Today's Topics

Wednesday, March 5, 2025 (Week 6, lecture 15) – Chapters 15, 16.

A. The Sun: a visual introduction

- B. Surface of the Sun
- C. Internal structure
- D. Solar fusion

Reminder: Problem Set #5 is due on ExpertTA on Friday, March 7, by 9:00 am.





Our Sun



Our Sun



Transit of Venus, 2012. *(visible light)*

Solar Equilibrium Gravity vs Fusion Heat



Hydrostatic Equilibrium: In the Sun (and any star), the **inward force of gravity** is **exactly balanced** at each point by the **outward force of gas pressure** due to **heat** from <u>nuclear fusion</u>.





[NASA: Solar Dynamics Observatory, UV light at 304 nm, 2010]



Sun view at 1600 nm.

Transit of Mercury, November 11, 2019



Transit of Mercury, November 11, 2019



Transit of Mercury, May 9, 2016



June 26, 2019 | The International Space Station passes in front of the sun as seen from Gyöngyös, Hungary. (Peter Komka/EPA-EFE/Shutterstock)

Our Sun

Blackbody Radiation Source



[Wikimedia Commons: Danmichaelo, public domain]

Our Sun: Surface (photosphere)

Properties

- Temperature = 5777 K (surface/photosphere)
- Substance: Plasma (electrons & nuclei are dissociated).
- Magnetosphere: ~ 1 Gauss at surface. (exception: sunspots at 3000 G)
- Rotation period: T_{equator} = 25 days, T_{poles} = 34 days.
- Rotation axis tilt: 7.25° with respect to ecliptic.



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Elemental Abundance in Photosphere

Solar Wind

Extension of the Corona



Solar Wind

Extension of the Corona

- Corona is very hot: 1 million Kelvin
- Solar wind consists of:
 - ➢ protons
 - ➤ electrons
 - alpha particles (He nuclei)
- Energy range: 0.01 10 keV
- Solar wind speed: 400 750 km/s
- Strongest emission is from coronal holes.



Our Sun's Surface

[NASA: Solar Dynamics Observatory, October 18, 2010]

Our Sun: Sunspots & Granules



Sunspot size ~ 10,000-20,000 km Sunspot = cooler surface region with strong magnetic field. → convection is impede by magnetic field. Granule size ~ 1500 km granule = convective cell

Solar Cycle: 11 year period



Sunspots: 11 year cycle

Monthly Average Sunspot Numbers



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The following all vary in-sync with the solar cycle:

- Number of **sunspots**.
- Solar flares and coronal mass ejections.
- Total solar irradiance (but only by 0.1 %).
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PollEv Quiz: PollEv.com/sethaubin

Our Sun: Structure

Structure determined from:

- Computer modelling.
- Helioseismology.
- **Neutrino** measurements.



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1.0





(see also Feb. 17 lecture)

9 billions years *weak force*



2 × 1.442 MeV

(Note: $1 \text{ eV} = 1.602 \times 10^{-19} \text{ J}$)

(see also Feb. 17 lecture)

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4 seconds strong force



(Note: $1 \text{ eV} = 1.602 \times 10^{-19} \text{ J}$)

 $2 \times 1.442 \text{ MeV}$

+ 2 × 5.49 MeV

(see also Feb. 17 lecture)

By Sarang - Own work, Public Domain, https://commons.wikimedia.org/w/index.php?curid=51118538



lecture)

$$Energy = E = mc^{2}$$

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Mass of a proton: $m_p = 1.6726 \times 10^{-27}$ kg Mass of 4 protons: $4 \times m_p = 6.6905 \times 10^{-27}$ kg Mass of ⁴He nucleus: $m_{He} = 6.6447 \times 10^{-27}$ kg

Note: $4m_p - m_{He} = 4.65 \times 10^{-29}$ kg difference is due to two positrons !