

Tuesday, October 21, 2025

Monte Carlo Wavefunctions — Quantum Trajectories (quantum jumps)

the "Schrodinger equation" for a 2-level dressed atom in a laser field with spontaneous emission is

$$i\hbar \frac{d}{dt} \begin{pmatrix} C_g \\ C_e \end{pmatrix} = \hbar \begin{pmatrix} \delta/2 & \Omega/2 \\ \Omega^*/2 & -\delta/2 \end{pmatrix} \begin{pmatrix} C_g \\ C_e \end{pmatrix} - \begin{pmatrix} 0 \\ i\hbar \frac{\gamma}{2} C_e \end{pmatrix}$$

$$= \hbar \begin{pmatrix} \delta/2 & \Omega/2 \\ \Omega^*/2 & -\delta/2 - i\gamma/2 \end{pmatrix} \begin{pmatrix} C_g \\ C_e \end{pmatrix}$$

effectively, $H = \hbar \begin{pmatrix} \delta/2 & \Omega/2 \\ \Omega^*/2 & -\delta/2 - i\gamma/2 \end{pmatrix}$

non-Hermitian

does not conserve probability

(complex eigenenergies)

time evolution: $e^{-i(\frac{E_2}{\hbar})t}$
 $= e^{\pm \frac{\gamma}{2}t}$

How do the states evolve in time?

Before a photon is spontaneously emitted:

$$|\psi_0(t)\rangle = \left[c_g(t) |g\rangle_a |N+1\rangle_e + c_e(t) |e\rangle_a |N\rangle_e \right] \otimes |0\rangle_{\text{all other vacuum photon states}}$$

$c_g(t)$ & $c_e(t)$ are given by the previous "Schrodinger equation".

After a ^{spontaneous} photon is emitted at time $t = t_s$, then

$$|\psi_1(t=t_s)\rangle = |g\rangle \otimes \left[\sum_{k, \epsilon} \alpha_{k, \epsilon} |k, \epsilon, 1\rangle \right]$$

all possible "1" photon states

photon polarization
momentum
accounts for dipole spatial emission distribution

Actually, the system (atom + laser + universe) could be in a superposition of these 2 states:

(leaked photon not yet detected)

$$|\psi(t)\rangle = \underbrace{\sqrt{1-\beta^2}}_{\text{for short times}} |\psi_0(t)\rangle + \underbrace{\beta}_{\text{1 photon leaks into an infinite superposition of photon momentum states}} |\psi_1(t)\rangle$$

no photon emitted

Loss of probability from H ends up here

$$\beta = \underbrace{\sqrt{\gamma |c_e|^2}}_{\text{for short times}} dt$$

Environment Detector

However, we are not dealing with a closed system, but an open quantum system. The "leaked photon" could be detected at any moment by the environment [environment is like a giant detector].

Let's assume that the "environment detector" is 100% efficient and makes a measurement in a time dt immediately after a photon leaks out $\rightarrow$ wavefunction collapses.

Immediately after collapse:
 "photon not emitted state"
 keeps evolving as before, but needs to be renormalized, since system is definitely here.

$$|\psi(t+dt)\rangle = \begin{cases} \psi_0(t+dt) & \text{with probability } p = 1 - \beta^2 \\ \text{or} \\ \psi_1(t+dt) = |g\rangle |h, \epsilon, 1\rangle & \text{with probability } p = \beta^2 \end{cases}$$

photon collapses to some h, ϵ state, but we don't pay attention to it.

$\hookrightarrow$ process repeats $\rightarrow$ propagate a time dt again.

This story suggests the following algorithm for calculating the time evolution of the "atom + laser + universe/environment detector" system though we only pay attention to the atomic state.

Step 1: At time $= t$, the state of the system is

$$|\psi(t)\rangle = |\psi_0(t)\rangle = c_g(t)|g\rangle + c_e(t)|e\rangle$$

Step 2: at time $= t+dt$, we have 2 possible trajectories for the wavefunction:

$$|\psi(t+dt)\rangle = \begin{cases} \text{case A: } \frac{1}{\sqrt{1-dp}} [c_g(t+dt)|g\rangle + c_e(t+dt)|e\rangle] & \text{with probability } 1-dp = 1-\beta^2 \\ \text{OR} \\ \text{case B: } |g\rangle & \text{with probability } dp \end{cases}$$

with $dp = \gamma |c_e|^2 dt = \beta^2$

re-normalization

Q: How do you pick between case A & case B?

↳ A: "flip a coin": choose a random number $\epsilon \in [0, 1]$

$$\begin{cases} \text{if } \epsilon > dp \Rightarrow |\psi(t+dt)\rangle = |\text{case A}\rangle \\ \text{if } \epsilon < dp \Rightarrow |\psi(t+dt)\rangle = |\text{case B}\rangle = |g\rangle \end{cases}$$

Condition on dt: $dt = \Delta t \ll \frac{1}{\gamma}, \frac{1}{\gamma}, \frac{1}{\delta}$

↑ not infinitesimal (though that would be ideal)



Random number generator must work well for very small numbers, i.e. $\epsilon \approx 0 (< dp)$!!!