

Today's Topics

Wednesday, February 4, 2026 (Week 2, lecture 6) – Chapter 3.

- A. Galilean Relativity
- B. Newton's Laws
- C. Momentum & energy
- D. Gravity *by Newton*

Galilean Relativity

Definition

An **inertial frame** is a *coordinate system* moving at **constant velocity**.
[constant velocity = constant speed & constant direction]

- Inertial frame = space that travels with you, e.g. car, airplane, rocket, etc ...
- Note: an accelerating/rotating system is NOT an inertial frame.

Galilean Relativity

Definition

An **inertial frame** is a *coordinate system* moving at **constant velocity**.
[constant velocity = constant speed & constant direction]

- Inertial frame = space that travels with you, e.g. car, airplane, rocket, etc ...
- Note: an accelerating/rotating system is NOT an inertial frame.

Galilean relativity posits that in any inertial frame:

“you cannot tell that you are moving based on local measurement.”

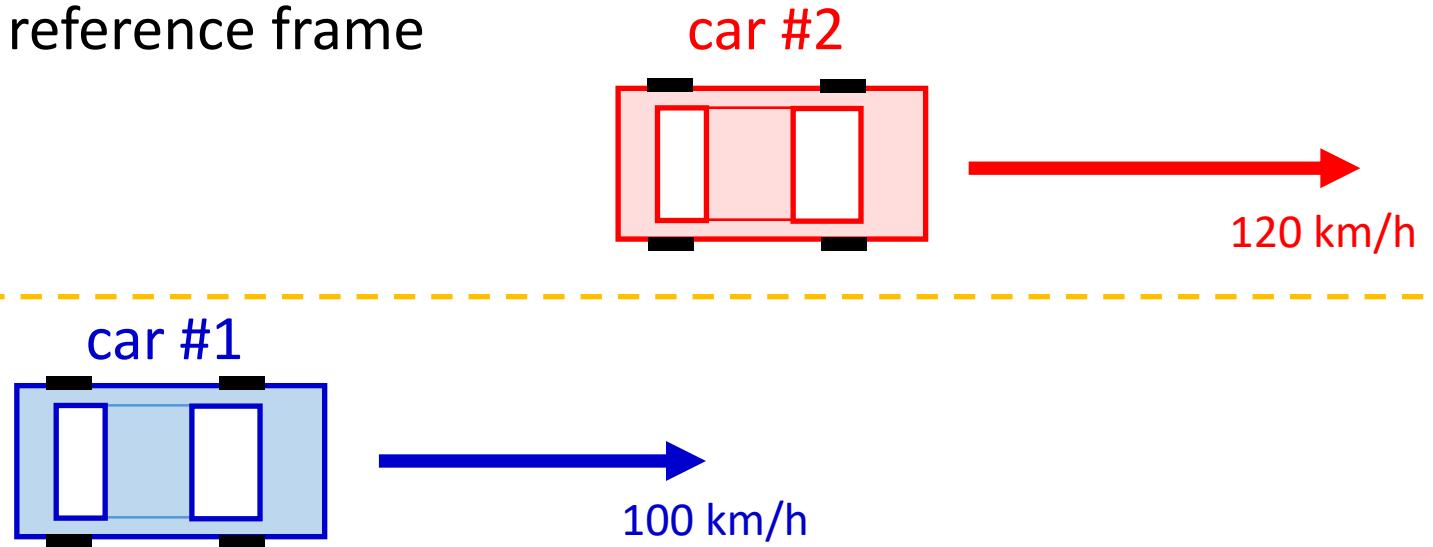
- i.e. an inertial frame locally behaves as if it is at rest (locally).
- **corollary**: an object in uniform motion will tend to stay in uniform motion.

Examples:

1. **Car**: You cannot tell that a car is moving (when at constant velocity) unless you look out window.
2. **Airplane**: You cannot tell an airplane is moving (when at constant velocity) unless you look out window (or hit turbulence).

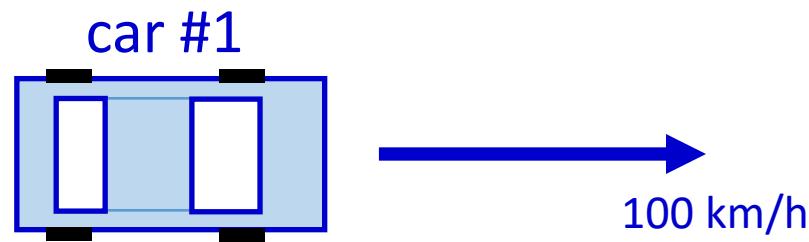
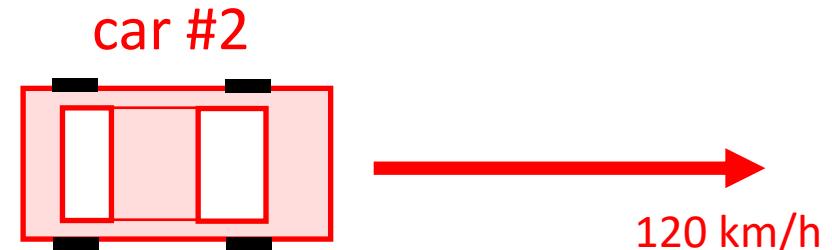
Galilean Relativity Example

Earth / road's reference frame

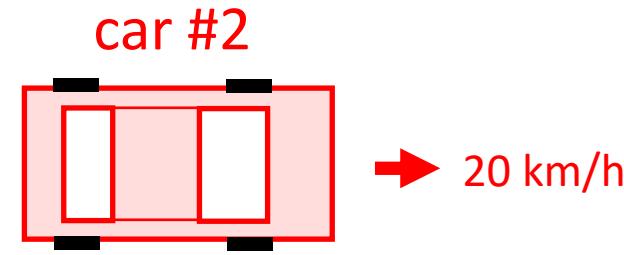


Galilean Relativity Example

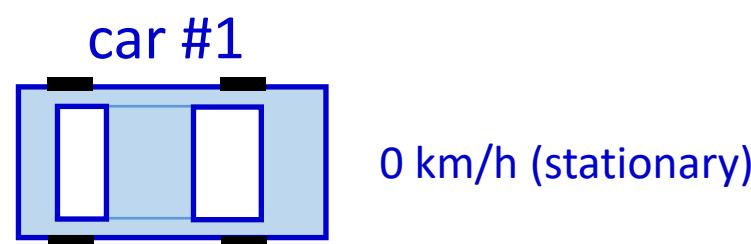
Earth / road's reference frame



Car # 1's reference frame

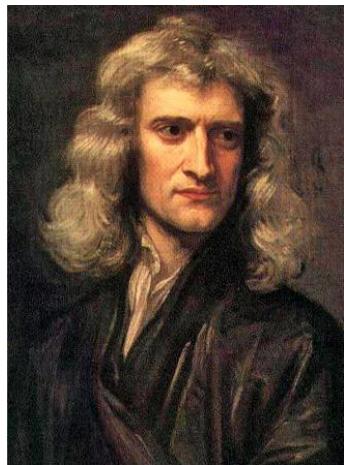


-100 km/h



PollEv Quiz: PollEv.com/sethaubin

Isaac Newton: Founder of Classical Mechanics

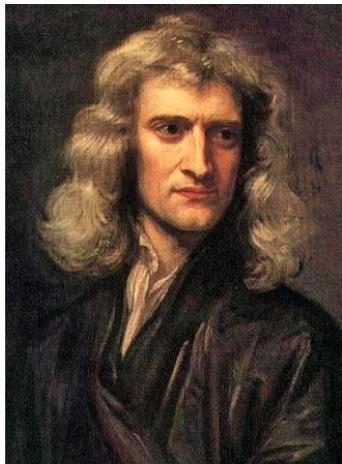


Sir Isaac Newton (1643-1727)

- Cambridge U.
- Founded **Classical Mechanics**.
- Discovered **Calculus**.
- Major contributions to **Optics & Astronomy**.

Newton (1689) [by G. Kneller]

Isaac Newton: Founder of Classical Mechanics



Sir Isaac Newton (1643-1727)

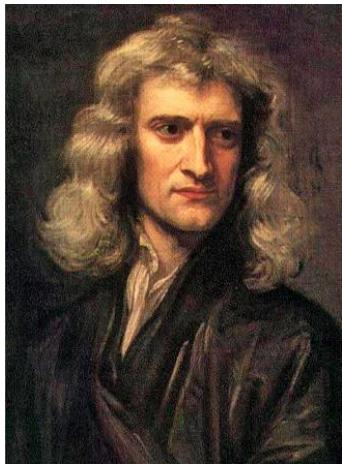
- Cambridge U.
- Founded **Classical Mechanics**.
- Discovered **Calculus**.
- Major contributions to **Optics & Astronomy**.

Newton (1689) [by G. Kneller]

Classical Mechanics

- “Newton’s Laws” of classical mechanics.
- Law of universal gravitation.
- Newton’s laws are used for ***calculating planetary & stellar motion.***
(+ Einstein’s “Special Relativity”)

Isaac Newton: Founder of Classical Mechanics



Sir Isaac Newton (1643-1727)

- Cambridge U.
- Founded **Classical Mechanics**.
- Discovered **Calculus**.
- Major contributions to **Optics & Astronomy**.

Newton (1689) [by G. Kneller]

Classical Mechanics

- “Newton’s Laws” of classical mechanics.
- Law of universal gravitation.
- Newton’s laws are used for **calculating planetary & stellar motion.**
(+ Einstein’s “Special Relativity”)

Astronomy

- **Optics:** white light & colors, refraction.
- Invented the **reflecting telescope**.

Newton's Laws of Classical Mechanics

1st Law: An object moves at constant velocity if there is no net force acting on it.

[fine print: in an inertial reference frame]

2nd Law: Force = mass \times acceleration.

3rd Law: For any force, there is always an equal and opposite reaction force.

Newton's 1st Law

An object moves at constant velocity if there is no net force acting on it.

[fine print: in an inertial reference frame]

Note: This law is a variation on the Galilean relativity statement.

Newton's 2nd Law

Force = Mass \times Acceleration

or

$$F = ma$$

F = net force

m = mass

a = acceleration

[**fine print:** in an inertial reference frame]

Newton's 2nd Law

Force = Mass \times Acceleration

or

$$F = ma$$

F = net force

m = mass

a = acceleration

[**fine print:** in an inertial reference frame]

Note 1: This equation is mostly useful if you know the net force applied.

Note 2: If the acceleration is zero, then the net force is zero.

Newton's 3rd Law

For any force, there is always an equal and opposite reaction force

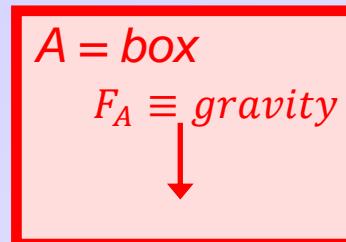
$$F_{A \rightarrow B} = -F_{B \rightarrow A}$$

Newton's 3rd Law

For any force, there is always an equal and opposite reaction force

$$F_{A \rightarrow B} = -F_{B \rightarrow A}$$

Example:



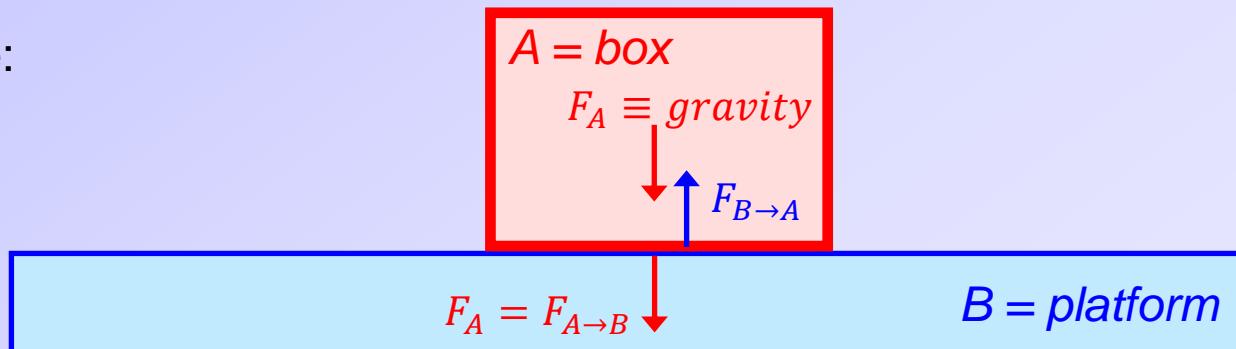
B = platform

Newton's 3rd Law

For any force, there is always an equal and opposite reaction force

$$F_{A \rightarrow B} = -F_{B \rightarrow A}$$

Example:

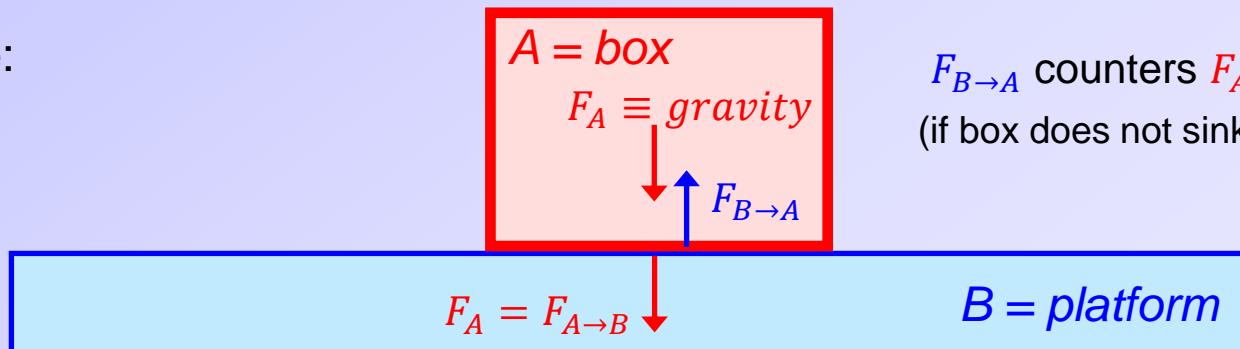


Newton's 3rd Law

For any force, there is always an equal and opposite reaction force

$$F_{A \rightarrow B} = -F_{B \rightarrow A}$$

Example:



Newton's 3rd Law: Rocket Thrust

A rocket accelerates by pushing on its exhaust.

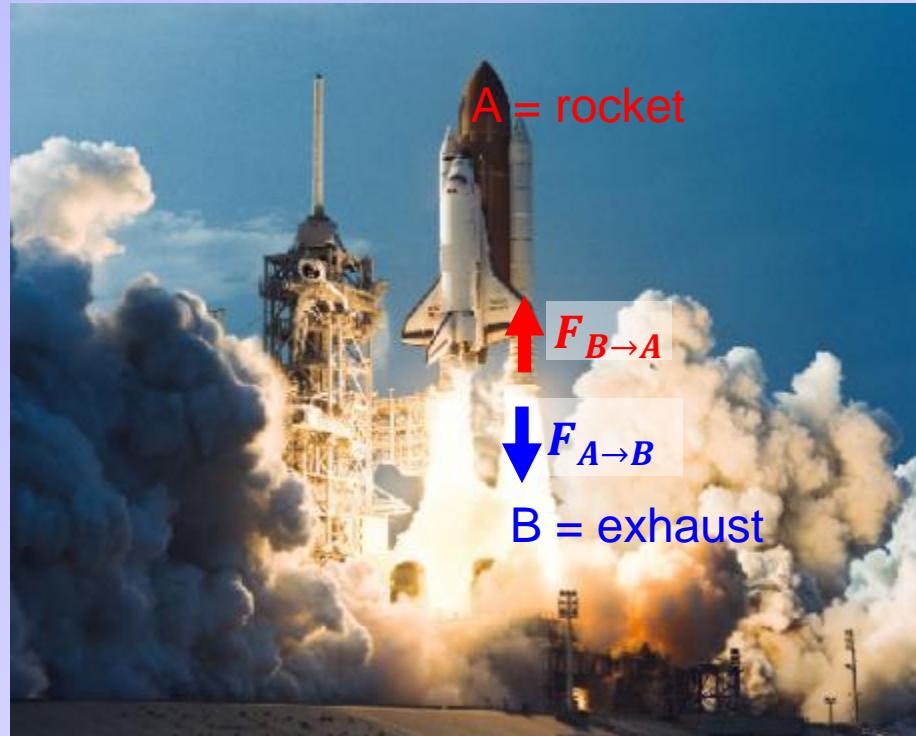


A rocket does NOT push on the air to accelerate.

A rocket does NOT push on its platform to accelerate.

Newton's 3rd Law: Rocket Thrust

A rocket accelerates by pushing on its exhaust.



A rocket does NOT push on the air to accelerate.

A rocket does NOT push on its platform to accelerate.

Newton's Laws of Classical Mechanics

1st Law: An object moves at constant velocity if there is no net force acting on it.

[fine print: in an inertial reference frame]

2nd Law: Force = mass \times acceleration.

3rd Law: For any force, there is always an equal and opposite reaction force.

Newton's Laws of Classical Mechanics

1st Law: An object moves at constant velocity if there is no net force acting on it.

[fine print: in an inertial reference frame]

2nd Law: Force = mass \times acceleration.

3rd Law: For any force, there is always an equal and opposite reaction force.



- Conservation of **Momentum**.
- Conservation of **Energy**.

Conservation of Momentum

momentum = mass \times velocity

total momentum

= sum of the momenta of all the sub-parts of a system

Conservation of Momentum

momentum = mass \times velocity

total momentum

= sum of the momenta of all the sub-parts of a system

Conservation Law

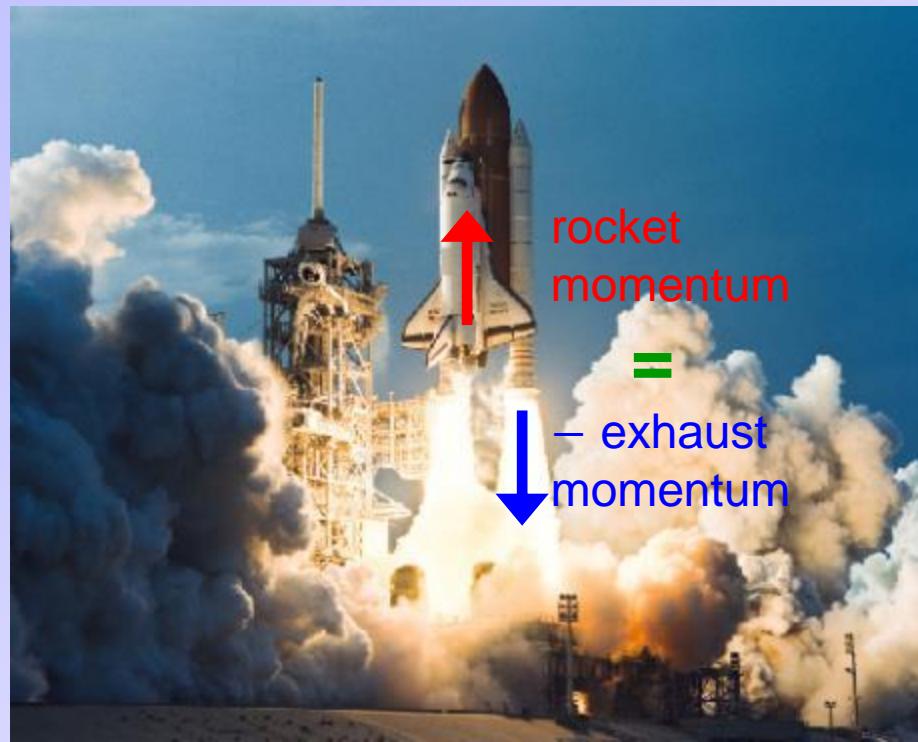
The **total momentum** of a closed system **never changes**.



*no external objects enter
no external forces*

Momentum Conservation: Rocket Thrust

$$\text{Momentum}_{\text{rocket}} + \text{Momentum}_{\text{exhaust}} = 0$$



Conservation of Energy

$$\text{Kinetic Energy} = E_k = \frac{1}{2}mv^2$$

m = mass
 v = speed

Potential Energy = “stored” energy

example: gravitational potential energy

Total Energy

= sum of the energies of all the sub-parts of a system

Conservation of Energy

$$\text{Kinetic Energy} = E_k = \frac{1}{2}mv^2$$

m = mass

v = speed

Potential Energy = “stored” energy

example: gravitational potential energy

Total Energy

= sum of the energies of all the sub-parts of a system

Conservation Law

The **total energy** of a closed system **never changes**.

Gravity

Newton figured out that the same force that is responsible for a *falling apple* is also responsible for keeping the ***Moon in orbit*** around the Earth.

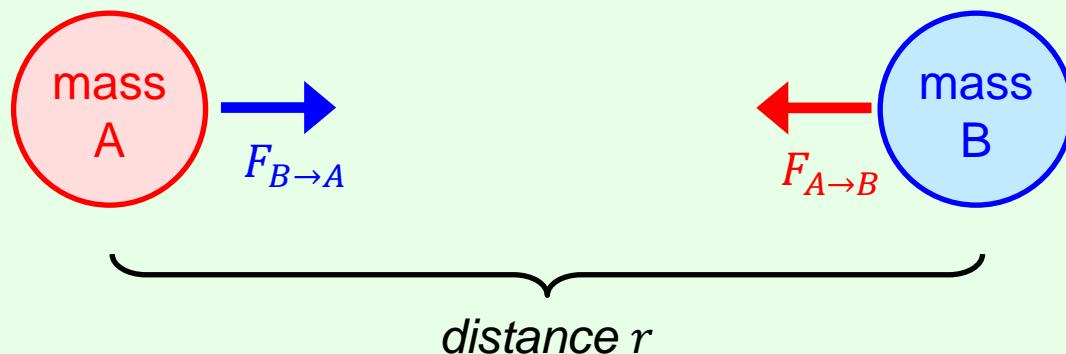
Gravity

Newton figured out that the same force that is responsible for a *falling apple* is also responsible for keeping the **Moon in orbit** around the Earth.

Newton's law of universal gravitation

All masses attract each other according to the following relation:

$$F_{A \rightarrow B} = -G \frac{M_A M_B}{r^2} = -F_{B \rightarrow A}$$



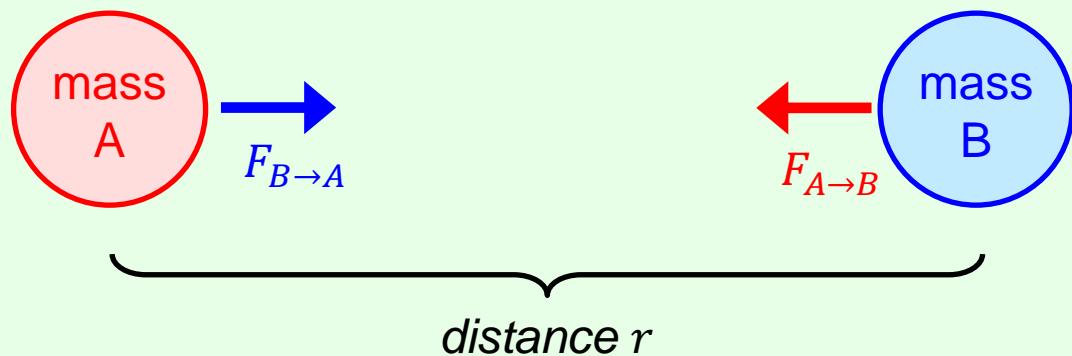
Gravity

Newton figured out that the same force that is responsible for a *falling apple* is also responsible for keeping the **Moon in orbit** around the Earth.

Newton's law of universal gravitation

All masses attract each other according to the following relation:

$$F_{A \rightarrow B} = -G \frac{M_A M_B}{r^2} = -F_{B \rightarrow A}$$



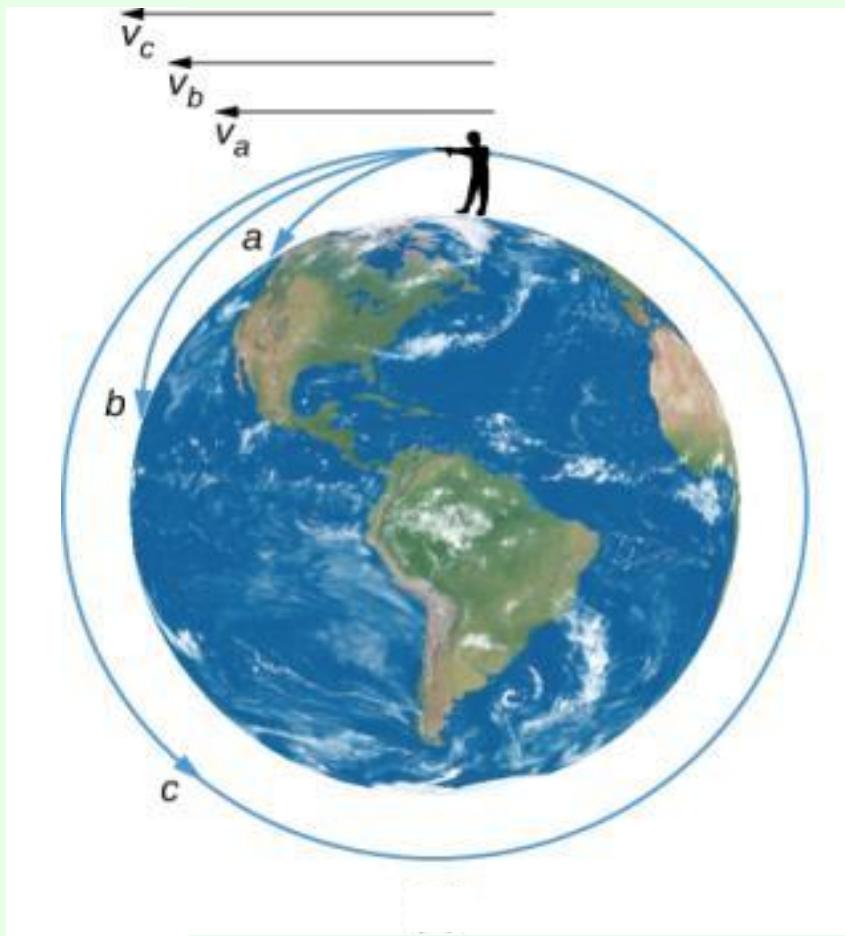
Properties

- Falls off as $1/r^2$.
- Proportional to M_A .
- Proportional to M_B .
- G = Newton's constant
 $= 6.67430(15) \times 10^{-11}$
 $m^3/Kg \cdot s^2$

**Why do all objects
fall
at the same rate?**

(to be covered in problem session)

Orbiting is free falling while missing Earth

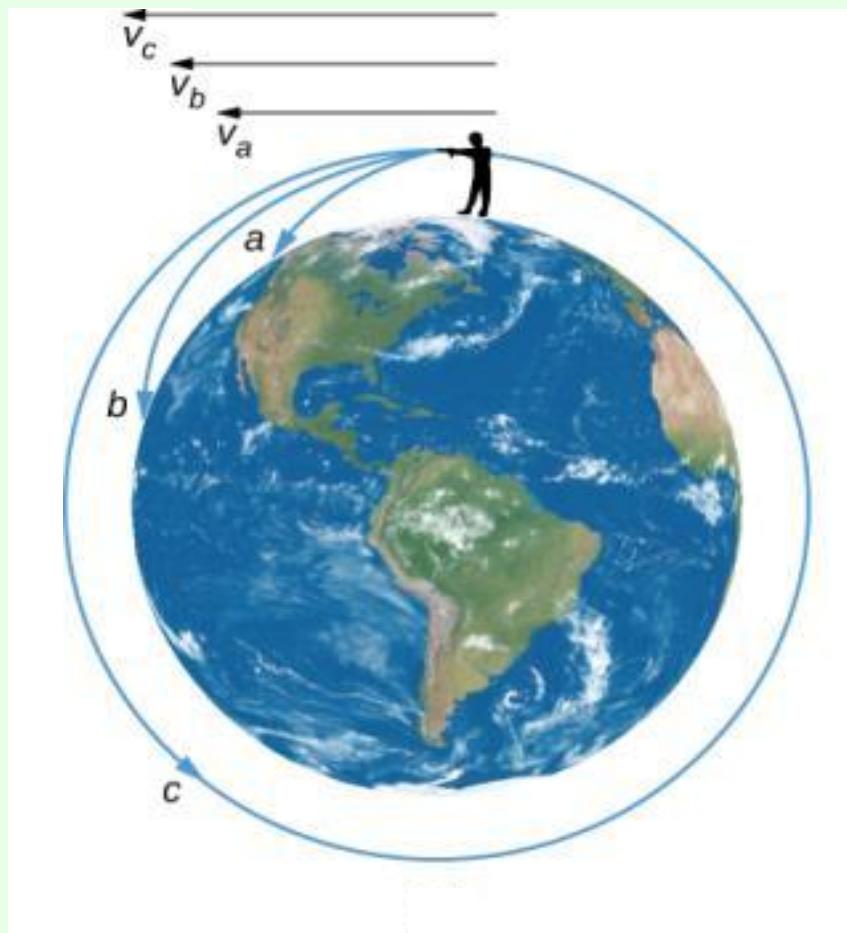


Paths a & b: Initial speeds are weak enough that Earth's gravity pulls the projectile back to the surface.

Path c: Initial speed is strong enough that Earth's gravity never pulls the projectile back to the surface.

[OpenStax: Astronomy]

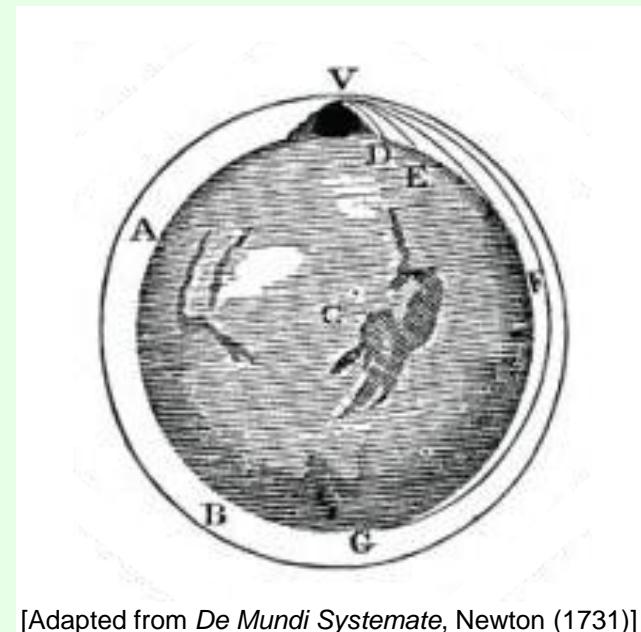
Orbiting is free falling while missing Earth



[OpenStax: Astronomy]

Paths a & b: Initial speeds are weak enough that Earth's gravity pulls the projectile back to the surface.

Path c: Initial speed is strong enough that Earth's gravity never pulls the projectile back to the surface.



[Adapted from *De Mundi Systemate*, Newton (1731)]

orbiting



“The knack of flying is learning how to throw yourself at the ground and miss”

- Hitchhikers Guide to the Galaxy

Weightless in Orbit



Clockwise from top left: Tracy Caldwell Dyson (NASA), Naoko Yamzaki (JAXA), Dorothy Metcalf-Lindenburger (NASA), and Stephanie Wilson (NASA). (credit: NASA)

Astronauts in Free Fall: While in space, astronauts are falling freely, so they experience “weightlessness.”