

Week 0: Thursday, August 28, 2025

Classical E-M quantities:

$$\text{Energy} = \frac{1}{2} \int_{\text{all space}} \left[\epsilon_0 \vec{E}^2 + \frac{1}{\mu_0} \vec{B}^2 \right] d^3r \stackrel{?}{=} H$$

$$\text{Momentum} = \vec{P} = \epsilon_0 \int_{\text{all space}} \vec{E} \times \vec{B} d^3r$$

$\vec{E} \& \vec{B}$: Wavefunctions or Operators?

Operators

- obey the superposition principle
- oscillate, wave-like, spread out over space
- Measurable quantities $\rightarrow$ operators

From relativistic mechanics, we know:

For photons: $E^2 = p^2 c^2 + \cancel{m_0^2 c^4} = 0 \text{ for photon}$

$E = \text{energy}$

$$\Rightarrow E = pc$$

[in E&M: $\mathcal{E} = pc$]

$\mathcal{E}$: energy density p : momentum density

In QM, particles have momentum

$$\boxed{\vec{p} = \hbar \vec{k}}$$

$$\Rightarrow p = \hbar \frac{2\pi f}{c} = \hbar \frac{\omega}{c}$$

$k = \text{wavevector}$
 $= \frac{2\pi}{\lambda} = \frac{2\pi f}{c}$
 $\lambda f = c$

$$\Rightarrow E_{\gamma} = pc = \hbar \frac{\omega}{c} \cdot c \Rightarrow$$

$$\boxed{E = \hbar \omega} \stackrel{?}{=} H_{\text{photon}}$$

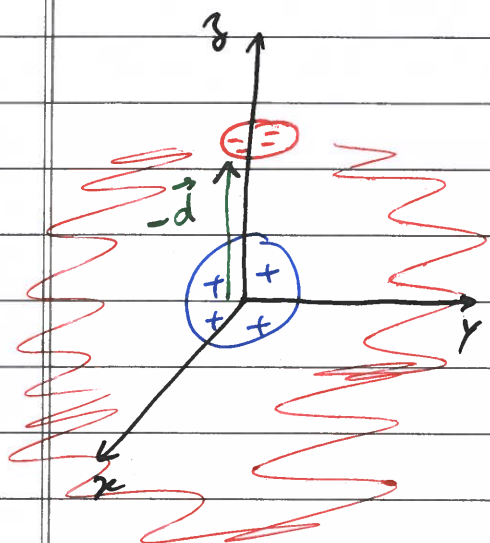
Semi-Classical Atom-light interaction

(I) Motivation

- Classical physics often contains the essential physics and then there are "quantum corrections".
- Later, when we do the quantum treatment, we will be able to distinguish more easily between classical & quantum properties of atom light interactions.
- Easier than full QM treatment.

(II) Model Atom

We consider a "semi-classical atom" with the



- { nucleus: "+" charge = $+e$
 $1 e^-$: "-" charge = $-e$

- no preferred orientation of dipole
 $\vec{d} = -e\vec{r}$

so we will use an isotropic e^- cloud (nucleus is very heavy)

$$\hookrightarrow \langle \vec{d} \rangle = 0 = -e \langle \vec{r} \rangle$$

- Mysteriously, e^- does not spiral into the nucleus, but remains at an equilibrium distance:

$$\sqrt{\langle \vec{r}^2 \rangle} = r_{\text{equilibrium}} \text{ (but } \langle \vec{r} \rangle = 0 \text{)}$$

- If the e^- is moved away from its equilibrium position, then there is a linear restoring force:

(at origin
on average)

$$\vec{F} = m \frac{d^2 \vec{r}}{dt^2} = -k \vec{r}$$

$$\Leftrightarrow \ddot{\vec{r}} + \frac{k}{m} \vec{r} = 0 \Leftrightarrow \ddot{\vec{r}} + \omega^2 \vec{r} = 0$$

with $\omega = \sqrt{\frac{k}{m}}$

(also $\ddot{d} + \frac{k}{m} d = 0$)

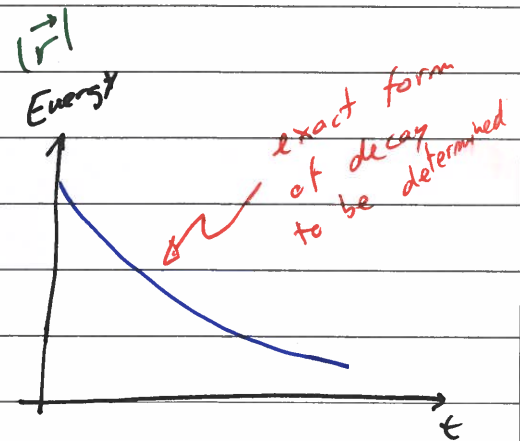
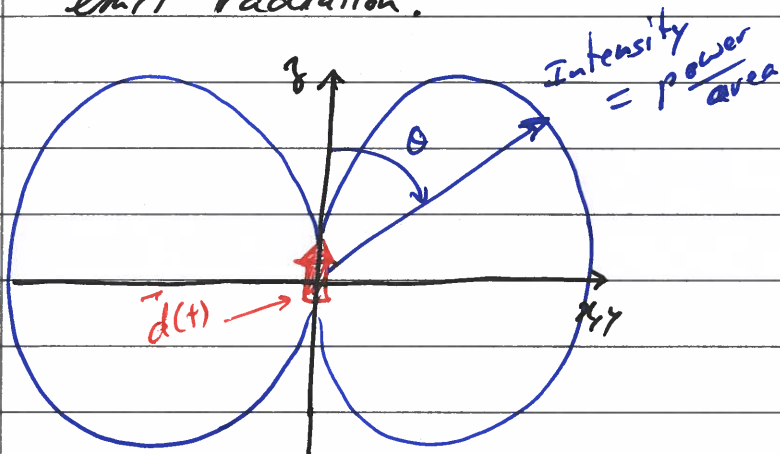
$\Rightarrow$ "simple harmonic atom"

Solution: $\vec{r}(t) = \vec{r}_0 \sin(\omega t) + \vec{r}'_0 \cos(\omega t)$

with $\vec{r}_0$ and $\vec{r}'_0$ defined by the initial conditions.

III Radiation Damping

Accelerating charges and oscillating dipoles (i.e., antennas) emit radiation.



$$\text{Intensity} = \langle \text{Poynting vector} \rangle = \langle \vec{S} \rangle = \frac{1}{\mu_0} (\vec{E} \times \vec{B})$$

$$= \frac{d_0^2}{32\pi \epsilon_0} \frac{\omega^4}{c^3} \frac{\sin^2 \theta}{r^2} \hat{r}$$

The total radiated power for an oscillating dipole $\vec{d} = \vec{d}_0 \sin(\omega t)$ is

$$\frac{d \langle \text{Energy} \rangle}{dt} = \langle \text{power} \rangle = -\frac{1}{4\pi \epsilon_0} \frac{d_0^2 \omega^4}{3c^3} = -\frac{1}{4\pi \epsilon_0} \frac{e^2 \omega^4 r_0^2}{3c^3}$$

(see Griffiths p401-405)

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What's the form of the damping?

Assume that the damping rate is much smaller than the oscillation rate.

In 1D: Energy of "harmonic" atom = $\frac{1}{2} k r_0^2 = \frac{1}{2} m \omega^2 r_0^2 = E_{H0}$

amplitude of harmonic atom oscillations

$$\text{Energy loss rate} = \frac{d}{dt} E_{H0} = -\frac{1}{4\pi \epsilon_0} \frac{e^2 \omega^4 r_0^2}{3c^3}$$

$$= -2 \underbrace{\frac{1}{4\pi \epsilon_0} \frac{e^2 \omega^2}{3mc^3}}_{\gamma} E_{H0}$$

$$\Leftrightarrow \frac{d}{dt} E_{H0} = -2\gamma E_{H0}$$

$\Rightarrow$

$$E_{H0}(t) = E_{H0} \Big|_{t=0} e^{-2\gamma t}$$

If $\gamma \ll \omega$, then $r_0(t)$ varies slowly compared to oscillations:

$$\Rightarrow r_0(t) = \sqrt{\frac{2 E_{H0, t=0}}{m \omega^2}} e^{-\gamma t}$$

$$\Rightarrow \text{electron oscillates as } r(t) = r_0(t) \sin(\omega t) = \underbrace{A e^{-\gamma t}}_{\text{damping \& oscillation}} \sin(\omega t)$$

damping & oscillation
are decoupled/independent