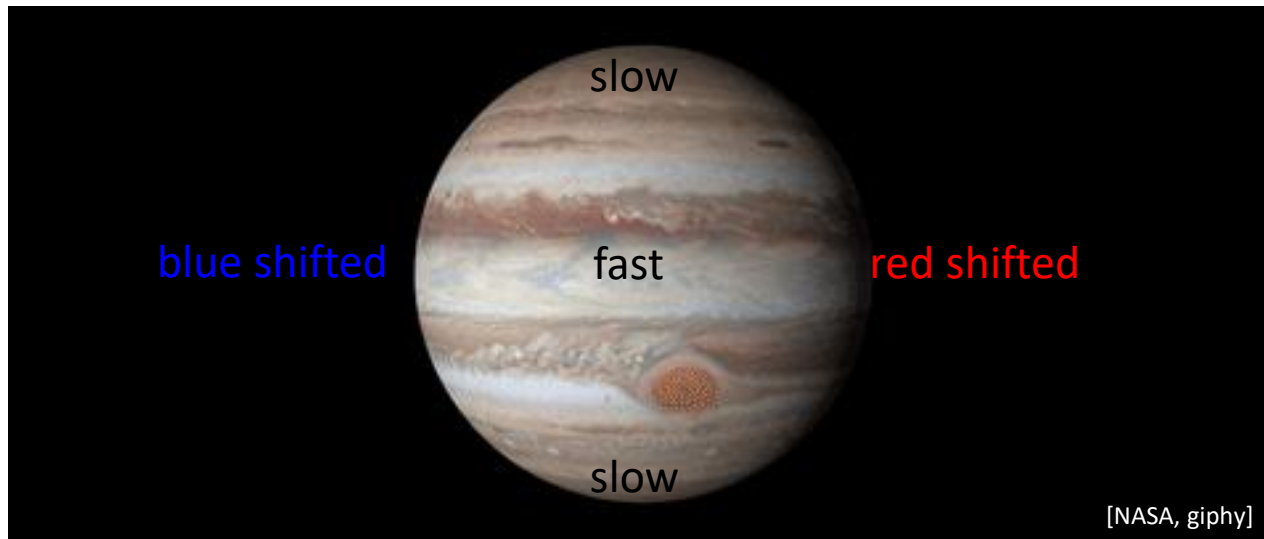
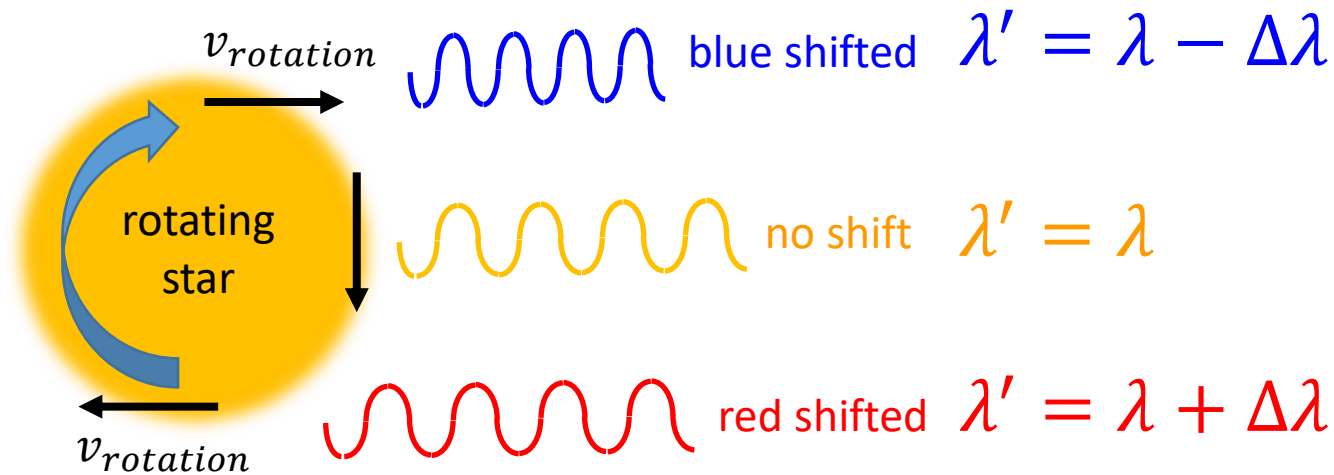
Doppler Shifts for Rotating Sources



Doppler Shifts for Rotating Sources



Doppler Shifts for Rotating Sources



Nuclear Particles

Photons are the easiest particles with which to view space, but they are not the only ones

- Protons (p or p^+), electrons (e or e^-), neutrons (n)
- Alpha particles (α)
- Neutrinos (ν)
- Anti-particles: Positrons (e^+) & anti-protons (p^-)
- Cosmic rays (high energy p^+ , p^- , e^+ , e^- , α , etc)

Particle Properties

Particle	Mass (kg)	Electric charge	Forces
Proton	1.67265×10^{-27}	+1	Strong, EM, weak, gravity
Neutron	1.67495×10^{-27} <i>$m_n \sim m_p$</i>	0	Strong, weak, gravity
Electron	9.11×10^{-31} <i>$m_e \sim 1/2000$ of m_p</i>	-1	EM, weak, gravity
Neutrino	$< 2 \times 10^{-36}$	0	weak, gravity



barely interacts with anything !!! (very hard to detect)

Particle Properties

Particle	Mass (kg)	Electric charge	Forces
Proton	1.67265×10^{-27}	+1	Strong, EM, weak, gravity
Neutron	1.67495×10^{-27} <i>$m_n \sim m_p$</i>	0	Strong, weak, gravity
Electron	9.11×10^{-31} <i>$m_e \sim 1/2000$ of m_p</i>	-1	EM, weak, gravity
Neutrino	$< 2 \times 10^{-36}$	0	weak, gravity

 barely interacts with anything !!! (very hard to detect)

Fundamental forces

There are only 4 fundamental forces that we know of:

Strong nuclear force, **electromagnetic force**, **weak nuclear force**, **gravity**.

(holds nucleus together)
[short range]

(generates radioactive decay) (very very weak)
[short range]