

Wednesday, March 26, 2025

#1

Example: Rotation frequency of a Sun-based white dwarf ... and "neutron star".

major assumption: no mass loss

the Sun cannot produce a neutron star, but this is a "what if" exercise.

note: $f_{\text{sun}} = \frac{1}{T_{\text{sun}}} = 4.63 \times 10^{-7} \text{ Hz}$

$$T_{\text{sun}} = 25 \text{ days} = 2.16 \times 10^6 \text{ s}$$

According to the Radius vs. mass plot, a 1 solar mass white dwarf has a radius of

$$R_{\text{wd}} = 0.0078 R_{\text{sun}} = 5400 \text{ km}$$

For the neutron star, we will use $R_{\text{ns}} = 14 \text{ km}$

$$= \frac{14}{6.96 \times 10^5} R_{\text{sun}}$$

$$\approx 2 \times 10^{-5} R_{\text{sun}}$$

When the Sun contracts into a white dwarf (no mass loss) angular momentum is conserved.

$$L = mvr$$

for a particle going in a circle

$$v = \frac{2\pi r}{T} = 2\pi r f$$

↑
period

↑
frequency

$$\Rightarrow L = 2\pi r^2 f m$$

for a rotating sphere: $L \propto 2\pi r^2 f m$

$$\text{or } L = \alpha 2\pi r^2 f m$$

↑ "proportional" constant
that accounts for the
geometry of the sphere

If angular momentum is conserved, then

$$L_{\text{sun}} = \alpha 2\pi r_{\text{sun}}^2 f_{\text{sun}} m$$

$$L_{\text{wd}} = \alpha 2\pi r_{\text{wd}}^2 f_{\text{wd}} m$$

$$L_{\text{ns}} = \alpha 2\pi r_{\text{ns}}^2 f_{\text{ns}} m$$

with $L_{\text{sun}} = L_{\text{wd}} = L_{\text{ns}}$

(m remains unchanged
by assumption)

$$\text{Thus } \frac{L_{\text{sun}}}{L_{\text{wd}}} = \frac{\alpha 2\pi r_{\text{sun}}^2 f_{\text{sun}} m}{\alpha 2\pi r_{\text{wd}}^2 f_{\text{wd}} m} \Leftrightarrow f_{\text{wd}} = \left(\frac{r_{\text{sun}}}{r_{\text{wd}}}\right)^2 f_{\text{sun}}$$

$$\Rightarrow f_{\text{wd}} = \left(\frac{1}{0.0078}\right)^2 f_{\text{sun}} = (16437) (4.63 \times 10^{-7}) = 0.00761 \text{ Hz}$$

$$\Rightarrow f_{\text{wd}} = 0.0076 \text{ Hz}$$

$$\Rightarrow T_{\text{wd}} = \frac{1}{f_{\text{wd}}} = 131 \text{ s} = 2 \text{ minutes } 11 \text{ seconds}$$

neutron star :

$$f_{ns} = \left(\frac{r_{sun}}{r_{ns}} \right)^2 f_{sun} = \left(\frac{1}{2 \times 10^{-5}} \right)^2 f_{sun}$$
$$= (2.5 \times 10^9) (4.63 \times 10^{-7})$$
$$= 1153 \text{ Hz}$$

$$\Rightarrow f_{ns} = 1153 \text{ Hz} \quad (\Rightarrow) \quad T_{ns} = 0.86 \text{ ms} = 0.86 \times 10^{-3} \text{ s}$$

↑
neutron star
rotation rate

↑
neutron star
rotation period